Spin evolution of cold atomic gases in $SU(2) \otimes U(1)$ fields

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We consider response function and spin evolution in spin-orbit coupled cold atomic gases in a synthetic gauge magnetic field influencing solely the orbital motion of atoms. We demonstrate that various regimes of spin-orbit coupling strength, magnetic field, and disorder can be treated within a single approach based on the representation of atomic motion in terms of auxiliary collective classical trajectories. Our approach allows a unified description of fermionic and bosonic gases.

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I. INTRODUCTION

Recent advances in the experimental physics of cold atomic gases led to the observation of new regimes of their spin dynamics, for both bosons [1,2] and fermions [3–7]. One of the most interesting features of these systems is a synthetic optically produced (pseudo)spin-orbit coupling playing the critical role there. The qualitative effects of this coupling seen in cold matter are strongly different from the effects observed in solids [8–11]. Recently, it became possible to realize either optically [12] or by a mechanical rotation [13] the synthetic magnetic fields influencing the orbital motion of cold atomic gases. Comprehensive reviews of the field can be found in Ref. [14].

A spin-orbit coupled three-dimensional system of particles with pseudospin 1/2 can be described by the Hamiltonian [15]

$$H = \int d^3r \frac{1}{2m} \Psi^\dagger (-i\partial_j - A_j - A_j^{\text{so}})^2 \Psi + \Psi^\dagger A^0 \Psi + W[\Psi^\dagger, \Psi], \quad (1)$$

where m is the particle mass, $\Psi^\dagger$ and Ψ are the particle-related two-component spinor field operators, and $W[\Psi^\dagger, \Psi]$ is determined by the spin-independent external potential and interaction between the particles. The spin-independent components of the gauge field A_j are due to the $U(1)$ synthetic magnetic field $\mathbf{B} = \nabla \times \mathbf{A}$. The spin-orbit coupling is represented here by generally a non-Abelian $SU(2)$ gauge field $(A^0, A_1^{\text{so}}, A_2^{\text{so}}, A_3^{\text{so}})$, where each component is a 2×2 matrix, the A^0 component corresponds to the Zeeman coupling, and spatial components describe the spin-orbit coupling. Thus, we consider a cold matter in a background gauge field with the $SU(2) \otimes U(1)$ symmetry. This general approach to the spin-orbit coupling in terms of an $SU(2)$ field leads to a deep understanding of the spin-related properties of condensed matter [16–22]. Here, the vector potential for the linear in the particle-momentum spin-orbit coupling has the form

$$A_j^{\text{so}} = \frac{1}{2} q_j (\mathbf{h}^{[j]} \cdot \boldsymbol{\sigma}), \quad (2)$$

where q_j is the momentum component determined by the spin-orbit coupling strength (with a typical value of the order of 10^4 cm^{-1}), $\mathbf{h}^{[j]}$ is the direction of the corresponding spin-orbit coupling field, and $\boldsymbol{\sigma}$ is the Pauli matrix vector. We consider

j -independent $\mathbf{h} \equiv \mathbf{h}^{[j]}$ corresponding to the single-particle spin-orbit coupling Hamiltonian in the form

$$H_{\text{so}} = \frac{i}{2m} q_j \partial_j (\mathbf{h} \cdot \boldsymbol{\sigma}), \quad (3)$$

such that the spin precession angle caused by the spin-orbit coupling depends only on the particle displacement along the $\mathbf{q}$ direction and the coupling can be gauged out by the corresponding coordinate-dependent spin rotation around the $\mathbf{h}$ axis. This circumstance allows the mapping of the spin dynamics onto the spin density evolution in real space in the absence of spin-orbit coupling, making it directly related to the generalized diffusion picture. Such a realization of spin-orbit coupling has recently been produced for cold three-dimensional Fermi gases of ^{40}K [5] and ^6Li [6] and corresponds to the formation of spin helix states [23–25]. For cold Fermi gases, q and the Fermi momentum p_F are of the same order of magnitude since they both are determined by $2\pi/\lambda$, where λ is the wavelength of an optical photon [5,6,26].

There are several ways to produce a nonequilibrium spin polarization and track the subsequent spin dynamics. One option is to excite a spin density in a certain region of the momentum space, as it is done in cold fermions [5–7] and in solids [27], and then to study its evolution. Another option is to produce a nonequilibrium spin polarization by an external, a uniform, or a weakly coordinate-dependent Zeeman-like field, let the system reach equilibrium in the applied field, and then switch this field off. We follow the second approach and consider spin polarization initially produced by a weak Zeeman δA^0 term, switched off at $t = 0$. Its subsequent evolution is determined by the spin diffusion kernel in the momentum space in the absence of spin-orbit coupling, however, taken at the momentum $\mathbf{q}$ in Eq. (2) [28,29]. With the spin-orbit coupling in Eq. (2), the spin projection at the $\mathbf{h}$ axis is conserved, and we will consider evolution of the orthogonal to $\mathbf{h}$ spin component, which we denote as $S(t)$. The time dependence of the spin can be presented in the form

$$\frac{S(t)}{S(0)} = \frac{\delta s_{\mathbf{q}}(t)}{\delta s_{\mathbf{q}}(0)}, \quad (4)$$

where $\delta s_{\mathbf{q}}(t)$ is the corresponding component of the spin density evolving as in the absence of the spin-orbit coupling. In other words, the coordinate-spin mapping can be expressed

in terms of the exact frequency and momentum-dependent diagonal spin susceptibility $\chi_{\sigma\sigma}(\mathbf{q}, \omega)$ as

$$\frac{S(t)}{S(0)} = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \left[\frac{\chi_{\sigma\sigma}(\mathbf{q}, \omega)}{\chi_{\sigma\sigma}(\mathbf{q}, 0)} - 1 \right] \frac{e^{-i\omega t}}{i\omega}. \quad (5)$$

