

## Branching fraction measurement of $B^+ \rightarrow \omega \ell^+ \nu$ decays

J. P. Lees, V. Poireau, and V. Tisserand

*Laboratoire d'Annecy-le-Vieux de Physique des Particules (LAPP),  
Université de Savoie, CNRS/IN2P3, F-74941 Annecy-Le-Vieux, France*

J. Garra Tico and E. Grauges

*Universitat de Barcelona, Facultat de Física, Departament ECM, E-08028 Barcelona, Spain*

A. Palano<sup>ab</sup>

*INFN Sezione di Bari<sup>a</sup>; Dipartimento di Fisica, Università di Bari<sup>b</sup>, I-70126 Bari, Italy*

G. Eigen and B. Stugu

*University of Bergen, Institute of Physics, N-5007 Bergen, Norway*

D. N. Brown, L. T. Kerth, Yu. G. Kolomensky, and G. Lynch

*Lawrence Berkeley National Laboratory and University of California, Berkeley, California 94720, USA*

H. Koch and T. Schroeder

*Ruhr Universität Bochum, Institut für Experimentalphysik 1, D-44780 Bochum, Germany*

D. J. Asgeirsson, C. Hearty, T. S. Mattison, J. A. McKenna, and R. Y. So

*University of British Columbia, Vancouver, British Columbia, Canada V6T 1Z1*

A. Khan

*Brunel University, Uxbridge, Middlesex UB8 3PH, United Kingdom*

V. E. Blinov, A. R. Buzykaev, V. P. Druzhinin, V. B. Golubev, E. A. Kravchenko, A. P. Onuchin,

S. I. Serebnyakov, Yu. I. Skovpen, E. P. Solodov, K. Yu. Todyshev, and A. N. Yushkov

*Budker Institute of Nuclear Physics, Novosibirsk 630090, Russia*

M. Bondioli, D. Kirkby, A. J. Lankford, and M. Mandelkern

*University of California at Irvine, Irvine, California 92697, USA*

H. Atmacan, J. W. Gary, F. Liu, O. Long, and G. M. Vitug

*University of California at Riverside, Riverside, California 92521, USA*

C. Campagnari, T. M. Hong, D. Kovalskyi, J. D. Richman, and C. A. West

*University of California at Santa Barbara, Santa Barbara, California 93106, USA*

A. M. Eisner, J. Kroseberg, W. S. Lockman, A. J. Martinez, B. A. Schumm, and A. Seiden

*University of California at Santa Cruz, Institute for Particle Physics, Santa Cruz, California 95064, USA*

D. S. Chao, C. H. Cheng, B. Echenard, K. T. Flood, D. G. Hitlin, P. Ongmongkolkul, F. C. Porter, and A. Y. Rakitin

*California Institute of Technology, Pasadena, California 91125, USA*

R. Andreassen, Z. Huard, B. T. Meadows, M. D. Sokoloff, and L. Sun

*University of Cincinnati, Cincinnati, Ohio 45221, USA*

P. C. Bloom, W. T. Ford, A. Gaz, U. Nauenberg, J. G. Smith, and S. R. Wagner

*University of Colorado, Boulder, Colorado 80309, USA*

R. Ayad\* and W. H. Toki

*Colorado State University, Fort Collins, Colorado 80523, USA*

B. Spaan

*Technische Universität Dortmund, Fakultät Physik, D-44221 Dortmund, Germany*

K. R. Schubert and R. Schwierz

*Technische Universität Dresden, Institut für Kern- und Teilchenphysik, D-01062 Dresden, Germany*

D. Bernard and M. Verderi

*Laboratoire Leprince-Ringuet, Ecole Polytechnique, CNRS/IN2P3, F-91128 Palaiseau, France*

P. J. Clark and S. Playfer

*University of Edinburgh, Edinburgh EH9 3JZ, United Kingdom*

D. Bettoni<sup>a</sup>, C. Bozzi<sup>a</sup>, R. Calabrese<sup>ab</sup>, G. Cibinetto<sup>ab</sup>, E. Fioravanti<sup>ab</sup>, I. Garzia<sup>ab</sup>,

E. Luppi<sup>ab</sup>, M. Munerato<sup>ab</sup>, M. Negrini<sup>ab</sup>, L. Piemontese<sup>a</sup>, and V. Santoro<sup>a</sup>

*INFN Sezione di Ferrara<sup>a</sup>; Dipartimento di Fisica, Università di Ferrara<sup>b</sup>, I-44100 Ferrara, Italy*

R. Baldini-Ferrolì, A. Calcaterra, R. de Sangro, G. Finocchiaro,

P. Patteri, I. M. Peruzzi,<sup>†</sup> M. Piccolo, M. Rama, and A. Zallo

*INFN Laboratori Nazionali di Frascati, I-00044 Frascati, Italy*

R. Contri<sup>ab</sup>, E. Guido<sup>ab</sup>, M. Lo Vetere<sup>ab</sup>, M. R. Monge<sup>ab</sup>, S. Passaggio<sup>a</sup>, C. Patrignani<sup>ab</sup>, and E. Robutti<sup>a</sup>

*INFN Sezione di Genova<sup>a</sup>; Dipartimento di Fisica, Università di Genova<sup>b</sup>, I-16146 Genova, Italy*

B. Bhuyan and V. Prasad

*Indian Institute of Technology Guwahati, Guwahati, Assam, 781 039, India*

C. L. Lee and M. Morii

*Harvard University, Cambridge, Massachusetts 02138, USA*

A. J. Edwards

*Harvey Mudd College, Claremont, California 91711*

A. Adametz and U. Uwer

*Universität Heidelberg, Physikalisches Institut, Philosophenweg 12, D-69120 Heidelberg, Germany*

H. M. Lacker and T. Lueck

*Humboldt-Universität zu Berlin, Institut für Physik, Newtonstr. 15, D-12489 Berlin, Germany*

P. D. Dauncey

*Imperial College London, London, SW7 2AZ, United Kingdom*

P. K. Behera and U. Mallik

*University of Iowa, Iowa City, Iowa 52242, USA*

C. Chen, J. Cochran, W. T. Meyer, S. Prell, and A. E. Rubin

*Iowa State University, Ames, Iowa 50011-3160, USA*

A. V. Gritsan and Z. J. Guo

*Johns Hopkins University, Baltimore, Maryland 21218, USA*

N. Arnaud, M. Davier, D. Derkach, G. Grosdidier, F. Le Diberder, A. M. Lutz,

B. Malaescu, P. Roudeau, M. H. Schune, A. Stocchi, and G. Wormser

*Laboratoire de l'Accélérateur Linéaire, IN2P3/CNRS et Université Paris-Sud 11,*

*Centre Scientifique d'Orsay, B. P. 34, F-91898 Orsay Cedex, France*

D. J. Lange and D. M. Wright

*Lawrence Livermore National Laboratory, Livermore, California 94550, USA*

C. A. Chavez, J. P. Coleman, J. R. Fry, E. Gabathuler, D. E. Hutchcroft, D. J. Payne, and C. Touramanis  
*University of Liverpool, Liverpool L69 7ZE, United Kingdom*

A. J. Bevan, F. Di Lodovico, R. Sacco, and M. Sigamani  
*Queen Mary, University of London, London, E1 4NS, United Kingdom*

G. Cowan  
*University of London, Royal Holloway and Bedford New College, Egham, Surrey TW20 0EX, United Kingdom*

D. N. Brown and C. L. Davis  
*University of Louisville, Louisville, Kentucky 40292, USA*

A. G. Denig, M. Fritsch, W. Gradl, K. Griessinger, A. Hafner, and E. Prencipe  
*Johannes Gutenberg-Universität Mainz, Institut für Kernphysik, D-55099 Mainz, Germany*

R. J. Barlow,<sup>‡</sup> G. Jackson, and G. D. Lafferty  
*University of Manchester, Manchester M13 9PL, United Kingdom*

E. Behn, R. Cenci, B. Hamilton, A. Jawahery, and D. A. Roberts  
*University of Maryland, College Park, Maryland 20742, USA*

C. Dallapiccola  
*University of Massachusetts, Amherst, Massachusetts 01003, USA*

R. Cowan, D. Dujmic, and G. Sciolla  
*Massachusetts Institute of Technology, Laboratory for Nuclear Science, Cambridge, Massachusetts 02139, USA*

R. Cheaib, D. Lindemann, P. M. Patel, and S. H. Robertson  
*McGill University, Montréal, Québec, Canada H3A 2T8*

P. Biassoni<sup>ab</sup>, N. Neri<sup>a</sup>, F. Palombo<sup>ab</sup>, and S. Stracka<sup>ab</sup>  
*INFN Sezione di Milano<sup>a</sup>; Dipartimento di Fisica, Università di Milano<sup>b</sup>, I-20133 Milano, Italy*

L. Cremaldi, R. Godang,<sup>§</sup> R. Kroeger, P. Sonnek, and D. J. Summers  
*University of Mississippi, University, Mississippi 38677, USA*

