

# 1 **Environmental magnetic fingerprinting of anthropogenic and** 2 **natural atmospheric deposition over southwestern Europe**

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## 26 27 **Abstract**

28 Here we present an environmental magnetic study of atmospheric deposition collected by  
29 a multi-site network in Spain that covers fourteen locations representative of urban,  
30 industrial, agricultural, and natural environments across southwestern Europe. We have  
31 combined magnetic methods with scanning electron microscopy and geochemical data to  
32 characterize the magnetic mineral assemblages of particular matter across different types  
33 of sites and phenomenological scenarios, and to unravel their environmental significance  
34 in terms of the most important anthropogenic and natural components of atmospheric

35 deposition. Our results enable identification of two magnetite/maghemite components of  
36 anthropogenic origin, derived mostly from vehicular traffic, plus a hematite component  
37 associated with a baseline supply of north African dust, in all the studied sites regardless  
38 of their type. The ubiquitous presence of anthropogenic magnetite/maghemite particles  
39 in pristine natural environments, albeit in lower concentrations, point to their arrival from  
40 neighbouring urban areas through atmospheric mixing processes. Samples including  
41 particulate matter deposited during distinctively intense periods of north African dust  
42 supply are characterized by a fourth component, represented also by coarser-grained  
43 hematite, that is likely derived from a different source area within the Sahara Desert. The  
44 concomitant increase observed in these cases in the amounts of magnetite/maghemite  
45 particles suggests strongly that part of the magnetite/maghemite load attributed to  
46 anthropogenic sources for the rest of the phenomenological scenarios is aeolian in origin.  
47 This seems to explain the overall moderate correlation observed between  
48 magnetite/maghemite contents and proxies for vehicular traffic, and demonstrates the  
49 need for caution when interpreting environmental magnetic proxies for  
50 magnetite/maghemite abundances in terms of anthropogenic loads. This is especially the  
51 case in southern European cities where a steady supply of north African dust occurs  
52 throughout most of the year. Our results show a good correlation between hematite  
53 abundances and geochemical proxies for north African dust, which collectively delineate  
54 broad maxima during the summer and large peaks during distinctive dust breakouts. Thus,  
55 environmental magnetic proxies of hematite abundances can be used to monitor the  
56 contribution of natural sources to atmospheric deposition.

57

58 **Key words:** Atmospheric deposition, particulate matter, environmental magnetism,  
59 pollution, north African dust, Spain.

60

## 61 **1. Introduction**

62 Rock magnetic methods provide a rapid, cost-effective and non-destructive means of  
63 identifying magnetic mineral particles and quantifying their concentration in a wide  
64 variety of materials, including atmospheric particulate matter (PM) (Verosub and  
65 Roberts, 1995; Petrovský and Ellwood, 1999; Evans and Heller, 2003; Liu et al., 2012).  
66 Most PM-producing anthropogenic activities, such as burning of fossil fuels (car,  
67 shipping and power-plant emissions), industrial processes involving high-temperature  
68 treatments (e.g., steel, cement and ceramic production), or abrasion of metallic (e.g.,  
69 brakes) and iron-bearing (e.g., tyres) vehicular parts, are invariably associated with the

70 production of magnetic particles (Petrovský and Ellwood, 1999; Evans and Heller, 2003;  
71 Liu et al., 2012; Hofman et al., 2017). The low-coercivity minerals magnetite ( $\text{Fe}_3\text{O}_4$ ) and  
72 maghemite ( $\gamma\text{-Fe}_2\text{O}_3$ ) are, by far, the most commonly reported magnetic minerals in  
73 anthropogenic PM (Muxworthy et al., 2002, 2003; Spassov et al., 2004; Sagnotti et al.,  
74 2006, 2009; Saragnese et al., 2011; Sagnotti and Winkler, 2012; Revuelta et al., 2014,  
75 Maher et al., 2016; Hofman et al., 2017). These minerals have similar magnetic properties  
76 and are typically reported as magnetite/maghemite or magnetite-like assemblages unless  
77 more diagnostic magnetic measurements are made (Muxworthy et al., 2002). The high-  
78 coercivity mineral hematite ( $\alpha\text{-Fe}_2\text{O}_3$ ) is also formed in industrial (e.g., steel  
79 manufacturing) and power-plant fly-ashes (Petrovský and Ellwood, 1999), but has been  
80 reported less frequently in anthropogenic PM. The weak spontaneous magnetization of  
81 hematite means that in mixed magnetic mineral assemblages its diagnostic signatures are  
82 often overwhelmed by the signals of more strongly magnetic magnetite-maghemite  
83 components (Saragnese et al., 2011, Hofman et al., 2017; Tan et al., 2018). Combined  
84 magnetic and scanning/transmission electron microscopy (SEM, TEM) observations  
85 indicate that anthropogenic magnetite and maghemite are typically found as isolated  
86 particles, spherules and irregular aggregates. These individual particles range in size from  
87 <20 nm in the so-called superparamagnetic (SP) grains, from 20 to 100 nm in stable single  
88 domain (SD), from 100 to 300 nm in pseudo-single domain (PSD) or vortex, to >300 nm  
89 in multi-domain (MD) particles (Sagnotti et al., 2009, Saragnese et al., 2011; Hofman et  
90 al., 2017; Tan et al., 2018). Anthropogenic hematite is often reported as SP to MD platy  
91 specularite or lamellar crystal (Petrovský and Ellwood, 1999), which range in size  
92 between a few nanometres and some tens of microns. Because of their grain size,  
93 magnetic particles are inherently linked to respirable anthropogenic PM (e.g., defined as  
94 PM with an aerodynamic diameter of  $\leq 10 \mu\text{m}$ ,  $\text{PM}_{10}$ ), thereby posing one of the biggest  
95 environmental concerns in urban areas (WHO, 2006; Maher et al., 2016). The ability of  
96 rock magnetic measurements to rapidly identify and quantify harmful atmospheric  
97 particles via their association with magnetic particles has led to widespread use of  
98 environmental magnetic techniques in PM studies over the last two decades (see Evans  
99 and Heller, 2003; Liu et al., 2012; Hofman et al., 2017).