This approach is valid for both fermionic and bosonic systems in the absence of a steady spin-related magnetization. Here we apply this mapping in the form of Eqs. (4) and (5) to cold gases in the $SU(2) \otimes U(1)$ -symmetric fields and study the qualitative features brought about by the orbital motion in the $U(1)$ pseudomagnetic field.

This paper is organized as follows. In Sec. II we introduce for three-dimensional systems the collective coordinate variables related to the motion of a particle with the momentum corresponding to the spin-orbit coupling in a synthetic magnetic field. This allows us to formulate the corresponding linear response theory to calculate $\delta s_{\mathbf{q}}(t)$ and to study the spin dynamics. This approach is applicable at any strength of the spin-orbit coupling. In Sec. III it will be applied to the spin dynamics, both in the reversible ballistic and irreversible collision-dominated regimes. The results will be summarized in Conclusions. In the Appendixes we show how the same results in the collisionless limit can be obtained by considering the real single-particle trajectories and briefly provide the results for two-dimensional systems.

II. DENSITY MATRIX AND AUXILIARY TRAJECTORIES

To introduce the approach based on the auxiliary trajectories determined by spin-orbit coupling, we consider an atomic gas in a uniform synthetic “magnetic” field parallel to the z axis and characterized by $\omega_c = \mathbf{z}\omega_c$, where $\omega_c = B/m$ is the cyclotron frequency or the macroscopic rotation frequency [13], as shown in Fig. 1. In the following we adopt the Landau gauge to describe the $U(1)$ magnetic field. Our aim is to calculate the spin response function $\chi_{\sigma\sigma}(\mathbf{q}, \omega)$ entering

FIG. 1. Synthetic magnetic field $\mathbf{B}$, the spin-orbit coupling momentum $\mathbf{q}$, and the corresponding spin-split Fermi surfaces with the Fermi momentum p_F . Vector $\mathbf{h}$ is the direction of the spin-orbit coupling field in Eq. (3) and σ shows the equilibrium magnetization of the components of the spin-split Fermi gas. This picture corresponds to ultrastrong coupling with $q > p_F$.

Eq. (5) by solving the equation of motion for the one-particle spin-density matrix $\rho(\mathbf{r}_1, \mathbf{r}_2, t)$. It is convenient to introduce the $U(1)$ gauge invariant density matrix as follows:

$$f(\mathbf{r}, \mathbf{R}) = \exp\left(i \frac{\mathbf{y} \cdot \mathbf{x}}{l_B^2}\right) \rho(\mathbf{r}, \mathbf{R}), \quad (6)$$

where $\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$ and $\mathbf{R} = (\mathbf{r}_1 + \mathbf{r}_2)/2$ are the relative and the center-of-mass coordinates, $\rho(\mathbf{r}, \mathbf{R})$ is the density matrix in the Landau gauge, and l_B is the corresponding magnetic length. The spin density is related to the diagonal element of the density matrix $\mathbf{s}(\mathbf{R}) = \text{Tr}[\sigma f(0, \mathbf{R})]$.

We consider first a collisionless, purely ballistic dynamics of noninteracting gas. This regime of dynamics is realized when the spin precession rate $\langle v \rangle q$ and/or the cyclotron frequency are larger than the collisional relaxation rate $1/\tau$ (here $\langle v \rangle$ is the characteristic velocity of particles). The equation of motion for the density matrix in the presence of external coordinate- and time-dependent Zeeman-like perturbation $\delta \mathcal{A}^0(\mathbf{R}, t)$, corresponding to the field orthogonal to the $\mathbf{h}$ axis, becomes

$$\left[i\partial_t - \left(\frac{\hat{P}_i}{m} + \omega_c \varepsilon_{zij} r_i \right) \hat{p}_j \right] f(\mathbf{r}, \mathbf{R}) = \left[\delta \mathcal{A}^0\left(\mathbf{R} + \frac{\mathbf{r}}{2}, t\right) - \delta \mathcal{A}^0\left(\mathbf{R} - \frac{\mathbf{r}}{2}, t\right) \right] f(\mathbf{r}, \mathbf{R}), \quad (7)$$

where $\hat{P}_i = -i\partial/\partial R_i$, $\hat{p}_j = -i\partial/\partial r_j$, and ε_{zij} is the corresponding component of the Levi-Civita fully antisymmetric tensor.