X. Nguyen, M. Simard, and P. Taras  
*Université de Montréal, Physique des Particules, Montréal, Québec, Canada H3C 3J7*

G. De Nardo<sup>ab</sup>, D. Monorchio<sup>ab</sup>, G. Onorato<sup>ab</sup>, and C. Sciacca<sup>ab</sup>  
*INFN Sezione di Napoli<sup>a</sup>; Dipartimento di Scienze Fisiche,  
 Università di Napoli Federico II<sup>b</sup>, I-80126 Napoli, Italy*

M. Martinelli and G. Raven  
*NIKHEF, National Institute for Nuclear Physics and High Energy Physics, NL-1009 DB Amsterdam, The Netherlands*

C. P. Jessop, J. M. LoSecco, and W. F. Wang  
*University of Notre Dame, Notre Dame, Indiana 46556, USA*

K. Honscheid and R. Kass  
*Ohio State University, Columbus, Ohio 43210, USA*

J. Brau, R. Frey, N. B. Sinev, D. Strom, and E. Torrence  
*University of Oregon, Eugene, Oregon 97403, USA*

E. Feltresi<sup>ab</sup>, N. Gagliardi<sup>ab</sup>, M. Margoni<sup>ab</sup>, M. Morandin<sup>a</sup>,  
 M. Posocco<sup>a</sup>, M. Rotondo<sup>a</sup>, G. Simi<sup>a</sup>, F. Simonetto<sup>ab</sup>, and R. Stroili<sup>ab</sup>  
*INFN Sezione di Padova<sup>a</sup>; Dipartimento di Fisica, Università di Padova<sup>b</sup>, I-35131 Padova, Italy*

S. Akar, E. Ben-Haim, M. Bomben, G. R. Bonneaud, H. Briand, G. Calderini,  
 J. Chauveau, O. Hamon, Ph. Leruste, G. Marchiori, J. Ocariz, and S. Sitt  
*Laboratoire de Physique Nucléaire et de Hautes Energies,  
 IN2P3/CNRS, Université Pierre et Marie Curie-Paris6,  
 Université Denis Diderot-Paris7, F-75252 Paris, France*

M. Biasini<sup>ab</sup>, E. Manoni<sup>ab</sup>, S. Pacetti<sup>ab</sup>, and A. Rossi<sup>ab</sup>  
*INFN Sezione di Perugia<sup>a</sup>; Dipartimento di Fisica, Università di Perugia<sup>b</sup>, I-06100 Perugia, Italy*

C. Angelini<sup>ab</sup>, G. Batignani<sup>ab</sup>, S. Bettarini<sup>ab</sup>, M. Carpinelli<sup>ab</sup>,<sup>¶</sup> G. Casarosa<sup>ab</sup>, A. Cervelli<sup>ab</sup>, F. Forti<sup>ab</sup>,  
 M. A. Giorgi<sup>ab</sup>, A. Lusiani<sup>ac</sup>, B. Oberhof<sup>ab</sup>, E. Paoloni<sup>ab</sup>, A. Perez<sup>a</sup>, G. Rizzo<sup>ab</sup>, and J. J. Walsh<sup>a</sup>  
*INFN Sezione di Pisa<sup>a</sup>; Dipartimento di Fisica, Università di Pisa<sup>b</sup>; Scuola Normale Superiore di Pisa<sup>c</sup>, I-56127 Pisa, Italy*

D. Lopes Pegna, J. Olsen, A. J. S. Smith, and A. V. Telnov  
*Princeton University, Princeton, New Jersey 08544, USA*

F. Anulli<sup>a</sup>, R. Faccini<sup>ab</sup>, F. Ferrarotto<sup>a</sup>, F. Ferroni<sup>ab</sup>, M. Gaspero<sup>ab</sup>, L. Li Gioi<sup>a</sup>, M. A. Mazzoni<sup>a</sup>, and G. Piredda<sup>a</sup>  
*INFN Sezione di Roma<sup>a</sup>; Dipartimento di Fisica,  
 Università di Roma La Sapienza<sup>b</sup>, I-00185 Roma, Italy*

C. Bünger, O. Grünberg, T. Hartmann, T. Leddig, H. Schröder,\*\* C. Voss, and R. Waldi  
*Universität Rostock, D-18051 Rostock, Germany*

T. Adye, E. O. Olaiya, and F. F. Wilson  
*Rutherford Appleton Laboratory, Chilton, Didcot, Oxon, OX11 0QX, United Kingdom*

S. Emery, G. Hamel de Monchenault, G. Vasseur, and Ch. Yèche  
*CEA, Irfu, SPP, Centre de Saclay, F-91191 Gif-sur-Yvette, France*

D. Aston, D. J. Bard, R. Bartoldus, J. F. Benitez, C. Cartaro, M. R. Convery, J. Dingfelder, J. Dorfan,  
 G. P. Dubois-Felsmann, W. Dunwoodie, M. Ebert, R. C. Field, M. Franco Sevilla, B. G. Fulsom, A. M. Gabareen,  
 M. T. Graham, P. Grenier, C. Hast, W. R. Innes, M. H. Kelsey, P. Kim, M. L. Kocian, D. W. G. S. Leith, P. Lewis,  
 B. Lindquist, S. Luitz, V. Luth, H. L. Lynch, D. B. MacFarlane, D. R. Muller, H. Neal, S. Nelson, M. Perl,  
 T. Pulliam, B. N. Ratcliff, A. Roodman, A. A. Salnikov, R. H. Schindler, A. Snyder, D. Su, M. K. Sullivan,  
 J. Va'vra, A. P. Wagner, W. J. Wisniewski, M. Wittgen, D. H. Wright, H. W. Wulsin, C. C. Young, and V. Ziegler  
*SLAC National Accelerator Laboratory, Stanford, California 94309 USA*

W. Park, M. V. Purohit, R. M. White, and J. R. Wilson  
*University of South Carolina, Columbia, South Carolina 29208, USA*

A. Randle-Conde and S. J. Sekula  
*Southern Methodist University, Dallas, Texas 75275, USA*

M. Bellis, P. R. Burchat, and T. S. Miyashita  
*Stanford University, Stanford, California 94305-4060, USA*

M. S. Alam and J. A. Ernst  
*State University of New York, Albany, New York 12222, USA*

R. Gorodeisky, N. Guttman, D. R. Peimer, and A. Soffer  
*Tel Aviv University, School of Physics and Astronomy, Tel Aviv, 69978, Israel*

P. Lund and S. M. Spanier  
*University of Tennessee, Knoxville, Tennessee 37996, USA*

J. L. Ritchie, A. M. Ruland, R. F. Schwitters, and B. C. Wray  
*University of Texas at Austin, Austin, Texas 78712, USA*

J. M. Izen and X. C. Lou  
*University of Texas at Dallas, Richardson, Texas 75083, USA*

F. Bianchi<sup>ab</sup> and D. Gamba<sup>ab</sup>  
*INFN Sezione di Torino<sup>a</sup>; Dipartimento di Fisica Sperimentale, Università di Torino<sup>b</sup>, I-10125 Torino, Italy*

L. Lanceri<sup>ab</sup> and L. Vitale<sup>ab</sup>  
*INFN Sezione di Trieste<sup>a</sup>; Dipartimento di Fisica, Università di Trieste<sup>b</sup>, I-34127 Trieste, Italy*

F. Martinez-Vidal and A. Oyanguren  
*IFIC, Universitat de Valencia-CSIC, E-46071 Valencia, Spain*

H. Ahmed, J. Albert, Sw. Banerjee, F. U. Bernlochner, H. H. F. Choi, G. J. King,  
 R. Kowalewski, M. J. Lewczuk, I. M. Nugent, J. M. Roney, R. J. Sobie, and N. Tasneem  
*University of Victoria, Victoria, British Columbia, Canada V8W 3P6*

T. J. Gershon, P. F. Harrison, T. E. Latham, and E. M. T. Puccio  
*Department of Physics, University of Warwick, Coventry CV4 7AL, United Kingdom*

H. R. Band, S. Dasu, Y. Pan, R. Prepost, and S. L. Wu  
*University of Wisconsin, Madison, Wisconsin 53706, USA*

We present a measurement of the  $B^+ \rightarrow \omega \ell^+ \nu$  branching fraction based on a sample of 467 million  $B\bar{B}$  pairs recorded by the *BABAR* detector at the SLAC PEP-II  $e^+e^-$  collider. We observe  $1125 \pm 131$  signal decays, corresponding to a branching fraction of  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu) = (1.21 \pm 0.14 \pm 0.08) \times 10^{-4}$ , where the first error is statistical and the second is systematic. The dependence of the decay rate on  $q^2$ , the invariant mass squared of the leptons, is compared to QCD predictions of the form factors based on a quark model and light-cone sum rules.