100 Most studies have concentrated on samples obtained in air-quality stations, where  
101  $\text{PM}_{10}$  (and sometimes also  $\text{PM}_{2.5}$ ) is collected actively by pumping air through appropriate  
102 filters (e.g., Muxworthy et al., 2001, 2003; Spassov et al., 2004; Sagnotti et al., 2006;  
103 Saragnese et al., 2011; Revuelta et al., 2014), or on biomaterials such as tree leaves, bark  
104 or mosses, where PM accumulates by passive fallout with no grain-size bias (Matzka and

105 Maher, 1999; Sagnotti et al., 2009; Sagnotti and Winkler, 2012; Hofman et al., 2017).  
106 Fewer studies have been based on direct collection of street dust (Tan et al., 2018), on  
107 PM accumulated in passive collectors designed to avoid grain-size biases (Muxworthy et  
108 al., 2002; Cao et al., 2015), or on PM accumulated in a combination of media (Abbasi et  
109 al., 2020). Although these studies often include material collected from rural sites  
110 regarded as the regional natural background (Spasov et al., 2003; Saragnese et al., 2011;  
111 Sagnotti et al., 2006), they typically focus on urban and industrial areas so that  
112 anthropogenic PM is sampled. MD and SP magnetite/maghemite grains are the most  
113 abundant magnetic particles reported in anthropogenic PM (Sagnotti et al., 2006, 2009;  
114 Lehndorff et al., 2006; Saragnese et al., 2011; Hofman et al., 2017), and a number of  
115 studies suggest that coarse MD magnetite grains with an outer layer of SP maghemite  
116 particles are representative of anthropogenic PM across a large number of urban and  
117 industrial areas (Muxworthy et al., 2003; Sagnotti et al., 2009; Saragnese et al., 2011;  
118 Revuelta et al., 2014). Such particles appear to derive from a magnetite precursor, formed  
119 by combustion or frictional heating processes, whose outer core has been oxidised to  
120 maghemite during subsequent cooling and atmospheric ageing (Muxworthy et al., 2003;  
121 Saragnese et al., 2011; Revuelta et al., 2014). A high-coercivity mineral, likely hematite,  
122 has been reported in urban and industrial PM (Saragnese et al., 2011; Revuelta et al.,  
123 2014), yet its origin and environmental significance has been seldom addressed due to  
124 difficulties in discerning its origin. Sampling duration also varies between different  
125 studies of anthropogenic PM, ranging from 24 hours for some studies of PM collected in  
126 air-quality stations (Spasov et al., 2004; Sagnotti et al., 2006) to 1 year in some studies  
127 dealing with PM accumulated in self-designed collectors (Cao et al., 2015), with sampling  
128 intervals ranging between 6 days and a full growing season when biological material is  
129 used (Hofman et al., 2017). Monitoring periods also differ, with lengths varying between  
130 some months and one year (Saragnese et al., 2011; Sagnotti et al., 2006). One of the prime  
131 goals of studies dealing with urban and industrial PM has been the development of  
132 magnetic parameters that can be employed, supported by additional geochemical data, as  
133 proxies for anthropogenic pollution load (see Hofman et al., 2017).

134 In comparison to anthropogenic PM, the magnetic characterization of natural PM has  
135 received very little attention. Locally-derived dust is expected to be relevant only in  
136 drylands (Abbasi et al., 2020) or agricultural regions or where the soil is, at least,  
137 temporally exposed. In contrast, desert dust sourced from north Africa (NAD) is known  
138 to be a substantial component of PM across many European regions, especially around  
139 the Mediterranean Sea (Pey et al., 2013, Varga et al., 2014). NAD contains large amounts

140 of magnetite, maghemite, hematite and goethite spanning a wide grain-size range  
141 (Larrasoña et al., 2015). It is, therefore, surprising that magnetite/maghemite grains  
142 derived from NAD have been reported only in a limited number of urban areas (Spasov  
143 et al., 2004; Sagnotti et al., 2006; Revuelta et al., 2014), possibly because it is obscured  
144 commonly by anthropogenic magnetite. Equally surprising is the lack of studies in which  
145 hematite has been reported as a constituent of natural PM (Sagnotti et al., 2006). Given  
146 the abundance of magnetic minerals in NAD, it is very likely that at least part of the  
147 magnetite/maghemite and hematite reported in urban and industrial PM actually has an  
148 aeolian origin. Yet, compelling evidence for this is still to be provided, and the  
149 implications for developing magnetic proxies for pollution remain unexplored. Although  
150 the knowledge gained from the magnetic characterization of PM has increased  
151 tremendously over the last decades, the large variety of methods, sampling times, and  
152 temporal and spatial resolutions employed have made it challenging to develop a  
153 comprehensive understanding of underlying anthropogenic and natural processes  
154 responsible for accumulation of PM at large temporal and geographical scales.