In the linear response regime the density matrix is weakly perturbed from its equilibrium form $f_0(\mathbf{r})$:

$$f(\mathbf{r}, \mathbf{R}) = f_0(\mathbf{r}) + \delta f(\mathbf{r}, \mathbf{R}). \quad (8)$$

The linearized version of the equation of motion reads

$$\left[i\partial_t - \left(\frac{\hat{P}_i}{m} + \omega_c \varepsilon_{zij} r_i \right) \hat{p}_j \right] \delta f(\mathbf{r}, \mathbf{R}) = \left[\delta \mathcal{A}^0\left(\mathbf{R} + \frac{\mathbf{r}}{2}, t\right) - \delta \mathcal{A}^0\left(\mathbf{R} - \frac{\mathbf{r}}{2}, t\right) \right] f_0(\mathbf{r}). \quad (9)$$

Since we are interested in the $\mathbf{q}$ -dependent response it is sufficient to consider a plane-wave perturbation

$$\delta \mathcal{A}^0(\mathbf{r}, t) = \delta \mathcal{A}_{\mathbf{q}}^0(t) e^{i\mathbf{q} \cdot \mathbf{r}}. \quad (10)$$

This perturbation induces the response of the form $\delta f(\mathbf{r}, \mathbf{R}, t) = \delta f_{\mathbf{q}}(\mathbf{r}, t) e^{i\mathbf{q} \cdot \mathbf{r}}$. By substituting $\delta \mathcal{A}^0(\mathbf{r}, t) = \mathcal{A}_{\mathbf{q}}^0(t) e^{i\mathbf{q} \cdot \mathbf{r}}$ and $\delta f(\mathbf{r}, \mathbf{R}, t) = \delta f_{\mathbf{q}}(\mathbf{r}, t) e^{i\mathbf{q} \cdot \mathbf{r}}$ into Eq. (9) we obtain the following final equation of motion:

$$\left[i\partial_t - \left(\frac{q_i}{m} + \omega_c \varepsilon_{zij} r_i \right) \frac{\partial}{\partial r_j} \right] \delta f_{\mathbf{q}}(\mathbf{r}, t) = 2\delta \mathcal{A}_{\mathbf{q}}^0(t) \sin\left[\frac{\mathbf{q} \cdot \mathbf{r}}{2}\right] f_0(\mathbf{r}). \quad (11)$$

The left-hand side of this equation is the Boltzmann operator, corresponding to the helix-like motion of particles having velocities $\mathbf{q}/m$ in a magnetic field. The right-hand side of this equation plays the role of the source of these particles [30]. To understand the role of the q_i parameters in helical

motion, we introduce the anomalous spin-dependent term in the velocity operator

$$\mathbf{v}_{\text{so}} \equiv i[H_{\text{so}}, \mathbf{r}] = -\frac{\mathbf{q}}{2m}(\mathbf{h} \cdot \boldsymbol{\sigma}), \quad (12)$$

and find that $\mathbf{q}/m$ is the difference of the eigenvalues of $\mathbf{v}_{\text{so}}$ corresponding to $\boldsymbol{\sigma}$ antiparallel and parallel to $\mathbf{h}$.

To solve Eq. (11) we introduce auxiliary trajectories $\boldsymbol{\xi}(\mathbf{r}, t)$, satisfying the following equation of motion:

$$\dot{\xi}_j(\mathbf{r}, t) = \frac{q_j}{m} + \omega_c \varepsilon_{zij} \xi_i(\mathbf{r}, t), \quad (13)$$

with the initial condition $\xi_j(\mathbf{r}, t) = r_j$. The trajectory $\boldsymbol{\xi}(\mathbf{r}, t)$ corresponds to a particle moving from the point $\mathbf{r}$ with the initial momentum $\mathbf{q}$ in the presence of a given magnetic field. By a direct substitution one easily checks that the solution to Eq. (11) can be expressed as

$$\delta f_q(\mathbf{r}, t) = 2 \int_{-\infty}^0 dt_1 \delta A_q^0(t_1 + t) \times \sin\left[\frac{\mathbf{q} \cdot \boldsymbol{\xi}(\mathbf{r}, t_1)}{2}\right] f_0(\boldsymbol{\xi}(\mathbf{r}, t_1)). \quad (14)$$

Taking into account that the spin density response corresponds to the density matrix at $\mathbf{r} = 0$, i.e., $\delta s_q(t) = \delta f_q(0, t)$, we obtain

$$\delta s_q(t) = 2 \int_{-\infty}^0 dt_1 A_q^0(t_1 + t) \sin\left[\frac{\mathbf{q} \cdot \boldsymbol{\xi}(t_1)}{2}\right] f_0(\boldsymbol{\xi}(t_1)), \quad (15)$$

where we introduced the notation $\boldsymbol{\xi}(t_1) \equiv \boldsymbol{\xi}(0, t_1)$.