PACS numbers: 13.20.He, 12.15.Hh, 12.38.Qk, 14.40.Nd

## I. INTRODUCTION

Most theoretical and experimental studies of exclusive  $B \rightarrow X_u \ell \nu$  decays have focused on  $B \rightarrow \pi \ell \nu$  decays, while  $B \rightarrow \rho \ell \nu$  and  $B^+ \rightarrow \omega \ell^+ \nu$  [1] decays involving the vector mesons  $\rho$  and  $\omega$  have received less attention. Here  $\ell$  is an electron or muon, and  $X$  refers to a hadronic state, with the subscript  $c$  or  $u$  signifying whether the state carries charm or is charmless. Measurements of the branching fraction of  $B \rightarrow \rho \ell \nu$  are impacted by an irreducible  $B \rightarrow X_u \ell \nu$  background, typically the dominant source of systematic uncertainty. In studies of  $B^+ \rightarrow \omega \ell^+ \nu$  that background can be suppressed to a larger degree, since the  $\omega$  width is about 15 times smaller than that of the  $\rho$ . Extractions of the CKM matrix element  $|V_{ub}|$  from  $B^+ \rightarrow \omega \ell^+ \nu$  and  $B \rightarrow \rho \ell \nu$  decay rates have greater uncertainties than those from  $B \rightarrow \pi \ell \nu$ ,

due to higher backgrounds and more complex form-factor dependencies. The persistent discrepancy between  $|V_{ub}|$  measurements based on inclusive and exclusive charmless decays is a motivation for the study of different exclusive  $B \rightarrow X_u \ell \nu$  decays [2, 3].

Measurements of  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)$  have been reported by Belle [4, 5]; a measurement by *BABAR* has been performed on a partial dataset [6]. In this analysis we use the full *BABAR* dataset to measure the total branching fraction  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)$  and partial branching fractions  $\Delta\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)/\Delta q^2$  in five  $q^2$  intervals, where  $q^2$  refers to the momentum transfer squared to the lepton system.

The differential decay rate for  $B^+ \rightarrow \omega \ell^+ \nu$  is given by [7]

$$\frac{d\Gamma(B^+ \rightarrow \omega \ell^+ \nu)}{dq^2} = |V_{ub}|^2 \frac{G_F^2 q^2 p_\omega}{96\pi^3 m_B^2 c_V^2} \times \left[ |H_0|^2 + |H_+|^2 + |H_-|^2 \right], \quad (1)$$

where  $p_\omega$  is the magnitude of the  $\omega$  momentum in the  $B$  rest frame,  $m_B$  is the  $B$  mass, and  $G_F$  is the Fermi coupling constant. The isospin factor  $c_V$  is equal to  $\sqrt{2}$  for  $B^+ \rightarrow \omega \ell^+ \nu$  [8]. As described in a related *BABAR* paper [9], the three helicity functions  $H_0$ ,  $H_+$ , and  $H_-$  can be expressed in terms of two axial vector form factors  $A_1$  and  $A_2$  and one vector form factor  $V$ , which describe

\*Now at the University of Tabuk, Tabuk 71491, Saudi Arabia

†Also with Università di Perugia, Dipartimento di Fisica, Perugia, Italy

‡Now at the University of Huddersfield, Huddersfield HD1 3DH, UK

§Now at University of South Alabama, Mobile, Alabama 36688, USA

¶Also with Università di Sassari, Sassari, Italy

\*\*Deceased

strong interaction effects,

$$H_{\pm}(q^2) = (m_B + m_{\omega}) \left[ A_1(q^2) \mp \frac{2m_B p_{\omega}}{(m_B + m_{\omega})^2} V(q^2) \right],$$

$$H_0(q^2) = \frac{m_B + m_{\omega}}{2m_{\omega}\sqrt{q^2}} \times \left[ (m_B^2 - m_{\omega}^2 - q^2) A_1(q^2) - \frac{4m_B^2 p_{\omega}^2}{(m_B + m_{\omega})^2} A_2(q^2) \right].$$

We compare the measured  $q^2$  dependence of the decay rate with form factor predictions based on light-cone sum rules (LCSR) [8] and the ISGW2 quark model [10]. We also use these form factor calculations and the measured branching fraction to extract  $|V_{ub}|$ .

## II. DETECTOR, DATA SET, AND SIMULATION

The data used in this analysis were recorded with the BABAR detector at the PEP-II  $e^+e^-$  collider operating at the  $\Upsilon(4S)$  resonance. We use a data sample of  $426 \text{ fb}^{-1}$ , corresponding to  $(467 \pm 5)$  million produced  $B\bar{B}$  pairs. In addition, we use  $44 \text{ fb}^{-1}$  of data collected 40 MeV below the  $B\bar{B}$  production threshold. This off-resonance sample is used to validate the simulation of the non- $B\bar{B}$  contributions whose principal source is  $e^+e^-$  annihilation to  $q\bar{q}$  pairs, where  $q = u, d, s, c$ .

The PEP-II collider and BABAR detector have been described in detail elsewhere [11]. Charged particles are reconstructed in a five-layer silicon tracker positioned close to the beam pipe and a forty-layer drift chamber. Particles of different masses are distinguished by their ionization energy loss in the tracking devices and by a ring-imaging Cerenkov detector. Electromagnetic showers from electrons and photons are measured in a finely segmented CsI(Tl) calorimeter. These detector components are embedded in a 1.5 T magnetic field of a superconducting solenoid; its steel flux return is segmented and instrumented with planar resistive plate chambers and limited streamer tubes to detect muons that penetrate the magnet coil and steel.

We use Monte Carlo (MC) techniques [12, 13] to simulate the production and decay of  $B\bar{B}$  and  $q\bar{q}$  pairs and the detector response [14], to estimate signal and background efficiencies and resolutions, and to extract the expected signal and background distributions. The size of the simulated sample of generic  $B\bar{B}$  events exceeds the  $B\bar{B}$  data sample by about a factor of three, while the MC samples for inclusive and exclusive  $B \rightarrow X_u \ell \nu$  decays exceed the data samples by factors of 15 or more. The MC sample for  $q\bar{q}$  events is about twice the size of the  $q\bar{q}$  contribution in the  $\Upsilon(4S)$  data.

The MC simulation of semileptonic decays uses the same models as in a recent BABAR analysis [9]. The simulation of inclusive charmless semileptonic decays  $B \rightarrow X_u \ell \nu$  is based on predictions of a heavy quark expansion [15] for the differential decay rates. For the simulation of  $B \rightarrow \pi \ell \nu$  decays we use the ansatz of [16] for

the  $q^2$  dependence, with the single parameter  $\alpha_{BK}$  set to the value determined in a previous BABAR analysis [17]. All other exclusive charmless semileptonic decays  $B \rightarrow X_u \ell \nu$ , including the signal, are generated with form factors determined by LCSR [8, 18]. For  $B \rightarrow D \ell \nu$  and  $B \rightarrow D^* \ell \nu$  decays we use parameterizations of the form factors [19, 20] based on heavy quark effective theory; for the generation of the decays  $B \rightarrow D^{**} \ell \nu$ , we use the ISGW2 model [10].

## III. CANDIDATE SELECTION

In the following, we describe the selection and kinematic reconstruction of signal candidates, the definition of the various background classes, and the application of neural networks to further suppress these backgrounds.

The primary challenge in studying charmless semileptonic  $B$  decays is to separate signal decays from Cabibbo-favored  $B \rightarrow X_c \ell \nu$  decays, which have a branching fraction approximately 50 times larger than that of  $B \rightarrow X_u \ell \nu$ . A significant background also arises due to multi-hadron continuum events.

Based on the origin of the candidate lepton we distinguish three categories of events: 1) *Signal* candidates with a charged lepton from a true  $B^+ \rightarrow \omega \ell^+ \nu$  decay; 2)  *$B\bar{B}$  background* with a charged lepton from all non-signal  $B\bar{B}$  events; 3) *Continuum background* from  $e^+e^- \rightarrow q\bar{q}$  events. The  $\omega$  meson is reconstructed in its dominant decay,  $\omega \rightarrow \pi^+ \pi^- \pi^0$ . For each of the three categories of events we distinguish correctly reconstructed  $\omega \rightarrow \pi^+ \pi^- \pi^0$  decays (true- $\omega$ ) from combinatorial- $\omega$  candidates, for which at least one of the reconstructed pions originates from a particle other than the  $\omega$ .

### A. Preselection

Signal candidates are selected from events with at least four charged tracks, since a  $B^+ \rightarrow \omega \ell^+ \nu$  decay leaves three tracks and the second  $B$  in the event is expected to produce at least one track. The magnitude of the sum of the charges of all reconstructed tracks is required to be less than two, helping to reject events with at least two undetected particles.

The preselection places requirements on the reconstructed lepton,  $\omega$  meson, and neutrino from the  $B^+ \rightarrow \omega \ell^+ \nu$  decay. At this stage in the analysis we allow for more than one candidate per event.