155 The DONAIRE network was designed to carry out an integrated study of atmospheric  
156 deposition across a large area of Spain that includes the Cantabrian-Pyrenean area, the  
157 Ebro basin, and the Mediterranean region (Pey et al., 2020). This area is characterized by  
158 variable climatic conditions that, combined with a complex topography (heights ranging  
159 from sea-level to over 3000 m a.s.l.) provide a wide variety of environmental conditions.  
160 Land-use is also highly variable, and includes a mixture of densely populated urban areas,  
161 industrial hot-spots, quasi-demographic deserts, agricultural regions, and pristine natural  
162 areas. Here we present an environmental magnetic study of atmospheric deposition  
163 collected in the framework of the DONAIRE network, which includes 14 representing  
164 urban, industrial, agricultural and natural areas, between June 2016 and June 2017. We  
165 have performed detailed magnetic characterization of the PM, assisted by SEM  
166 observations, across a set of samples selected to be representative for the different areas  
167 and phenomenological scenarios of atmospheric deposition. This information has been  
168 combined with bulk environmental magnetic properties and selected geochemical data  
169 for the entire set of samples. By integrating all results, the significance of bulk  
170 environmental magnetic properties in terms of the most important anthropogenic and  
171 natural components of atmospheric deposition, and their temporal variations, has been  
172 determined for an area representative of different anthropized and natural settings in SW  
173 Europe. The forthcoming completion of the full geochemical dataset will enable a  
174 detailed source apportionment of the atmospheric deposition collected in the framework

175 of the DONAIRE network, constraining the reliability of the magnetic proxies considered  
176 in this contribution.

177

## 178 **2. Materials and methods**

### 179 2.1. Sites, samples, and atmospheric deposition scenarios

180 Samples were collected using home-produced collectors consisting of a plastic funnel  
181 (with a diameter ranging from 26 to 34 cm) that acts as the collecting surface, a 10L  
182 plastic container, and a silicone tube that connects the funnel with the container. The  
183 silicone tube is mounted inside a rigid PVC tube that, along with the plastic container, is  
184 placed within a dark box to keep the samples in opaque conditions. Eight of the collectors  
185 were placed in the roof of air quality monitoring sites or other buildings (Fig. 1). The  
186 remaining collectors were placed at least 2 m above the ground in order to minimize the  
187 impact of local resuspension. A total of 14 contrasting micro-environments form our  
188 monitoring network, which includes four urban areas of variable sizes with their  
189 accompanying peripheral industrial activity (Pamplona, Zaragoza, Barcelona, and  
190 Palma), a heavily-polluted industrial town (Monzón), two agricultural territories (Ejea de  
191 los Caballeros and Almudévar), and seven pristine natural environments that serve as  
192 regional backgrounds (Enol, Orgi, Ordesa, Chiprana, Montseny, Joan March and  
193 Frigiliana) (Fig. 1, Table 1). Samples were collected every second week throughout the  
194 whole survey, which started at the beginning of June 2016 (when the collectors were  
195 deployed) and lasted till the end of June 2017 (when maintenance and sampling of the  
196 collectors could not be extended for a number of sites). Exceptions to this are the sites of  
197 Enol (ENO, background site), Ejea de los Caballeros (EDC, agricultural site) and  
198 Frigiliana (FRI, background site); ENO was sampled every two months due to logistic  
199 constraints, EDC was removed in March 2017 after being vandalized twice, and FRI was  
200 only incorporated to the network in October 2016. Although a 1-year monitoring period  
201 might not be representative of the climatology of the studied areas, it enables robust  
202 magnetic characterization of atmospheric deposition throughout a large region of SW  
203 Europe and a reliable assessment of their fluctuations as a function of seasonal variations  
204 and underlying sources. Details on the location, type and monitoring period for the sites  
205 considered in this study can be found in Table 1 and in Pey et al. (2020).

206 Sample collection involved: 1) cleaning of the funnel using 500 ml of bi-distilled water  
207 and a plastic brush, so that the resulting solution was combined with any material that had  
208 accumulated in the container either in the form of dry (free fall) or wet (washed out by  
209 rainfall) deposition; and 2) the replacement of the containers by new ones. Once in the

210 laboratory, samples were stored in dark conditions at 4°C before filtering, which involved  
211 two steps. The first filtering step employed a 100 µm mesh to remove biological particles  
212 (mostly insect and plant fragments). The second filtering employed 47 µm-diameter  
213 quartz fibre filters (PALL tissuquartz 2500QAT-UP), which had been previously kept in  
214 controlled conditions ( $T^a$  22-24 °C; RH 25-35%) and weighed two times on different days.  
215 For each sample, between two and six sister filters were obtained depending on the  
216 amount of material collected (FRI-3.1 would indicate the first filter from sample 3  
217 collected in Frigiliana). The filtered water was subjected to chemical analyses to provide  
218 information concerning the soluble fraction of atmospheric deposition (Pey et al., 2020),  
219 which is not considered here. After drying at room temperature, each filter was weighed  
220 two times on different days. The insoluble PM collected by the filters, is the subject of  
221 the environmental magnetic analyses presented here.