Without loss of generality the spin-orbit coupling momentum can be taken as $\mathbf{q} = (q_{\perp}, 0, q_{\parallel})$, where $q_{\parallel} = q \cos \theta$ and $q_{\perp} = q \sin \theta$ with θ being the angle between the direction of $\mathbf{q}$ and the magnetic field as shown in Fig. 1. In this parametrization the auxiliary helix trajectories take the following explicit form:

$$\begin{aligned} \xi_x(t) &= \frac{q_{\perp}}{m\omega_c} \sin(\omega_c t), & \xi_y(t) &= \frac{q_{\perp}}{m\omega_c} [1 - \cos(\omega_c t)], \\ \xi_z(t) &= \frac{q_{\parallel}}{m} t, \\ \xi(t) &= \sqrt{4(R_q^{\perp})^2 \sin^2 \frac{\omega_c t}{2} + \left(\frac{q_{\parallel}}{m} t\right)^2}, \end{aligned} \quad (16)$$

where $R_q^{\perp} = q_{\perp}/m\omega_c$ and

$$\mathbf{q} \cdot \boldsymbol{\xi}(t) = \frac{q_{\perp}^2}{m\omega_c} \sin(\omega_c t) + \frac{q_{\parallel}^2}{m} t. \quad (17)$$

Up to this point we specify neither statistics of the particles nor the temperature. All this information is encoded in the equilibrium density matrix $f_0(\mathbf{r})$. Below we consider the experimentally relevant quasiclassical limit with respect to the U(1) magnetic field, that is, $\omega_c \ll \langle E \rangle$, where $\langle E \rangle$ is the characteristic particle energy. In this limit $f_0(\mathbf{r})$ reduces to the Fourier transform of the equilibrium distribution function $f(p)$

$$f(p) = \frac{1}{\exp[(p^2/2m - \mu)/T] \pm 1}, \quad (18)$$

where $\pm$ corresponds to the Fermi or Bose statistics of the particles.

III. COLLISIONLESS MOTION AND SPIN RELAXATION

A. General solution

The special feature of a three-dimensional gas is that for a nonzero $q_{\parallel}$ the trajectories $\boldsymbol{\xi}(\mathbf{r}, t)$ are open, and a particle can move infinitely far from the initial position, making the behavior qualitatively different from that expected in a two-dimensional gas. For the subsequent discussion of spin dynamics it is convenient to represent the Fourier transform of the distribution function in the following form:

$$f_0(\mathbf{r}) = \int \frac{d^3 p}{(2\pi)^3} e^{i\mathbf{p} \cdot \mathbf{r}} f(p) = -\frac{1}{\pi r} \frac{d}{dr} \mathcal{N}(r), \quad (19)$$

where

$$\mathcal{N}(r) \equiv \int_0^{\infty} \cos(pr) f(p) \frac{dp}{2\pi}. \quad (20)$$

The function $\mathcal{N}(\xi(t))$ will play an important role in the following analysis. First of all we notice that $\lim_{t \rightarrow \infty} \mathcal{N}(\xi(t))_{q_{\parallel} \neq 0} = 0$, which follows from the openness of the particle trajectories. We also show the functions $\mathcal{N}(\xi(t))$ for two limiting cases of interest. For a degenerate Fermi gas with $f(p) = \theta(p_F - p)$ we find

$$\mathcal{N}(\xi(t)) = \int \frac{dp}{2\pi} \theta(p_F - p) \cos[p\xi(t)] = \frac{1}{2\pi} \frac{\sin[(p_F \xi(t))]}{\xi(t)}. \quad (21)$$

For highly degenerate Bose gases slightly above the Bose-Einstein condensation transition, where the chemical potential $-\mu$ is much less than the temperature T , the low-momentum part of the distribution function reduces to the form $f(p) = 2p_T^2/(p^2 + p_{\mu}^2)$, where $p_{\mu}^2 \equiv 2m|\mu| \ll p_T^2$ and $p_T^2 \equiv mT$. In this limit we obtain

$$\mathcal{N}(\xi(t)) = \int \frac{dp}{2\pi} \frac{2p_T^2}{p^2 + p_{\mu}^2} \cos[p\xi(t)] = \frac{p_T^2}{2p_{\mu}} \exp[-p_{\mu} \xi(t)]. \quad (22)$$

Note that Eq. (22), based on the small-momentum low- T form of the Bose distribution, is exact only at sufficiently large displacements such that $p_T \xi(t) \gg 1$ [31].

Now we proceed further with the transformation of the density matrix $f_0(\boldsymbol{\xi}(t))$ entering Eq. (15). Using the trajectory equation (13) we find the following identity:

$$\frac{d}{dt} \xi(t) = \frac{\dot{\boldsymbol{\xi}}(t) \cdot \boldsymbol{\xi}(t)}{\xi(t)} = \frac{\mathbf{q} \cdot \dot{\boldsymbol{\xi}}(t)}{m\xi(t)}. \quad (23)$$

Making use of this identity we obtain

$$\frac{d\mathcal{N}(\xi(t))}{d\xi(t)} = \frac{m\xi(t)}{\mathbf{q} \cdot \dot{\boldsymbol{\xi}}(t)} \frac{d\mathcal{N}(\xi(t))}{dt}.$$