The lepton is identified as either an electron or muon. The electron identification efficiency is greater than 90% and constant as a function of momentum above 1 GeV, while the muon identification efficiency is between 65%–75% for momenta of 1.5–3 GeV. The pion misidentification rates are about 0.1% for the electron selector and 1% for the muon selector. The lepton is required to have a momentum in the center-of-mass (c.m.) frame greater than 1.6 GeV. This requirement significantly reduces

the background from hadrons that are misidentified as leptons, and also removes a large fraction of true leptons from secondary decays or photon conversions and from  $B \rightarrow X_c \ell \nu$  decays. The acceptance of the detector for leptons covers momentum polar angles in the range  $0.41 \leq \theta \leq 2.54$  rad.

For the reconstruction of the decay  $\omega \rightarrow \pi^+ \pi^- \pi^0$ , we require that the candidate charged pions are not identified as leptons or kaons. The reconstructed  $\omega$  mass must be in the range  $680 < m_{3\pi} < 860$  MeV, and the  $\pi^0$  candidate is required to have an invariant mass of  $115 < m_{\gamma\gamma} < 150$  MeV. To reduce combinatorial  $\omega$  background, we require minimum momenta for the three pion candidates,  $p_{\pi^\pm} > 200$  MeV and  $p_{\pi^0} > 400$  MeV, and also energies of at least 80 MeV for photons from the  $\pi^0$  candidate.

The charged lepton candidate is combined with an  $\omega$  candidate to form a so-called  $Y$  candidate. The charged tracks associated with the  $Y$  candidate are fitted to a common vertex  $Y_{\text{vtx}}$ . This vertex fit must yield a  $\chi^2$  probability  $\text{Prob}(\chi^2, Y_{\text{vtx}}) > 0.1$ . To further reduce backgrounds without significant signal losses, we impose two-dimensional restrictions on the momenta of the lepton and  $\omega$ . Each  $Y$  candidate must satisfy at least one of the following conditions on the c.m. momentum of the lepton and  $\omega$ :  $p_\omega^* > 1.3$  GeV, or  $p_\ell^* > 2.0$  GeV, or  $p_\ell^* + p_\omega^* > 2.65$  GeV, where quantities with an asterisk refer to the c.m. frame. These requirements reject background candidates that are inconsistent with the phase space of the signal decay. The condition  $|\cos \theta_{BY}| \leq 1.0$ , where  $\cos \theta_{BY} = (2E_B^* E_Y^* - M_B^2 - M_Y^2)/(2p_B^* p_Y^*)$  is the cosine of the angle between the momentum vectors of the  $B$  meson and the  $Y$  candidate, should be fulfilled for a well-reconstructed  $Y$  candidate originating from a signal decay [21]. The energy  $E_B^*$  and momentum  $p_B^*$  of the  $B$  meson are not measured event by event. Specifically,  $E_B^* = \sqrt{s}/2$ , where  $\sqrt{s}$  is the c.m. energy of the colliding beams, and the  $B$  momentum is derived as  $p_B^* = \sqrt{E_B^{*2} - m_B^2}$ . To allow for the finite resolution of the detector, we impose the requirement  $-1.2 < \cos \theta_{BY} < 1.1$ .

The neutrino four-momentum is inferred from the missing energy and momentum of the whole event:  $(E_{\text{miss}}, \vec{p}_{\text{miss}}) = (E_{e^+e^-}, \vec{p}_{e^+e^-}) - (\sum_i E_i, \sum_i \vec{p}_i)$ , where  $E_{e^+e^-}$  and  $\vec{p}_{e^+e^-}$  are the energy and momentum of the colliding beam particles, and the sums are performed over all tracks and all calorimeter clusters without an associated track. If all tracks and clusters in an event are well-measured, and there are no undetected particles besides a single neutrino, then the measured distribution of the missing mass squared,  $m_{\text{miss}}^2 = E_{\text{miss}}^2 - p_{\text{miss}}^2$ , peaks at zero. We require the reconstructed neutrino mass to be consistent with zero,  $|m_{\text{miss}}^2/(2E_{\text{miss}})| < 2.5$  GeV, and the missing momentum to exceed 0.5 GeV. The polar angle of the missing momentum vector is also required to pass through the fiducial region of the detector,  $0.3 < \theta_{\text{miss}} < 2.2$  rad.

Other restrictions are applied to suppress  $q\bar{q}$  back-

ground, which has a two-jet topology in contrast to  $B\bar{B}$  events with a more uniform angular distribution of the tracks and clusters. Events must have  $R_2 \leq 0.5$ , where  $R_2$  is the second normalized Fox-Wolfram moment [22], determined from all charged and neutral particles in the event. We also require  $\cos \Delta\theta_{\text{thrust}} \leq 0.9$ , where  $\Delta\theta_{\text{thrust}}$  is the angle between the thrust axis of the  $Y$  candidate's decay particles and the thrust axis of all other detected particles in the event. We require  $L2 < 3.0$  GeV, with  $L2 = \sum_i p_i^* \cos^2 \theta_i^*$ , where the sum runs over all tracks in the event excluding the  $Y$  candidate, and  $p_i^*$  and  $\theta_i^*$  refer to the c.m. momenta and the angles measured with respect to the thrust axis of the  $Y$  candidate.

We reject candidates that have a charged lepton and a low-momentum charged pion consistent with a  $B^0 \rightarrow D^{*-} \ell^+ \nu$ ,  $D^{*-} \rightarrow \bar{D}^0 \pi_{\text{slow}}^-$  decay as described in [23].

The kinematic consistency of the candidate decay with a signal  $B$  decay is ascertained by restrictions on two variables, the beam-energy substituted  $B$  mass  $m_{\text{ES}}$ , and the difference between the reconstructed and expected energy of the  $B$  candidate  $\Delta E$ . In the laboratory frame these variables are defined as  $m_{\text{ES}} = \sqrt{(s/2 + \vec{p}_B \cdot \vec{p}_{e^+e^-})^2 / E_{e^+e^-}^2 - p_B^2}$  and  $\Delta E = (P_{e^+e^-} \cdot P_B - s/2) / \sqrt{s}$ , where  $P_B = (E_B, \vec{p}_B)$  and  $P_{e^+e^-} = (E_{e^+e^-}, \vec{p}_{e^+e^-})$  are the four-momenta of the  $B$  meson and the colliding beams, respectively. For correctly reconstructed signal  $B$  decays, the  $\Delta E$  distribution is centered at zero, and the  $m_{\text{ES}}$  distribution peaks at the  $B$  mass. We restrict candidates to  $-0.95 < \Delta E < 0.95$  GeV and  $5.095 < m_{\text{ES}} < 5.295$  GeV.

## B. Neural Network Selection

To separate signal candidates from the remaining background we employ two separate neural networks (NN), to suppress  $q\bar{q}$  background and  $B \rightarrow X_c \ell \nu$  background. The  $q\bar{q}$  NN is trained on a sample passing the preselection criteria, while the  $B \rightarrow X_c \ell \nu$  NN is trained on a sample passing both the preselection and the  $q\bar{q}$  neural network criteria. The training is performed with signal and background MC samples. These NN are multilayer perceptrons that have two hidden layers with seven and three nodes.

The variables used as inputs to the  $q\bar{q}$  NN are  $R_2$ ,  $L2$ ,  $\cos \Delta\theta_{\text{thrust}}$ ,  $\cos \theta_{BY}$ ,  $m_{\text{miss}}^2/(2E_{\text{miss}})$ ,  $\text{Prob}(\chi^2, Y_{\text{vtx}})$ , the polar angle of the missing momentum vector in the laboratory frame, and the Dalitz plot amplitude  $A_{\text{Dalitz}} = \alpha |\vec{p}_{\pi^+} \times \vec{p}_{\pi^-}|$ , with the  $\pi^+$  and  $\pi^-$  momenta measured in the  $\omega$  rest frame and scaled by a normalization factor  $\alpha$ . True  $\omega$  mesons typically have larger values of  $A_{\text{Dalitz}}$  than combinatorial  $\omega$  candidates reconstructed from unrelated pions. The  $B \rightarrow X_c \ell \nu$  NN uses the same variables, except for  $\cos \Delta\theta_{\text{thrust}}$ , which is replaced by  $\cos \theta_{W\ell}$ , the helicity angle of the lepton, defined as the angle between the momentum of the lepton in the rest frame of the virtual  $W$  and the momentum of the  $W$  in

the rest frame of the  $B$ . The data and MC simulation agree well for the NN input variables at each stage of the selection. The NN discriminators are chosen by maximizing  $\sqrt{\epsilon_{\text{sig}}^2 + (1 - \eta_{\text{bkg}})^2}$ , where  $\epsilon_{\text{sig}}$  is the efficiency of the signal and  $\eta_{\text{bkg}}$  is the fraction of the background misidentified as signal.