222 Atmospheric deposition across the study area occurs under seasonally contrasting  
223 meteorological scenarios (Pey et al., 2013, 2020; Revuelta et al., 2014). The winter is  
224 characterized by either the advection of W and NW Atlantic low-pressure systems that  
225 bring rainfall, or by anticyclonic conditions with negligible precipitation. The spring and  
226 fall are dominated by highly dynamic atmospheric conditions associated to the arrival of  
227 both Atlantic and Mediterranean depressions. This situation contrasts with that prevailing  
228 during summer, when recirculation and mixing of local air masses, and the arrival of north  
229 African warm winds, dominate. The passage of Atlantic and Mediterranean depressions  
230 during the fall, winter and spring drive episodes dominated by wet deposition that often  
231 involve the transport of locally-produced particles away from their source. Atlantic  
232 depressions are sometimes associated with the arrival of southerly winds sourced in north  
233 Africa, a situation that is accompanied by high NAD loads that are washed out by rain  
234 mainly over the eastern half of the Iberian Peninsula and the Western Mediterranean  
235 basin. The most prominent NAD outbreaks reported during the monitoring period of the  
236 DONAIRE network occurred in November 2016 and February 2017, and affected most  
237 of the studied sites to variable extents (see Pey et al., 2020 and  
238 <http://www.bsc.es/projects/earthscience/visor/dust/med8/dep/archive/>). In contrast,  
239 winter anticyclonic conditions, cause nearly stagnant conditions throughout much of the  
240 studied area (Pey et al., 2020; Revuelta et al., 2014). Atmospheric dispersion is then  
241 limited, and locally-derived particles constitute the bulk of the then dominant dry  
242 deposition. One such long anticyclonic period occurred during December 2016 and the  
243 first two weeks of January 2017. Dry deposition also dominates over the summer, but the  
244 recirculation of local air masses results in mixing and redistribution of atmospheric

245 particles with different origins. Penetration of north African air masses also results in  
246 regular delivery of dust particles throughout the warm season, which accumulate mainly  
247 as dry deposition following a north-south decreasing gradient across the Iberian Peninsula  
248 (e.g., Pey et al., 2013). On some particular occasions during the summer (e.g., August  
249 2016), the advection of north African air masses can reach a considerable magnitude and  
250 result in the accumulation of significantly higher amounts of desert dust. All the  
251 aforementioned meteorological scenarios typically last between a few days and some  
252 weeks, and their signature is prone to be mixed at the studied sites given the biweekly  
253 sampling.

254

## 255 2.2. Rock magnetic characterization

256 Samples representative for all types of studied sites and atmospheric deposition  
257 scenarios were selected for detailed rock magnetic measurements, which were conducted  
258 with a Princeton Measurements Corporation vibrating sample magnetometer at the Black  
259 Mountain Paleomagnetic Laboratory of the Australian National University. Isothermal  
260 remanent magnetisation (IRM) acquisition curves were measured at more than 50 steps,  
261 approximately equally spaced on a logarithmic scale, from 1 mT up to 1 T (Kruiver et al.,  
262 2001). IRM curves were unmixed using the MAX UnMix web application (Maxbauer et  
263 al., 2016) to determine the main magnetic minerals contributing to the IRM. Hysteresis  
264 loops were measured with a maximum field of 1 T, after which a backfield magnetization  
265 was applied in a succession of fields of up to 1 T. Hysteresis parameters  $M_s$  (saturation  
266 magnetisation),  $M_{rs}$  (saturation remanent magnetisation),  $H_c$  (coercivity) and  $H_{cr}$   
267 (coercivity of remanence) were used to plot the results in the Day diagram (Day et al.,  
268 1977), which provides a rough inference on the magnetic domain state (and hence grain  
269 size) of the studied magnetic assemblages (Roberts et al., 2018). Due to the limited  
270 amount of material present in the filters, reliable IRM acquisition curves and hysteresis  
271 data could be only obtained with measurement averaging times greater than 10 s. Reliable  
272 results could not, however, be obtained for samples with the lowest atmospheric  
273 deposition loads, which correspond to background sites without prominent NAD  
274 contributions. A subset of samples, representative of the most relevant types of sites and  
275 deposition scenarios, were selected for measurement of irregular first order reversal curve  
276 (FORC) distributions (Zhao et al., 2015) following the approach of Zhao et al. (2017),  
277 which enables distinction of different domain state signatures.

278



### 279 2.3. SEM observations

280 Samples representative for all types of studied sites and atmospheric deposition  
281 scenarios were selected for the electron microscopy study to determine type, size and  
282 shape of particulate matter deposited on each filter. The observations have been carried  
283 out on small pieces (10 x 0.5 mm) taken from the central part of sister filters. The filters  
284 were placed with conductive adhesive over Al-stubs and coated with a Cressington 208  
285 carbon coater. Analyses were performed with a JSM-6010PLUS/LA InTouchScope™  
286 scanning electron microscope at IGME. The observation conditions involve high vacuum,  
287 20 kv, and 10 mm of working distance (WD). Most of the images have been taken with  
288 the backscattered electron detector to better estimate compositional variations of the  
289 particles of interest. Energy dispersive analysis (EDX) was performed on individual  
290 particles to obtain qualitative compositions.