Finally, using the representation of Eq. (19) we find for the equilibrium density matrix evaluated at the particle trajectory

$$f_0(\boldsymbol{\xi}(t)) = -\frac{1}{\pi \xi(t)} \frac{d\mathcal{N}(\xi(t))}{d\xi(t)} = -\frac{m}{\pi} \frac{1}{\mathbf{q} \cdot \dot{\boldsymbol{\xi}}(t)} \frac{d\mathcal{N}(\xi(t))}{dt}. \quad (24)$$

As a result, the spin density response to the perturbation $\mathcal{A}_q^0(t_1)$ takes the form

$$\delta s_q(t) = \frac{2m}{\pi} \int_{-\infty}^0 dt_1 \delta \mathcal{A}_q^0(t_1 + t) \times \frac{\sin\left[\frac{\mathbf{q} \cdot \xi(t_1)}{2}\right]}{\mathbf{q} \cdot \xi(t_1)} \frac{d\mathcal{N}(\xi(t_1))}{dt_1}. \quad (25)$$

Below we apply this equation to study the reversible collisionless spin evolution and spin relaxation in the collision-dominated regime.

B. Collisionless spin dynamics as switch-off response

In general to find the spin evolution from Eq. (5) one needs the spin response function $\chi_{\sigma\sigma}(\mathbf{q}, \omega)$. The latter requires calculation of the response to a time periodic perturbation $\sim e^{-i\omega t}$. However, the reversible collisionless spin dynamics can be extracted directly from Eq. (25) by considering the switch-off process with the driving field of the following form:

$$\delta \mathcal{A}_q^0(t) = \exp(\delta t) \theta(-t), \quad (26)$$

where the limit $\delta \rightarrow +0$ is assumed. In other words, the Zeeman field is switched on adiabatically to polarize the system (in the equilibrium at $t < 0$), and then switched off at $t = 0$ and follows the dynamics for $t > 0$. The corresponding spin density response acquires the form

$$\delta s_q(t) = \frac{2m}{\pi} \int_t^\infty dt_1 \frac{\sin\left[\frac{\mathbf{q} \cdot \xi(t_1)}{2}\right]}{\mathbf{q} \cdot \xi(t_1)} \frac{d\mathcal{N}(\xi(t_1))}{dt_1}. \quad (27)$$

The characteristic scale of the function $\mathcal{N}(p)$ is determined by the mean velocity $\langle v \rangle$. This allows the further simplification of Eq. (27) in the limit $\langle v \rangle \gg q/m$. Integrating by parts and separating the fast motion we obtain the approximate expression, formally valid for small spin-orbit momentum q :

$$\frac{S(t)}{S(0)} = 2 \frac{\sin\left[\frac{\mathbf{q} \cdot \xi(t)}{2}\right]}{\mathbf{q} \cdot \xi(t)} \frac{\mathcal{N}(\xi(t))}{\mathcal{N}(0)}. \quad (28)$$

The results of calculation with the exact Eq. (27) are presented in Fig. 2 for Fermi and Bose statistics. For θ close to $\pi/2$ one expects revivals of the spin density after a cyclotron period: $S(2\pi n/\omega_c) = S(0)$. These revivals, however, are strongly suppressed at smaller angles by the motion along the magnetic field, as can be seen in Fig. 2.

It is interesting to note that although the approximation of Eq. (28) is formally valid in the small- q limit, it describes the spin evolution highly accurately even for $q \sim m\langle v \rangle$. Therefore the main features of the reversible dynamics can be understood from the analytic formula (28). Let us consider first the degenerate Fermi gas with $\mathcal{N}(\xi)$ given by Eq. (21) and substitute the trajectory of Eq. (16). Then Eq. (28) acquires the following explicit form:

$$\frac{S(t)}{S(0)} = 2 \frac{\sin\left[p_F \sqrt{4(R_q^\perp)^2 \sin^2 \frac{\omega_c t}{2} + \left(\frac{q_\parallel}{m} t\right)^2}\right]}{p_F \sqrt{4(R_q^\perp)^2 \sin^2 \frac{\omega_c t}{2} + \left(\frac{q_\parallel}{m} t\right)^2}} \times \frac{\sin\left[\frac{1}{2}\left(\frac{q_\perp^2}{m\omega_c} \sin(\omega_c t) + \frac{q_\parallel^2}{m} t\right)\right]}{\frac{q_\perp^2}{m\omega_c} \sin(\omega_c t) + \frac{q_\parallel^2}{m} t}. \quad (29)$$

FIG. 2. Dependence of the dimensionless ratio $S(t)/S(0)$ for the total spin calculated with exact Eqs. (4) and (27) on the dimensionless product $\omega_c t$ for different angles θ (see Fig. 1) as marked near the plots. (a) For fermions with the Fermi momentum p_F , spin-orbit coupling $q = p_F/2$, and cyclotron $\omega_c = E_F/8$. (b) For bosons with the momentum p_μ , spin-orbit coupling $q = p_\mu/2$, and $\omega_c = |\mu|/8$. An approximation with Eq. (28) (not shown here) describes all evolutions very accurately. The insets show the time dependences for $\theta = \pi/4$.