The selection efficiencies for the various stages of the candidate selection for the signal and background components are given in Table I. After the preselection and NN selection, 21% of events in data contribute multiple  $B^+ \rightarrow \omega \ell^+ \nu$  candidates. The candidate with the largest value of  $\text{Prob}(\chi^2, Y_{\text{vtx}})$  is retained. For the remaining candidates, the reconstructed 3-pion mass is required to be consistent with the  $\omega$  nominal mass [24],  $|m_{3\pi} - m_\omega| < 23 \text{ MeV}$ . The overall signal efficiency is 0.73% if the reconstructed candidate includes a true  $\omega$  and 0.21% if it includes a combinatorial  $\omega$ . The efficiencies of the  $B\bar{B}$  and  $q\bar{q}$  backgrounds are suppressed by several orders of magnitude relative to the signal.

TABLE I: Successive efficiencies (in %) predicted by MC simulation for each stage of the selection, for true- and combinatorial- $\omega$  signal, and backgrounds from  $B\bar{B}$  and  $q\bar{q}$  events.

Source	true- $\omega$ signal	comb.- $\omega$ signal	$B\bar{B}$	$q\bar{q}$
Preselection	1.9	4.8	0.0094	0.00073
Neural nets	43	17	7.9	11
3-pion mass	88	26	24	30
Total (product)	0.73	0.21	0.00018	0.000024

### C. Data-MC Comparisons

The determination of the number of signal events relies heavily on the MC simulation to correctly describe the efficiencies and resolutions, as well as the distributions for signal and background sources. Therefore a significant effort has been devoted to detailed comparisons of data and MC distributions, for samples that have been selected to enhance a given source of background.

Specifically, we have studied the MC simulation of the neutrino reconstruction for a control sample of  $B^0 \rightarrow D^{*-} \ell^+ \nu$  decays, with  $D^{*-} \rightarrow \bar{D}^0 \pi_{\text{slow}}^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^- \pi^0$ . This final state is similar to that of the  $B^+ \rightarrow \omega \ell^+ \nu$  decay, except for the addition of the slow pion  $\pi_{\text{slow}}^-$  and the substitution of a  $K^+$  for a  $\pi^+$ . This control sample constitutes a high-statistics and high-purity sample on which to test the neutrino reconstruction. We compare data and MC distributions for the control sample and find good agreement for the variables used in the preselection and as inputs to the NN. We have also used this sample to study the resolution of the neutrino reconstruction and its impact on  $q^2$ ,  $m_{\text{ES}}$ , and  $\Delta E$ .

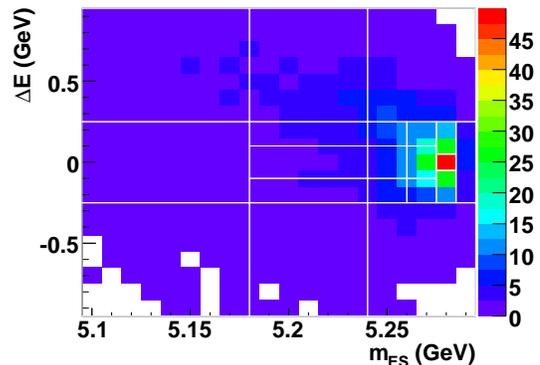


FIG. 1: Distribution of  $\Delta E$  versus  $m_{\text{ES}}$  for true- $\omega$  signal MC. The 20 bins into which the plane is divided for the fit histogram are overlaid.

## IV. SIGNAL EXTRACTION

### A. Fit Method

We determine the signal yields by performing an extended binned maximum-likelihood fit to the observed three-dimensional  $\Delta E$ - $m_{\text{ES}}$ - $q^2$  distributions. The fit technique [25] accounts for the statistical fluctuations of the data and MC samples.

For this fit the  $\Delta E$ - $m_{\text{ES}}$  plane is divided into 20 bins, as shown in Fig. 1, and the data are further subdivided into five bins in  $q^2$ , chosen to contain roughly equal numbers of signal events. The  $q^2$  resolution is dominated by the neutrino reconstruction. It can be improved by substituting the missing energy with the magnitude of the missing momentum and by rescaling  $\vec{p}_{\text{miss}}$  to force  $\Delta E = 0$ ,  $q_{\text{corr}}^2 = [(E_\ell, \vec{p}_\ell) + \delta \cdot (p_{\text{miss}}, \vec{p}_{\text{miss}})]^2$ , where  $\delta = 1 - \Delta E/E_{\text{miss}}$ . This correction to  $q^2$  is used in the fit.

We describe the measured  $\Delta E$ - $m_{\text{ES}}$ - $q^2$  distribution as a sum of four contributions:  $B^+ \rightarrow \omega \ell^+ \nu$  signal (both true- $\omega$  and combinatoric- $\omega$ ), true- $\omega$   $B\bar{B}$ , true- $\omega$   $q\bar{q}$ , and the sum of the combinatorial- $\omega$  background from  $B\bar{B}$  and  $q\bar{q}$  events.

While the  $\Delta E$ - $m_{\text{ES}}$  shapes for the signal and true- $\omega$   $B\bar{B}$  and  $q\bar{q}$  sources are taken from MC samples, we choose to represent the dominant combinatorial- $\omega$  background by the distributions of data events in the  $m_{3\pi}$  sidebands, thereby reducing the dependence on MC simulation of these backgrounds. The normalization of these background data is taken from a fit to the 3- $\pi$  mass distribution in the range  $0.680 < m_{3\pi} < 0.880 \text{ GeV}$ . To obtain a sample corresponding to the combinatorial- $\omega$  background from  $B\bar{B}$  and  $q\bar{q}$  events only, we subtract the MC simulated  $m_{3\pi}$  contribution of the small combinatorial- $\omega$   $B^+ \rightarrow \omega \ell^+ \nu$  signal sample. To the resulting  $m_{3\pi}$  distribution we fit the sum of a relativistic Breit-Wigner convolved with a normalized Gaussian function, and the combinatorial background described by a second degree

polynomial. The resulting fit to the  $m_{3\pi}$  distribution for the all- $q^2$  sample is shown in Fig. 2. The  $\chi^2$  per number of degrees of freedom (dof) for the fits are within the range expected for good fits. The fitted background function is used to determine the weights to apply to the upper and lower sidebands to scale them to the expected yield of combinatorial- $\omega$   $B\bar{B}$  and  $q\bar{q}$  background in the  $m_{3\pi}$  peak region.

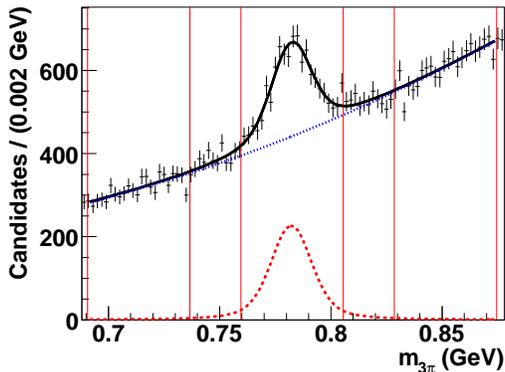


FIG. 2: Fit to the distribution of  $m_{3\pi}$  for data from the all- $q^2$  sample, with MC combinatorial- $\omega$  signal subtracted. The dashed (red) and dotted (blue) curves describe the fitted peaking and combinatorial background functions, respectively, and the solid (black) curve is their sum. The peak and sideband regions are also indicated.

The peak and two sideband regions are chosen to have a width of 46 MeV and are separated by 23 MeV, as indicated in Fig. 2. Since the normalization of the combinatorial- $\omega$  signal contribution depends on the fitted signal yield, which is a priori unknown, this component is determined iteratively.

The fit has seven free parameters, five for the signal yields in each  $q^2$  bin, and one each for the yields of the true- $\omega$   $B\bar{B}$  and  $q\bar{q}$  backgrounds, the shapes of the distributions are taken from MC simulations. The fitted yields are expressed as scale factors relative to the default yields of the MC simulation. The total signal yield is taken as the sum of the fitted yields in the individual  $q^2$  bins, taking into account correlations.

## B. Fit Results

The fitting procedure has been validated on pseudo-experiments generated from the MC distributions. We find no biases and the uncertainties follow the expected statistical distribution.

The yields of the signal, true- $\omega$   $B\bar{B}$ , and true- $\omega$   $q\bar{q}$  components obtained from the binned maximum-likelihood fit to  $\Delta E$ - $m_{\text{ES}}$ - $q^2$  are presented in Table II. Projections of the fitted distributions of  $m_{\text{ES}}$  for the all- $q^2$  fit and for the five  $q^2$  bins fit are shown in Fig. 3. The agreement

between the data and fitted MC samples is reasonable for distributions of  $\Delta E$ ,  $m_{\text{ES}}$ , and  $q^2$ , as indicated by the  $\chi^2/\text{dof}$  of the fit, 106/93, which has a probability of 16%. The fixed combinatorial- $\omega$  background yield accounts for 83% of all backgrounds. The correlations among the parameters are listed in Table III. The strongest correlation is  $-72\%$ , between the signal and  $q\bar{q}$  yields in the first  $q^2$  bin, which contains most of the  $q\bar{q}$  background. The correlation between signal and  $B\bar{B}$  background is strongest in the last  $q^2$  bin,  $-40\%$ , because of a large contribution from other  $B \rightarrow X_u \ell \nu$  decays. Correlations among signal yields are significantly smaller.