291

### 292 2.4. Bulk environmental magnetic measurements

293 Bulk magnetic properties were measured, after folding and placing one sister filter for  
294 each sample in standard paleomagnetic boxes, at the Paleomagnetic Laboratory of the  
295 Institute Geosciences Barcelona (CCiTUB-CSIC), Spain. The magnetic susceptibility  
296 was measured with a KLY-2 magnetic susceptibility bridge with a field of 0.1 mT at a  
297 frequency of 470 Hz. The data provided represent the average of three measurements for  
298 each filter. Initial attempts to measure frequency-dependence of magnetic susceptibility  
299 (using a Bartington Instruments MS3 susceptibility meter coupled to a MS2B sensor)  
300 yielded unreliable results due to the very small amount of material accumulated on the  
301 filters (typically < 50 mg, Pey et al., 2020). We then measured two IRMs applied with an  
302 IRM10-30 pulse magnetizer (ASC Scientific) at fields of 0.3 and 1.2 T (IRM@0.3T and  
303 IRM@1.2T), which were measured using a 2G Enterprises SRM755 superconducting  
304 rock magnetometer. Due to technical problems, the anhysteretic remanent magnetisation  
305 (ARM) could not be measured prior to IRM analyses. A set of 5 blank samples was  
306 prepared with empty filters to measure their magnetic susceptibility, IRM@0.3T and  
307 IRM@1.2T. The average of these measurements was subtracted from the measured  
308 magnetic properties of all samples to account for the combined signal of the quartz filter  
309 and the paleomagnetic boxes, whose magnetic moment is (on average) an order of  
310 magnitude lower than that of the PM even in regional background sites. Bulk magnetic  
311 properties have been normalized by the mass of the material on the filters, but have been  
312 also considered as an absolute signal (e.g., magnetic moment) to provide an indication of  
313 the total load of magnetic particles. The total magnetic moment of each sample has been

314 calculated by assuming that the measured sister filters are representative for the magnetic  
315 properties of all the filters obtained for each sample.

316 We have used mass-specific magnetic susceptibility ( $\chi$ ) as an initial indicator of the  
317 concentration of magnetic minerals on the filters (Liu et al., 2012). In order to assess the  
318 relative contribution of low- (magnetite, maghemite) versus high-coercivity (hematite,  
319 goethite) minerals to the remanence, we have computed the forward S-ratio  
320 ( $\text{IRM}@0.3\text{T}/\text{IRM}@1.2\text{T}$ ) of Kruiver and Passier (2001).  $\text{IRM}@0.3\text{T}$  and the “hard” IRM  
321 ( $\text{HIRM} = \text{IRM}@1.2\text{T} - \text{IRM}@0.3\text{T}$ ) have been used as proxies for the concentration of  
322 low- and high-coercivity minerals, respectively (Liu et al., 2012). Finally, the ratio  
323  $\text{IRM}@1.2\text{T}/\chi$  has been used as an indicator of the grain size of the magnetic minerals in  
324 the filters (Peters and Dekkers, 2003; Liu et al., 2012).

325

## 326 2.5. Geochemical data

327 One sister filter for each sample was used to quantify the concentration of major and  
328 trace elements. We followed the method of Querol et al. (2001), which involves complete  
329 digestion of the filters in a  $\text{HF}:\text{HNO}_3:\text{HClO}_4$  solution, then drying and re-dissolution of  
330 the resulting mixture in  $\text{HNO}_3$  (5% concentration), and its analysis by Inductively  
331 Coupled Plasma Atomic Emission Spectroscopy (ICP-AES) and Inductively Coupled  
332 Plasma Mass Spectrometry (ICP-MS). Blank filters were analysed along with each set of  
333 samples in order to account for filter contribution. We have used the concentrations of  
334 Al, Ti, Cu and Sb for comparison with environmental magnetic properties in order to  
335 establish their link with two of the most relevant natural and anthropogenic sources of  
336 atmospheric deposition expected in the studied sites, namely NAD (Al and Ti, Revuelta  
337 et al., 2014) and vehicular traffic (Cu and Sb, Amato et al., 2009; Revuelta et al., 2014).  
338 Geochemical measurements were performed at the Geochemistry Laboratory of IGME at  
339 Madrid using a VARIAN VISTA-MPX ICP-EAS and an AGILIENT 7500 CE ICP-MS.

340

## 341 3. Results

### 342 3.1. Rock magnetic characterization

343 Unmixing of IRM acquisition curves requires fitting the data with up to four distinctive  
344 components whose characteristics are consistent across the studied samples (Fig. 2, Table  
345 2). Urban, industrial and agricultural samples are characterized by three components  
346 (Figs. 2a-d). Component 1 (C1) has a coercivity ( $B_h$ ) of less than 10 mT ( $\log_{10} B < 1$  in  
347 Fig. 2) and represents an average extrapolated contribution (EC) of 6% to the total IRM  
348 (Table 2). Component 2 (C2) has a coercivity ranging between 19 and 54 mT ( $\log_{10} B =$