This expression shows that the initial stage is well described in the limit of a straight motion with $\xi(t) = qt/m$, producing a universal evolution which is ω_c and θ independent. Taking into account the short-time expansion $qt/m - \xi(t) = qt/m(\omega_c t)^2 \sin^2 \theta/24$, we obtain that the cyclotron motion modifies this universal behavior at times $q(p_F/m)t \sin^2 \theta(\omega_c t)^2/24 \sim 1$, that is, $t \sim 3/\omega_c(\sin^2 \theta p_F q/m\omega_c)^{1/3}$. Note that if the angle θ is sufficiently small, that is, $\theta \ll (m\omega_c/p_F q)^{1/2}$, the cyclotron motion has only a weak influence on the spin dynamics since the effective displacement along the field direction is always sufficiently larger than the cyclotron radius. Another time scale of the problem is determined by the second, p_F independent factor in Eq. (29). At $\theta = 0$, it leads to additional decay factor in the spin density as $\sin(q_\parallel^2 t/2m)/(q_\parallel^2 t/m)$, while at $\theta = \pi/2$ it leads to a modulation of the $\sin(p_F q t)/p_F q t$ decay. The revivals show a moderate suppression by the along-the-field motion as $\sin(p_F n q_\parallel t_c)/p_F n q_\parallel t_c$.

The same arguments can be applied to the low-temperature Bose gas [32] with $\mathcal{N}(\xi)$ of Eq. (22). In this case Eq. (28)

reduces to the form

$$\frac{S(t)}{S(0)} = 2 \exp \left(-p_\mu \sqrt{4(R_q^\perp)^2 \sin^2 \frac{\omega_c t}{2} + \left(\frac{q_\parallel}{m} t \right)^2} \right) \times \frac{\sin \left[\frac{1}{2} \left(\frac{q_\perp^2}{m\omega_c} \sin(\omega_c t) + \frac{q_\parallel^2}{m} t \right) \right]}{\frac{q_\perp^2}{m\omega_c} \sin(\omega_c t) + \frac{q_\parallel^2}{m} t}. \quad (30)$$

Because of a broad distribution of the momenta, the revivals at full cyclotron periods here are suppressed much stronger than for the Fermi gas, that is, exponentially as $\exp(-np_\mu q_\parallel t_c)$, which is clearly visible in Fig. 2(b).

C. Susceptibility, diffusion, and spin relaxation in collision-dominated limit

Equation (27) describes collisionless reversible spin dynamics which originate from the inhomogeneous spin precession of particles moving ballistically along helicoidal trajectories in the magnetic field. Now we consider the opposite limit of collision-dominated diffusive dynamics corresponding to the situation when the spin precession rate is smaller than the rate of collisional relaxation $q(v) \ll 1/\tau$. To take into account the effect of collisions we apply our general formula, Eq. (5). For the response function $\chi_{\sigma\sigma}(\mathbf{q}, \omega)$ we use the density-conserving relaxation time approximation that has been introduced by Mermin [33] and adapted recently [34] to the problems of spin dynamics. In this approximation the spin response function is calculated as

$$\chi_{\sigma\sigma}(\mathbf{q}, \omega) = \frac{\chi_0(\mathbf{q}, \omega + i/\tau)}{1 - \frac{1}{1-i\omega\tau} \left[1 - \frac{\chi_0(\mathbf{q}, \omega + i/\tau)}{\chi_0(\mathbf{q}, 0)} \right]}, \quad (31)$$

where $\chi_0(\mathbf{q}, \omega)$ is a bare susceptibility of the system in the absence of collisions. The latter quantity is obtained straightforwardly from Eq. (25) by considering the response to a time-periodic Zeeman field of the form $\delta\mathcal{A}_\mathbf{q}^0(t) = \delta\mathcal{A}_\mathbf{q}^0 e^{-i\omega t}$:

$$\chi_0(\mathbf{q}, \omega) = \frac{2m}{\pi} \int_0^\infty dt e^{i\omega t} \frac{\sin \left[\frac{\mathbf{q} \cdot \xi(t)}{2} \right]}{\mathbf{q} \cdot \xi(t)} \frac{d\mathcal{N}(\xi(t))}{dt}. \quad (32)$$

Equations (5), (31), and (32) describe the spin dynamics for any q, ω_c , and τ , provided the magnetic field can be treated semiclassically.

In the collision-dominated regime, only small frequencies $\omega \ll 1/\tau$ contribute to the integral in Eq. (5). In this limit the integrand in Eq. (5) acquires a simple pole on the imaginary axis

$$\frac{1}{\omega} \left[\frac{\chi_{\sigma\sigma}(\mathbf{q}, \omega)}{\chi_{\sigma\sigma}(\mathbf{q}, 0)} - 1 \right] \approx \frac{1}{\omega + i\Omega(\mathbf{q})}, \quad (33)$$

$$\Omega(\mathbf{q}) = \frac{1}{\tau} \frac{\chi_0(\mathbf{q}, i/\tau)}{\chi_0(\mathbf{q}, 0)}, \quad (34)$$

which is a clear signature of the diffusive dynamics [35,36]. To find the position of the diffusive pole we notice that the collision-dominated regime occurs only in the small- q limit. Therefore a simplified form of the bare susceptibility [Eq. (32)]

can be used as

$$\begin{aligned} \chi_0(\mathbf{q}, \omega) &= \frac{m}{\pi} \int_0^\infty dt e^{i\omega t} \frac{d\mathcal{N}(\xi(t))}{dt} \\ &= -\frac{m}{\pi} \mathcal{N}(0) - i\omega \frac{m}{\pi} \int_0^\infty dt e^{i\omega t} \mathcal{N}(\xi(t)). \end{aligned} \quad (35)$$