The branching fraction,  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)$ , averaged over electron and muon channels, is defined as  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu) = \sum_i (N_i^{\text{sig}} / \epsilon_i^{\text{sig}}) / (4f_{\pm} N_{B\bar{B}})$ , where  $N_i^{\text{sig}}$  refers to the number of reconstructed electron and muon signal events in  $q^2$  bin  $i$ ,  $\epsilon_i^{\text{sig}}$  is the reconstruction efficiency,  $f_{\pm}$  is the fraction of  $B^+ B^-$  decays in all  $B\bar{B}$  events, and  $N_{B\bar{B}}$  is the number of produced  $B\bar{B}$  events. The factor of 4 comes from the fact that  $\mathcal{B}$  is quoted as the average of  $\ell = e$  and  $\mu$  samples, not the sum, and the fact that either of the two  $B$  mesons in the  $B^+ B^-$  event may decay into the signal mode. The  $q^2$  resolution in the signal region is  $0.36 \text{ GeV}^2$ , smaller than the width of the  $q^2$  bins. To account for the finite  $q^2$  resolution, the background-subtracted, efficiency-corrected spectrum is adjusted by deriving from the signal MC the ratio of the true and reconstructed  $q^2$  spectra,  $(d\mathcal{B}/dq_{\text{true}}^2) / (d\mathcal{B}/dq_{\text{reco}}^2)$ . This ratio differs by 9% at low  $q^2$  and considerably less at higher  $q^2$ . The partial and total branching fractions listed in Table IV are corrected for the effects of finite  $q^2$  resolution and efficiency.

## V. SYSTEMATIC UNCERTAINTIES

Table V summarizes the contributions to the systematic uncertainty. The event reconstruction systematic uncertainties are most sensitive to the neutrino reconstruction, which depends on the detection of all of the particles in the event. To assess the impact of the uncertainty of the measured efficiencies for charged tracks, the MC signal and background samples are reprocessed and the analysis is repeated, after tracks have been eliminated at random with a probability determined by the uncertainty in the tracking efficiency. Similarly, we evaluate the impact from uncertainties in the photon reconstruction efficiency by eliminating photons at random as a function of the photon energy. Since a  $K_L^0$  leaves no track and deposits only a small fraction of its energy in the calorimeter, the reconstruction of the neutrino is impacted. The uncertainty on the  $K_L^0$  MC simulation involves the shower energy deposited by the  $K_L^0$  in the calorimeter, the  $K_L^0$  detection efficiency, and the inclusive  $K_L^0$  production rate as a function of momentum from  $B\bar{B}$  events.

The impact of the changes to the simulated background distributions which enter the fit are smaller than for

TABLE II: Number of events and their statistical uncertainties, as determined from the fit, compared with the number of observed events in data. The combinatorial- $\omega$  background yields are fixed in the fit; the quoted uncertainties are derived from the sideband subtraction.

$q^2$ range (GeV <sup>2</sup> )	0–4	4–8	8–10	10–12	12–21	0–21
All signal	257 ± 72	238 ± 44	161 ± 32	177 ± 32	293 ± 57	1125 ± 131
True- $\omega$ signal	238	209	136	137	168	869
Comb.- $\omega$ signal	19	28	25	40	125	256
$B\bar{B}$ (true- $\omega$ )	105 ± 19	192 ± 34	154 ± 27	195 ± 34	411 ± 73	1057 ± 187
$q\bar{q}$ (true- $\omega$ )	409 ± 96	145 ± 34	65 ± 15	34 ± 8	64 ± 15	716 ± 167
Comb.- $\omega$ bkgd.	1741 ± 23	1818 ± 24	1240 ± 20	1520 ± 22	3913 ± 35	10232 ± 57
Data	2504 ± 50	2433 ± 49	1605 ± 40	1858 ± 43	4738 ± 69	13138 ± 115

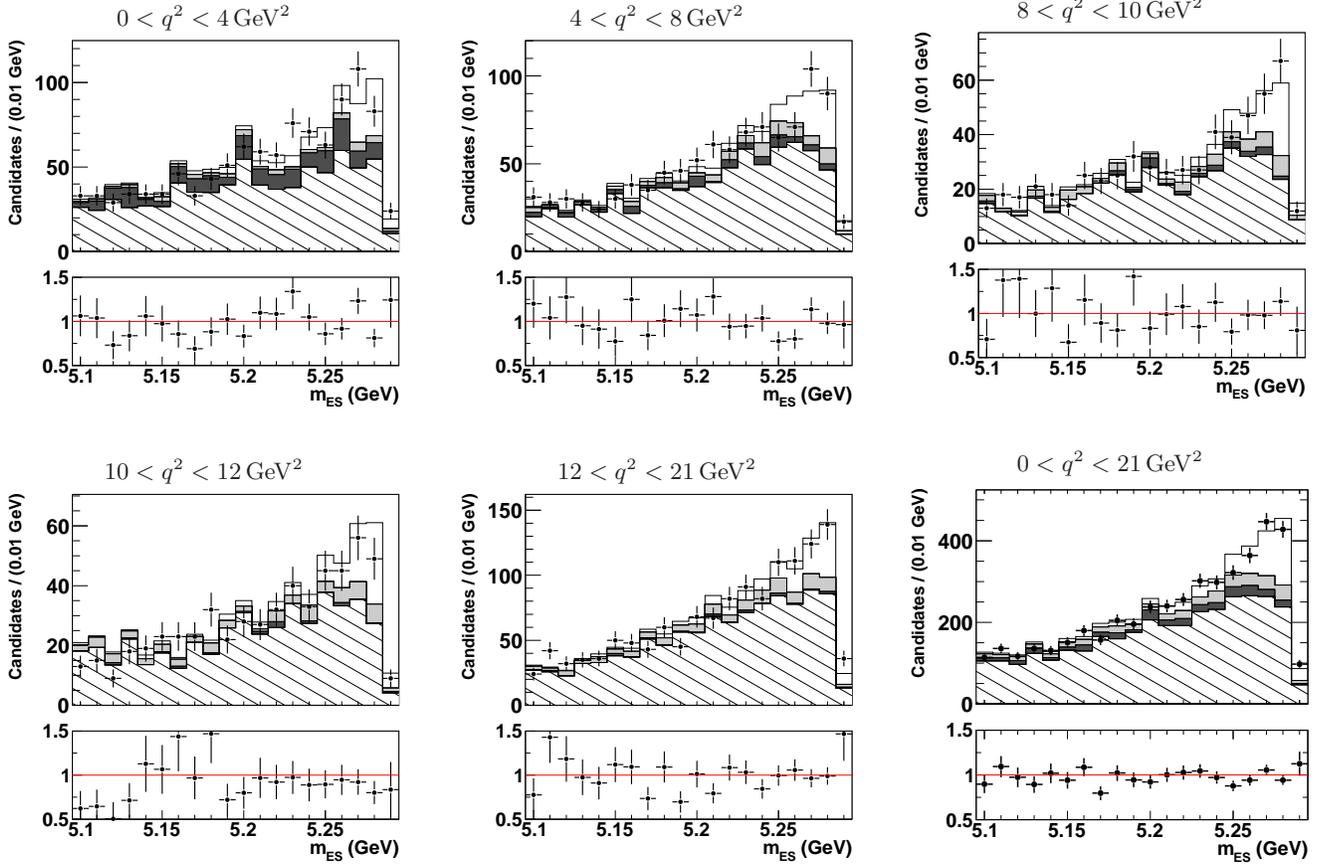


FIG. 3: Distributions of  $m_{ES}$  after the fit and the ratio of the data to the fitted predictions, for five separate  $q^2$  bins and the full  $q^2$  range, in the  $\Delta E$  signal band,  $-0.25 < \Delta E \leq 0.25$  GeV. The points represent data with statistical uncertainties, while the stacked histograms represent the sum of fitted source components, signal (white), true- $\omega$   $B\bar{B}$  (light gray), true- $\omega$   $q\bar{q}$  (dark gray), and combinatorial- $\omega$  background (diagonally hatched).

the signal, since the large combinatorial backgrounds are taken from data, rather than MC simulations. As an estimate of the impact of these variations of the MC simulated distributions on the  $q^2$  dependent signal yield, we combine the observed reduction in the signal distribution with the impact of the changes to  $q\bar{q}$  and  $B\bar{B}$  backgrounds on the signal yield, taking into account the correlations obtained from the fit (see Table III). Since the correlations between signal and backgrounds are small at

high  $q^2$ , the impact of the uncertainties in the background are also modest. This procedure avoids large statistical fluctuations of the fit procedure that have been observed to be larger than the changes in the detection efficiencies. However, this procedure does not account for the small changes in the shape of the distributions, and we therefore sum the magnitude of the changes for signal and background, rather than adding them in quadrature or taking into account the signs of the correlations of the

TABLE III: Correlations among the fit scale factors  $p_k^s$  for the simulated source  $s$  and  $q^2$  bin  $k$ . The scale factors for  $q\bar{q}$  and  $B\bar{B}$  apply to the full  $q^2$  range.