349 1.2-1.8) and constitutes the bulk of the IRM, with an average contribution that is slightly  
350 larger in urban and industrial samples (78-96%) compared to the remaining settings (70-  
351 92%) (Table 2). Component 3 (C3) has an average coercivity of 760 mT ( $\log_{10} B > 2.7$ ),  
352 and represent a fraction of the IRM ranging between 1 and 9 % (Table 2). The only  
353 exception to this pattern is one industrial sample, where C1 was not identified (MON-4.1,  
354 Table 2). In all but one sample associated with prominent NAD events, a fourth  
355 component (C4) with an average coercivity of  $\sim 125$  mT ( $\log_{10} B = 2-2.2$ ) is required to  
356 provide an adequate fit to the data regardless of the site type (Fig. 2e-h). In these cases,  
357 the relative contribution of C3 increases slightly (5-9%) at the expense of C2, which  
358 shows a consistent shift to slightly lower coercivities ranging between 19 and 42 mT (Fig.  
359 2) (Table 2). The four identified components exhibit dispersion parameters (DP) ranging  
360 typically from 0.25 to 0.4, although some values smaller and larger than 0.2 and 0.4,  
361 respectively, are also observed. S values range from 0.83 to 1.3 for the four components.  
362 We attribute the somewhat large DP and S values to the very small amount of material  
363 present in the filters, which often results in noisy IRM acquisition curves (see grey dots  
364 in Fig. 2) that needed to be smoothed before unmixing. These values call for a cautious  
365 interpretation of results, with the possibility of a mixed mineralogy within a single  
366 component being the most likely flaw (Heslop et al., 2004). Based on their coercivities,  
367 C1 and C2 can be associated with either magnetite or maghemite. C3 corresponds to  
368 hematite and might include a small fraction of goethite that is magnetized by the 1.2 T  
369 field. C4 has a coercivity that may correspond either to magnetite/maghemite or to  
370 hematite.

371 Hysteresis curves for all the studied samples are characterized, after correction of the  
372 high-field paramagnetic slope, by narrow loops that saturate between 300-500 mT (Fig.  
373 3a-e).  $B_{cr}$  values range between 17 and 33 mT and are somewhat higher in urban and  
374 industrial sites compared to agricultural, background, or urban samples with a prominent  
375 NAD load (Table 2). These values are similar to those reported in many studies of  
376 anthropogenic PM (e.g., Sagnotti et al., 2006, 2009; Saragnese et al., 2011), and indicate  
377 that low-coercivity minerals such as magnetite or maghemite dominate the magnetic  
378 properties of the samples. A minor contribution of high-coercivity minerals (e.g.  
379 hematite) to hysteresis properties could be inferred from the somewhat wasp-waisted  
380 shape of some hysteresis loops (Roberts et al., 1995), especially those of samples with a  
381 significant NAD load (Fig. 3b, d). Representation of hysteresis data on the Day plot (Day  
382 et al., 1977) indicates that urban samples fall just above the MD field, suggesting the  
383 dominance of coarse magnetite/maghemite particles (Fig. 3f). Urban samples with a

384 prominent NAD contribution move towards lower  $B_{cr}/B_c$  values along the SD-MD mixing  
385 curve of Dunlop (2002), thereby indicating a relatively major contribution of finer SD or  
386 PSD (vortex) magnetic grains. Samples from industrial, agricultural and background sites  
387 with a prominent NAD load fall within the PSD domain, also indicating the relative  
388 importance of SD or vortex magnetic grains. Noticeably, all the studied samples fall  
389 slightly above the SD-MD mixing line of Dunlop (2002) regardless of sample type, which  
390 indicates the widespread contribution of ultrafine SP particles. Overall, hysteresis data  
391 confirm the complex mineralogical assemblage of the studied samples demonstrated by  
392 the IRM acquisition curves.

393 FORC distributions in urban site samples (Fig. 4a) are characterized by slightly  
394 asymmetric contour lines that diverge at a certain angle along the  $B_i$  axis and include a  
395 low coercivity (<25 mT) tail along the  $B_c$  axis. This feature is intermediate between those  
396 shown by MD particles, whose contour lines show an important divergence along the  $B_i$   
397 axis, and by vortex particles, whose contour lines tend to intersect the  $B_i$  axis at right  
398 angles (Zhao et al., 2017). The associated remFORC distributions have a positive peak  
399 along the  $B_c$  axis and a second, significantly smaller peak along the  $B_i < 0$  mT axis. iFORC  
400 distributions are characterised by two positive and negative peaks along the  $B_i > 0$  and  $B_i$   
401  $< 0$  axis, respectively, plus some weaker, additional positive and negative peaks in the  
402 remainder of the diagram. Finally, the tFORC distributions are characterized by two  
403 positive peaks close to the  $B_i > 0$  and  $B_i < 0$  axis forming a butterfly-shaped pattern. The  
404 iFORC and tFORC features indicate the presence of both vortex and MD particles,  
405 although the relative width of the butterfly-shaped pattern suggests a greater contribution  
406 from vortex particles (Zhao et al., 2017). The small peak along the  $B_i < 0$  mT axis in the  
407 remFORC points to thermal activation effects related with the presence of SP particles in  
408 urban samples. FORC distributions for samples from the MON industrial site are  
409 comparable to those found in urban sites, although the more divergent distribution along  
410 the  $B_i$  axis in the FORC diagram and the narrower butterfly-shaped pattern in the tFORC  
411 distribution points to a greater fraction of MD particles. In contrast to urban and industrial  
412 site samples, samples from background locations with significant NAD loads appear to  
413 be dominated by a SD fraction that produces closed contour lines around a peak at <20  
414 mT in the FORC distribution (Fig. 4b). Such SD particles are also inferred from the  
415 iFORC diagram, where a characteristic negative-positive-negative triple peak near the  $B_i$ -  
416  $B_c$  intersection can be identified despite of the surrounding noisy signal. The tFORC  
417 distribution is characterized by a butterfly-shaped pattern, which, along with the

418 asymmetric divergent distribution along the  $B_i$  axis observed in the FORC diagram,  
419 indicates vortex particles (Zhao et al., 2017).