Using this expression and the definition of $\mathcal{N}(\xi(t))$ in Eq. (20) we find that $\chi_0(\mathbf{q}, 0)$ can be expressed in terms of the derivative of the density n of particles with respect to the chemical potential,

$$\chi_0(\mathbf{q}, 0) = -\frac{\partial n}{\partial \mu}, \quad (36)$$

in agreement with the compressibility sum rule [37] corresponding to an adiabatic modification of the density by a smooth external potential. By substituting $\omega = i/\tau$ into Eq. (35) and expanding $\mathcal{N}(\xi)$ to the second order in ξ we compute $\chi_0(\mathbf{q}, i/\tau)$ in the diffusive limit

$$\begin{aligned} \chi_0(\mathbf{q}, i/\tau) &= -\frac{mn}{2\tau} \int_0^\infty \exp(-t/\tau) \xi^2(t) dt \\ &= -\frac{\tau^2 n}{m} \left(q_\parallel^2 + \frac{q_\perp^2}{1 + \omega_c^2 \tau^2} \right). \end{aligned} \quad (37)$$

In the second line of this equation we used the explicit form of $\xi(t)$ for the helix trajectory, Eq. (16). The anisotropic $\mathbf{q}$ dependence reflects the fact that at long times the circular motion is restricted as $\xi_\perp^2 \leq 4(R_q^\perp)^2 \sim (q_\perp/m\omega_c)^2$, and, therefore, cannot give a large contribution into the spin evolution, while the parallel motion with $\xi_\parallel^2 = (q_\parallel/m)^2 t^2$ is unrestricted and influences this evolution much stronger.

Equations (34), (36), and (37) determine the position of the diffusive pole in the response function

$$\Omega(\mathbf{q}) = D \left(q_\parallel^2 + \frac{q_\perp^2}{1 + \omega_c^2 \tau^2} \right), \quad (38)$$

where $D = \tau n(\partial\mu/\partial n)/m$ is the diffusion coefficient. Finally, Eqs. (5) and (33) yield the modified Dyakonov-Perel spin relaxation [38–40] dependent on $\mathbf{q}$ direction as follows:

$$\frac{S(t)}{S(0)} = \exp \left[-D \left(q_\parallel^2 + \frac{q_\perp^2}{1 + \omega_c^2 \tau^2} \right) t \right]. \quad (39)$$

Importantly, the statistics of the particles and the temperature enter this equation only through the diffusion coefficient. Qualitatively in the collision-dominated regime the behavior of the spin is universal [41]. For $q_\perp = 0$, the magnetic field does not influence the particle trajectory, and we return to the conventional Dyakonov-Perel relaxation in zero magnetic field. In the case of a strong magnetic field and $q_\parallel = 0$, the relaxation rate can be understood as follows. Particles move over circular trajectories of a small radius $R_c = v/\omega_c$ with $R_c \ll 1/q$. The trajectory relatively rarely (after the collision time $\tau \gg 1/\omega_c$) experiences considerable changes by displacement of the order of R_c . Therefore, the guiding center of the trajectory diffuses with characteristic time-dependent displacement $\rho_{gc}^2(t) \sim R_c^2(t/\tau)$. Taking into account that $R_c^2 \sim D/\omega_c^2 \tau$, the condition $\rho_{gc}^2(t_{sr}) \sim 1/q^2$ [28] yields the spin relaxation time t_{sr} in agreement with Eq. (39).

IV. CONCLUSIONS

We have studied (pseudo)spin evolution in cold atomic gases, fermionic and bosonic, in synthetic $SU(2) \otimes U(1)$ fields. Our approach, based on the representation of motion in terms of auxiliary trajectories satisfying the kinetic equation, utilizes the mapping of spin dynamics onto the spin density evolution in systems without spin-orbit coupling. This mapping is possible for the considered spin-orbit coupling determined by a single momentum-independent precession axis. With this mapping, valid for any strength of the spin-orbit coupling, we describe reversible collisionless evolution as well as irreversible spin relaxation in the collision-dominated limit.

The collisionless limit shows the dependence of the reversible spin evolution on the angle between the synthetic $U(1)$ field $\mathbf{B}$ and the single momentum $\mathbf{q}$ characterizing the $SU(2)$ coupling. For the angle between this spin-orbit coupling momentum and the synthetic magnetic field close to $\pi/2$, the total spin shows periodic behavior with returns after each cyclotron period. The periodicity disappears at smaller angles due to the motion of atoms along the $\mathbf{B}$ field. At low temperatures, the spin dynamics of Fermi gases tends to the limit determined by the Fermi-momentum, spin-orbit coupling, and the strength of the synthetic magnetic field. At a sufficiently strong spin-orbit coupling, this evolution can show an oscillating behavior determined by the spin precession. For the Bose statistics the oscillations are smeared out by the thermal Bose distribution. In addition, the spin evolution of bosons vanishes due to accumulation of the particles in the vicinity of the zero-energy point at temperatures slightly above the Bose-Einstein condensation temperature.