	$p^{q\bar{q}}$	$p^{B\bar{B}}$	$p_1^{\omega\ell\nu}$	$p_2^{\omega\ell\nu}$	$p_3^{\omega\ell\nu}$	$p_4^{\omega\ell\nu}$	$p_5^{\omega\ell\nu}$
$p^{q\bar{q}}$	1.000	-0.466	-0.724	-0.106	-0.031	0.051	0.088
$p^{B\bar{B}}$		1.000	0.223	-0.249	-0.253	-0.284	-0.401
$p_1^{\omega\ell\nu}$			1.000	0.121	0.061	0.001	-0.011
$p_2^{\omega\ell\nu}$				1.000	0.105	0.094	0.128
$p_3^{\omega\ell\nu}$					1.000	0.088	0.121
$p_4^{\omega\ell\nu}$						1.000	0.125
$p_5^{\omega\ell\nu}$							1.000

TABLE IV: Measured  $B^+ \rightarrow \omega\ell^+\nu$  branching fraction and partial branching fractions in bins of  $q^2$  with statistical and systematic uncertainties.

$q^2$ (GeV <sup>2</sup> )	$\Delta\mathcal{B}$ ( $\times 10^{-4}$ )
0-4	$0.214 \pm 0.060 \pm 0.024$
4-8	$0.200 \pm 0.037 \pm 0.010$
8-10	$0.147 \pm 0.029 \pm 0.010$
10-12	$0.169 \pm 0.031 \pm 0.010$
12-21	$0.482 \pm 0.093 \pm 0.038$
0-12	$0.730 \pm 0.083 \pm 0.054$
0-21	$1.212 \pm 0.140 \pm 0.084$

signal and backgrounds in a given  $q^2$  bin.

We assign an uncertainty on the identification efficiency of electrons and muons, as well as on the lepton and kaon vetoes of the  $\omega$  daughter pions, based on the change in signal yield after varying the selector efficiencies within their uncertainties.

The uncertainty in the calculation of the LCSR form factors impacts the uncertainty on the branching fraction because it affects the predicted  $q^2$  distribution of the signal and thereby the fitted signal yield. We assess the impact by varying the form factors within their uncertainties. We include the uncertainty on the branching fraction of the  $\omega$  decay,  $\mathcal{B}(\omega \rightarrow \pi^+\pi^-\pi^0) = (89.2 \pm 0.7) \times 10^{-2}$  [24]. To evaluate the uncertainty from radiative corrections, candidates are reweighted by 20% of the difference between the spectra with and without PHOTOS [26], which models the final state radiation of the decay.

The uncertainty on the true- $\omega$  backgrounds has a small impact on the signal yield since these components represent a small fraction of the total sample. To assess the uncertainty of the  $\Delta E$ - $m_{ES}$ - $q^2$  shapes of the true- $\omega$   $q\bar{q}$  and true- $\omega$   $B\bar{B}$  samples, the fit is repeated after the events are reweighted to reproduce the inclusive  $\omega$  momentum distribution measured in  $B\bar{B}$  and  $q\bar{q}$  events. We also assess the uncertainty on the modeling of the semileptonic backgrounds by varying the branching fractions and form factors of the exclusive and inclusive  $B \rightarrow X_u\ell\nu$  [24] and  $B \rightarrow X_c\ell\nu$  backgrounds [3] within their uncertainties.

To assess the uncertainties that result from the MC prediction of the  $m_{3\pi}$  distribution of the combinatorial- $\omega$  signal, we use the uncorrected distribution, in which

the combinatorial- $\omega$  signal is not subtracted from the  $m_{3\pi}$  sidebands, and the signal fit parameter is set to scale only the true- $\omega$  signal contribution. Twenty percent of the difference between the nominal and uncorrected results is taken as the systematic uncertainty; it is largest for  $12 < q^2 < 21$  GeV<sup>2</sup> because the fraction of combinatorial- $\omega$  signal in this  $q^2$  bin is large. The sideband event yields determined from the  $m_{3\pi}$  fit are varied within their fit errors to determine the statistical uncertainty on the combinatorial- $\omega$  background. The uncertainty in the chosen  $m_{3\pi}$  ansatz is assessed by repeating the  $m_{3\pi}$  fits, replacing the nominal functions for the peak and background components. For the background component, we use a third instead of a second degree polynomial. For the peaking component, we use a Gaussian function in place of a relativistic Breit-Wigner convoluted with a Gaussian function. The systematic error from the  $m_{3\pi}$  ansatz is taken as the sum in quadrature of the change in signal yield for each of these functional variations.

The branching fraction depends inversely on the value of  $N_{B\bar{B}}$ , which is determined with a precision of 1.1% [27]. At the  $\Upsilon(4S)$  resonance, the fraction of  $B^+B^-$  events is measured to be  $f_{\pm} = 0.516 \pm 0.006$  [24], with an uncertainty of 1.2%.

TABLE V: Systematic uncertainties in % on the branching fraction.

$q^2$ range (GeV <sup>2</sup> )	0-4	4-8	8-10	10-12	12-21	0-21
<b>Event reconstruction</b>						
Tracking efficiency	3.9	1.5	2.8	2.3	1.1	2.0
Photon efficiency	2.0	1.7	3.3	1.1	0.6	1.5
$K_L$ prod./interactions	4.8	1.8	2.5	1.1	1.4	1.9
Lepton identification	1.6	1.5	1.5	1.2	1.2	1.3
$K/\ell$ veto of $\omega$ daughters	1.7	1.7	1.7	1.7	1.8	1.7
<b>Signal simulation</b>						
Signal form factors	6.3	1.5	1.1	2.9	4.6	4.8
$\mathcal{B}(\omega \rightarrow \pi^+\pi^-\pi^0)$	0.8	0.8	0.8	0.8	0.8	0.8
Radiative corrections	0.4	0.3	0.2	0.1	0.2	0.2
<b>True-<math>\omega</math> background</b>						
$q\bar{q}$ $\Delta E$ - $m_{ES}$ - $q^2$ shapes	2.6	0.1	0.4	0.2	0.3	0.5
$B\bar{B}$ $\Delta E$ - $m_{ES}$ - $q^2$ shapes	2.0	0.9	1.8	0.2	0.1	0.8
$B \rightarrow X_c\ell\nu$ $\mathcal{B}$ and FF	0.2	0.6	0.3	0.2	0.2	0.2
$B \rightarrow X_u\ell\nu$ $\mathcal{B}$ and FF	0.3	0.4	0.4	0.3	0.5	0.4
<b>Comb.-<math>\omega</math> sources</b>						
Signal $m_{3\pi}$ distribution	0.6	0.5	0.4	1.1	3.7	1.5
Bkgd. yield, stat. error	4.2	1.0	0.9	0.9	2.0	1.7
Bkgd. yield, ansatz error	1.7	2.2	2.7	2.7	3.5	0.9
<b><math>B</math> production</b>						
$B\bar{B}$ counting	1.1	1.1	1.1	1.1	1.1	1.1
$f_{\pm}$	1.2	1.2	1.2	1.2	1.2	1.2
<b>Syst. uncertainty</b>	11.1	5.2	6.8	5.8	7.9	6.9
<b>Stat. uncertainty</b>	28.1	18.7	20.0	18.1	19.4	11.6
<b>Total uncertainty</b>	30.2	19.4	21.1	19.0	20.9	13.5

## VI. RESULTS AND CONCLUSIONS

We have measured the branching fraction,

$$\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu) = (1.21 \pm 0.14 \pm 0.08) \times 10^{-4}, \quad (2)$$

where the first error is statistical and the second is systematic, based on  $1125 \pm 131$  observed signal candidates. Here,  $\ell$  indicates the electron or muon decay mode and not the sum over them. The measured partial branching fractions are presented in Table IV and are compared to the predictions from two form factor calculations in Fig. 4. These QCD predictions have been normalized to the measured branching fraction.

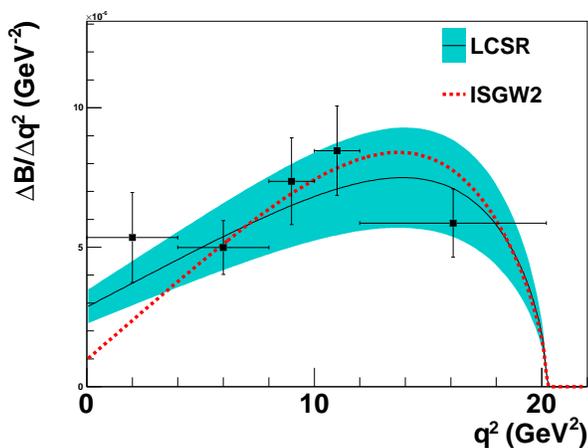


FIG. 4: Partial branching fractions (points with error bars) with respect to  $q^2$ . The data are compared with the predictions from light-cone sum rules (LCSR) [8] and a quark-model calculation (ISGW2) [10]. The uncertainty band (shaded) is given for the LCSR calculation.