420

### 421 3.2. SEM observations

422 SEM images and EDS analyses demonstrate the frequent occurrence of iron oxides  
423 with different morphologies and grain sizes (Fig. 5). The most common form of iron oxide  
424 in urban and industrial samples is constituted by irregular aggregates that range between  
425 2 and 40  $\mu\text{m}$  in length and are formed by individual particles with irregular or spherical  
426 shapes and sizes varying between 0.2 and 2  $\mu\text{m}$  (Fig. 5a-c). They also appear as isolated  
427 particles (Fig. 5d) or spherules that range between 0.5 and 10  $\mu\text{m}$  and sometimes show  
428 the characteristic dendritic surface attributed to magnetite spherules formed by rapid  
429 quenching (Isobe and Gondo, 2013) (Fig. 5e). At MON industrial site, we observed some  
430 spherical aggregates with evidence for cracking (Fig. 5b). Aggregates and spherules often  
431 contain minor amounts of Zn, Mn or Cu, which, along with the observed textures, point  
432 to magnetite particles of anthropogenic origin (Petrovský and Ellwood, 1999; Sagnotti et  
433 al., 2006, Saragnese et al., 2011). Minor amounts of such anthropogenic irregular  
434 aggregates and spherules have also been found in background and agricultural sites as  
435 well. Background and urban samples encompassing distinctive NAD outbreaks are  
436 characterized by iron oxides with different morphologies and grain sizes of about 0.5 to  
437 8  $\mu\text{m}$ . Some iron oxides show the platy, hexagonal habit diagnostic of hematite crystals,  
438 which sometimes appear attached to the surface of large phyllosilicate aggregates (Fig.  
439 5f). Other iron and iron-titanium oxide particles display elongated or equidimensional  
440 shapes with irregular, but smooth, surfaces (Fig. 5c). These features (Rodríguez-Navarro  
441 et al., 2018) point, along with the connection between these samples and distinctive NAD  
442 events, to an aeolian origin for these particles, that are also present in samples not related  
443 with important NAD events.

444

### 445 3.3. Bulk environmental magnetic properties

446 S-ratios from urban and industrial samples range typically between 0.94 and 1, in clear  
447 contrast with the lower (0.86-0.96) values found for background and agricultural sites  
448 (Fig. 6a) (Supplementary Table 1). Site median S-ratios between 0.96 and 0.98 in  
449 industrial and urban environments are significantly higher, based on their inter-quartile  
450 (IQ) ranges, than those of background and agricultural settings. These data indicate that  
451 a low-coercivity, magnetite-like mineral is dominating the magnetic properties of

452 atmospheric deposition at all sites, with a smaller contribution by hematite in urban and  
453 industrial sites that is relatively increased in background and agricultural areas.

454 The mass specific magnetic susceptibility ( $\chi$ ) of urban samples ranges mostly between  
455 3 and  $8 \times 10^{-6} \text{ m}^3/\text{kg}$ , which is in contrast to those found in background sites (typically  $<2$   
456  $\times 10^{-6} \text{ m}^3/\text{kg}$  and often exhibiting negative values) (Fig. 6b). This points to higher  
457 concentrations of magnetic minerals in urban areas compared to background contexts,  
458 where  $\chi$  is often dominated by diamagnetic contributions. Remanence values support  
459 these inferences, higher IRM@0.3T values, and hence magnetite abundances, are found  
460 in urban sites compared with those in background areas (Fig. 6c). The exception to this  
461 pattern is Ebro valley, where S-ratios,  $\chi$ , and IRM@0.3T values of the urban (ZGZ) and  
462 industrial (MON) sites are somewhat smaller than in other cities, while also being  
463 comparatively higher in background (and agricultural) sites. This results in the overlap of  
464 IQ ranges between urban/industrial and background/agricultural sites. HIRM tends to be  
465 higher in urban/industrial than in background/agricultural sites although the differences  
466 are small (Fig. 6d), indicating that the concentration of hematite, as opposed to magnetite,  
467 is comparable across the set of studied environments.

468 IRM@1.2T/ $\chi$  values of urban samples cluster around site median values that range  
469 between 4.6 and 7.7 kA/m (Fig. 6e). This contrasts with IRM@1.2T/ $\chi$  of background and  
470 agricultural sites, which varies between -19 and  $>29$  kA/m. The combination of negative  
471 and exceedingly large values is an artefact caused by the either negative or near-zero  
472 susceptibilities in these samples, which results in unreliable IRM@1.2T/ $\chi$  estimates with  
473 no physical meaning. In contrast, the narrow range of IRM@1.2T/ $\chi$  found in urban  
474 samples, where  $\chi$  is dominated by large concentrations of magnetite, point to a narrow  
475 grain size either spanning the SP/SD boundary ( $<30$  nm) or the coarse MD state (10-20  
476  $\mu\text{m}$ ) (see Peters and Dekkers, 2003). Combined hysteresis, FORC and SEM data suggest  
477 that the latter situation is much more likely. Slightly larger IRM@1.2T/ $\chi$  values in  
478 industrial samples (median of 12 kA/m) point to finer MD magnetite grains between 2  
479 and 10  $\mu\text{m}$  (Peters and Dekkers 2003).