In the collision-dominated limit the spin relaxation described by the anisotropic Dyakonov-Perel formula depends on the angle between the momentum corresponding to the spin-orbit coupling and the synthetic magnetic field. Here the spin relaxation rate is proportional to the density diffusion coefficient. At a given momentum relaxation time, the diffusion coefficient for fermions reaches a finite zero-temperature limit proportional to the Fermi energy. However, for bosons this diffusion coefficient decreases to zero, and the spin relaxation slows down as the system approaches the Bose-Einstein condensation.

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APPENDIX A: REAL SINGLE-PARTICLE TRAJECTORIES FOR THREE-DIMENSIONAL SYSTEMS

Here we will show with an example that the same results for the collisionless dynamics can be obtained in the weak spin-orbit coupling limit by using the real rather than auxiliary helix trajectories for particles in a magnetic field. This trajectory for a particle with initial velocity $\mathbf{v} = (v_\perp, v_z)$, where $v_\perp = v \sin \theta$ and $v_z = v \cos \theta$ satisfies conditions $z = vt \cos \theta$ and

$\rho = 2R_c \sin \theta \sin |\omega_c t/2|$, where ρ is the distance from the origin. As a result, we obtain the collisionless limit for the diffusion kernel in the real space as

$$\mathcal{D}(\rho, z; t) = \delta[z - vt \cos \theta] \frac{\delta[\rho - 2R_c \sin \theta \sin |\frac{\omega_c t}{2}|]}{4\pi R_c \sin \theta |\sin \frac{\omega_c t}{2}|}, \quad (\text{A1})$$

satisfying the normalization condition

$$\int \mathcal{D}(\rho, z; t) d^2 \rho dz = 1. \quad (\text{A2})$$

According to Ref. [29], taking the Fourier transform of $\mathcal{D}(\rho, z; t)$ at $\mathbf{q} = (q_\perp, 0, q_\parallel)$ and averaging over the Fermi surface, that is, over the angles θ , we arrive at

$$\frac{S(t)}{S(0)} = \frac{\sin \sqrt{(q_\parallel vt)^2 + [2q_\perp R_c \sin(\frac{\omega_c t}{2})]^2}}{\sqrt{(q_\parallel vt)^2 + [2q_\perp R_c \sin(\frac{\omega_c t}{2})]^2}}, \quad (\text{A3})$$

same as Eq. (29) for $v = p_F/m$.

APPENDIX B: TWO-DIMENSIONAL GASES

The approach developed in this paper is valid for two-dimensional gases as well. For completeness, we present here the collisionless time dependence of spin for three main realizations of statistics while in the collision-dominated regime the relaxation is described by the modified Dyakonov-Perel formula.

For a degenerate Fermi gas the evolution

$$\begin{aligned} \frac{S(t)}{S(0)} &= \frac{2m\omega_c}{q^2 \sin(\omega_c t)} \sin \left[\frac{q^2 \sin(\omega_c t)}{2m\omega_c} \right] \\ &\times J_0 \left(2 \frac{p_F q}{m\omega_c} \sin \frac{\omega_c t}{2} \right), \end{aligned} \quad (\text{B1})$$

in the limit of a weak spin-orbit coupling and $U(1)$ field, shows a slower oscillating decrease ($\sim 1/\sqrt{t}$) than in the three-dimensional (3D) gas since in 3D the absolute values of the in-plane components of the Fermi momentum are distributed over the Fermi sphere between 0 and p_F while in 2D all particles have the same p_F . In the weak coupling limit this expression coincides with the result of Glazov in Ref. [42].

For a highly degenerate Bose gas close to the Bose-Einstein condensation one obtains

$$\begin{aligned} \frac{S(t)}{S(0)} &= \frac{2p_\mu}{q} \sin \left[\frac{q^2 \sin(\omega_c t)}{2m\omega_c} \right] \frac{1}{\cos(\omega_c t/2)} \\ &\times K_1 \left[2 \frac{qp_\mu}{m\omega_c} \sin \frac{\omega_c t}{2} \right], \end{aligned} \quad (\text{B2})$$

with K_1 being the Bessel function of the second kind.

For nondegenerate gases with $f(p) = \exp(-p^2/2p_T^2)$ the evolution is described by

$$\begin{aligned} \frac{S(t)}{S(0)} &= \frac{2m\omega_c}{q^2 \sin(\omega_c t)} \sin \left[\frac{q^2 \sin(\omega_c t)}{2m\omega_c} \right] \\ &\times \exp \left[-\frac{q^2 p_T^2}{2m^2 \omega_c^2} \sin^2 \frac{\omega_c t}{2} \right]. \end{aligned} \quad (\text{B3})$$

As expected, in all these 2D realizations the returns of the spin to the initial $S(0)$ value occur after each cyclotron period.

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