Neglecting the theoretical uncertainties, the  $\chi^2/\text{NDF}$  of the measured distribution relative to the LCSR prediction [8] is 2.4/4, corresponding to a  $\chi^2$  probability of 67%; relative to the ISGW2 prediction [10] the  $\chi^2/\text{NDF}$  is 4.2/4, with a  $\chi^2$  probability of 40%. Within the large experimental uncertainties, both the LCSR and ISGW2 form factor calculations are consistent with the data. The uncertainties of the ISGW2 form factor calculation are not available. The uncertainties of the LCSR calculation were estimated by the authors to vary linearly as a function of  $q^2$ ; i.e.,  $\sigma_{\Delta\mathcal{B}/\Delta q^2}/(\text{d}\mathcal{B}/\text{d}q^2) = 21\% + 3\% \times q^2/(14 \text{ GeV}^2)$ , for the  $B \rightarrow \rho \ell \nu$  decays [28]. It is assumed that this estimate is also valid for  $B^+ \rightarrow \omega \ell^+ \nu$  decays.

The value of  $|V_{ub}|$  can be determined from the measured partial branching fraction, the  $B^+$  lifetime  $\tau_+ = (1.638 \pm 0.011) \text{ ps}$  [24], and the integral  $\Delta\zeta$  of the predicted differential decay rate:

$$|V_{ub}| = \sqrt{\frac{\Delta\mathcal{B}(q_{\min}^2, q_{\max}^2)}{\tau_+ \Delta\zeta(q_{\min}^2, q_{\max}^2)}}, \quad (3)$$

$$\Delta\zeta(q_{\min}^2, q_{\max}^2) = \frac{1}{|V_{ub}|^2} \int_{q_{\min}^2}^{q_{\max}^2} \frac{\text{d}\Gamma_{\text{theory}}}{\text{d}q^2} \text{d}q^2.$$

Table VI lists the values of  $\Delta\zeta$  and  $|V_{ub}|$  for LCSR and ISGW2 in different ranges of  $q^2$ . LCSR calculations are more accurate at low  $q^2$ , while ISGW2 predictions are more reliable at high  $q^2$ . Both form factor calculations arrive at very similar values for  $|V_{ub}|$ . These values of  $|V_{ub}|$  are consistent with the more precisely measured values from  $B \rightarrow \pi \ell \nu$  decays [29].

TABLE VI:  $|V_{ub}|$ , determined from two form factor calculations of  $\Delta\zeta$ , in different ranges of  $q^2$ . The first uncertainty is experimental (the sum in quadrature of statistical and systematic); the second uncertainty is from theory, and is only available for LCSR.

	$q^2$ (GeV <sup>2</sup> )	$\Delta\zeta$ (ps <sup>-1</sup> )	$ V_{ub} $ ( $\times 10^{-3}$ )
LCSR [8]	0–12	$3.9 \pm 0.9$	$3.37 \pm 0.23 \pm 0.38$
	12–21	$3.2 \pm 0.8$	$3.04 \pm 0.32 \pm 0.37$
	0–21	$7.1 \pm 1.7$	$3.23 \pm 0.22 \pm 0.38$
ISGW2 [10]	0–12	3.6	$3.51 \pm 0.24$
	12–21	3.4	$2.94 \pm 0.31$
	0–21	7.0	$3.24 \pm 0.22$

The value of  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)$  measured in this analysis supersedes the previous *BABAR* measurement [6] based on a smaller data sample, and is in excellent agreement with a recent result [30] based on the full *BABAR* data set. The principal difference between this analysis and the previous ones is that the combinatorial- $\omega$  background is taken from the sideband of the data  $m_{3\pi}$  distribution rather than from MC simulation. Although the dominant systematic uncertainties from event reconstruction cannot be avoided, this procedure substantially reduces the reliance on the MC simulation of this largest source of background.

Currently, the QCD predictions of the form factors, and in particular their uncertainties, have limited precision for  $B^+ \rightarrow \omega \ell^+ \nu$  and  $B \rightarrow \rho \ell \nu$  decays. These form factor uncertainties impact  $|V_{ub}|$  derived from  $\mathcal{B}(B^+ \rightarrow \omega \ell^+ \nu)$ . In the future, form factor calculations with reduced uncertainties combined with improved branching fraction measurements would enable tests and discrimination among different predictions as a function of  $q^2$ , and thereby improve the determination of  $|V_{ub}|$ .

## VII. ACKNOWLEDGMENTS

We are grateful for the extraordinary contributions of our PEP-II colleagues in achieving the excellent luminosity and machine conditions that have made this work possible. The success of this project also relies critically on

the expertise and dedication of the computing organizations that support *BABAR*. The collaborating institutions wish to thank SLAC for its support and the kind hospitality extended to them. This work is supported by the US Department of Energy and National Science Foundation, the Natural Sciences and Engineering Research Council (Canada), the Commissariat à l’Energie Atomique and Institut National de Physique Nucléaire et de Physique des Particules (France), the Bundesministerium für Bildung und Forschung and Deutsche Forschungsgemein-

schaft (Germany), the Istituto Nazionale di Fisica Nucleare (Italy), the Foundation for Fundamental Research on Matter (The Netherlands), the Research Council of Norway, the Ministry of Education and Science of the Russian Federation, Ministerio de Ciencia e Innovación (Spain), and the Science and Technology Facilities Council (United Kingdom). Individuals have received support from the Marie-Curie IEF program (European Union), the A. P. Sloan Foundation (USA) and the Binational Science Foundation (USA-Israel).

- 
- [1] The charge conjugate decay mode is implicitly included.
  - [2] J. Beringer *et al.*, Phys. Rev. **86**, 010001 (2012), page 1111.
  - [3] D. Asner *et al.* (Heavy Flavor Averaging Group) (2010), e-Print: arXiv:1010.1589 [hep-ex].
  - [4] C. Schwanda *et al.* (Belle Collaboration), Phys. Rev. Lett. **93**, 131803 (2004).
  - [5] I. Adachi *et al.* (Belle Collaboration) (2008), 0812.1414.
  - [6] B. Aubert *et al.* (*BABAR* Collaboration), Phys. Rev. **D79**, 052011 (2009).
  - [7] J. D. Richman and P. R. Burchat, Rev. Mod. Phys. **67**, 893 (1995).
  - [8] P. Ball and R. Zwicky, Phys. Rev. **D71**, 014029 (2005).
  - [9] P. del Amo Sanchez *et al.* (*BABAR* Collaboration), Phys. Rev. **D83**, 032007 (2011).
  - [10] D. Scora and N. Isgur, Phys. Rev. D **52**, 2783 (1995).
  - [11] B. Aubert *et al.*, Nucl. Instrum. Methods A **479**, 1 (2002).
  - [12] D. J. Lange, Nucl. Instrum. Methods **A462**, 152 (2001).
  - [13] T. Sjöstrand, Comp. Phys. Commun. **82**, 74 (1994).
  - [14] S. Agostinelli *et al.*, Nucl. Instrum. Methods A **506**, 250 (2003).
  - [15] F. D. Fazio and M. Neubert, J. of High Energy Phys. **6**, 17 (1999).
  - [16] D. Becirevic and A. B. Kaidalov, Phys. Lett. B **478** (2000).
  - [17] B. Aubert *et al.* (*BABAR* Collaboration), Phys. Rev. Lett. **98**, 091801 (2007).
  - [18] P. Ball and R. Zwicky, Phys. Rev. D **71**, 014015 (2005).
  - [19] N. Isgur and M. B. Wise, Phys. Lett. B **232**, 113 (1989).
  - [20] L. L. I. Caprini and M. Neubert, Nucl. Phys. **B530** (1998).
  - [21] B. H. Behrens *et al.* (CLEO Collaboration), Phys. Rev. D **61**, 052001 (2000).
  - [22] G. C. Fox and S. Wolfram, Nucl. Phys. **B149**, 413 (1979).
  - [23] B. Aubert *et al.*, Phys. Rev. Lett. **89**, 011802 (2002).
  - [24] K. Nakamura *et al.* (Particle Data Group), J. Phys. **G37**, 075021 (2010).
  - [25] R. J. Barlow and C. Beeston, Comput. Phys. Commun. **77**, 219 (1993).
  - [26] E. Barberio and Z. Was, Comput. Phys. Commun. **79**, 291 (1994).
  - [27] G. D. McGregor, *B Counting at BABAR*, MSc Thesis, University of British Columbia, Canada (2008), SLAC-R-912, e-Print: arXiv:0812.1954 [hep-ex].
  - [28] R. Zwicky (2005), private communication.
  - [29] V. G. Lüth, Ann. Rev. Nucl. Part. Sci. **61**, 119 (2011).
  - [30] J. P. Lees *et al.* (*BABAR* Collaboration), Phys. Rev. **D86**, 092004 (2012).