480 Time variations in mass-specific magnetic parameters in cities and their closest  
481 background sites reveal significant features (Fig. 7). Thus,  $\chi$  shows little variation  
482 throughout the studied period at all urban sites with the exception of the winter season,  
483 when  $\chi$  increases significantly. Significant  $\chi$  peaks are also observed during fall and  
484 spring in BCN and PMI. As expected,  $\chi$  variations are broadly mimicked by changes in  
485 IRM@0.3T. The situation is different in background sites, where the only remarkable

486 features in  $\chi$  are the decreases observed during winter at all sites except HJM. Minor  
487 IRM@0.3T variations are observed throughout the year at all background sites with the  
488 exception of CHI, where a significant increase is observed during the winter. No  
489 distinctive HIRM variations are observed at any of the sites throughout the year. It is  
490 noticeable that most variations observed in  $\chi$  and IRM@0.3T are contained within the  
491 protracted period of winter anticyclonic conditions experienced by the Iberian Peninsula  
492 between December 2016 and January 2016, which led to stagnant conditions and dry  
493 accumulations of very low amounts of PM. It is also noticeable that prominent NAD  
494 outbreaks, signalled by clear peaks in the amount of PM, did not alter significantly the  $\chi$ ,  
495 IRM@0.3T and HIRM values.

496

### 497 3.4. Geochemical data

498 Cu loads, expressed as fluxes to account for differences in sampling time and collector  
499 area, span an order of magnitude regardless of site type (Fig. 8). The correlation between  
500 Cu and IRM@0.3T, also expressed as a flux, is moderate, poor or even non-existent in  
501 most cases, although at some sites it appears to be meaningful (Fig. 8D) (Table 3). The  
502 correlation of IRM@0.3T with Sb, another proxy for vehicular PM (Revuelta et al., 2014,  
503 more), is somewhat weaker, even for urban sites (Table 3). The correlation between Al  
504 and HIRM is significantly stronger in most cases, regardless the type of site, a situation  
505 that is also found for the correlation between HIRM and Ti, another proxy for NAD loads  
506 (Revuelta et al., 2014). These correlations are better elucidated when Cu, IRM@0.3T, Al  
507 and HIRM fluxes are plotted as a function of time for all the studied sites (Figs. 9 and  
508 10). The high correlation coefficients (R) observed between Al and HIRM are reflected  
509 in the close similarity observed in the variations of both parameters (Fig. 9). Thus, they  
510 do not only consistently delineate the most distinctive NAD events, but also smaller  
511 variations throughout the rest of the monitored period (Fig. 9). Cu and IRM@0.3T do not  
512 covary at most sites as clearly as Al and HIRM do (Fig. 10), and also appear to be affected  
513 by some of the most distinctive NAD events.

514

## 515 **4. Discussion**

### 516 4.1. Origin of magnetic minerals in atmospheric deposition

517 IRM component C1 is similar to a low coercivity ( $\sim 10$  mT) component associated by  
518 Sagnotti et al. (2006) to large MD magnetite/maghemite grains that result from the  
519 abrasion of metallic parts of cars, especially brakes and tyres in Rome. Hysteresis (Fig.  
520 3), FORC (Fig. 4a,b) and SEM (Fig. 5d) data demonstrate the presence of large MD grains

521 in urban and industrial sites and, hence, validate the association of C1 with MD  
522 magnetite/maghemite grains derived mainly from vehicular traffic in urban and industrial  
523 areas. C1 is also present in agricultural and background sites (Fig. 2); while C1 in the  
524 former might be related to agricultural vehicles, its presence in background sites suggests  
525 that at least part of the emissions produced by traffic in urban and industrial environments  
526 reaches these sites via atmosphere mixing.

527       Regarding IRM component C2, which is characterized by coercivities between 19 and  
528 54 mT and is therefore linked to magnetite/maghemite, our data indicate that it constitutes  
529 the bulk of the magnetization in all studied samples (Fig. 2). The observation that  
530 IRM@0.3T values are distinctly higher in urban and industrial sites compared to  
531 agricultural and background areas (Fig. 6) suggests that this dominant  
532 magnetite/maghemite component is related to anthropogenic activity. This interpretation  
533 is further supported by hysteresis (Fig. 3), FORC (Fig. 4a,b) and SEM (Fig. 5a-e) data,  
534 which collectively show that vortex and MD magnetite/maghemite grains of  
535 anthropogenic origin dominate the magnetic assemblage of the studied samples. Given  
536 that vehicular exhausts are the main source of pollution in urban areas (Muxworthy et al.,  
537 2003; Sagnotti et al., 2006, 2009; Saragnese et al., 2011; Revuelta et al., 2014), and that  
538 C1 is related to abrasion of vehicular brakes and tyres, the most likely interpretation is  
539 that C2 is linked to vehicular exhaust emissions. It should be noted, however, that IRM  
540 components with average coercivities similar to those of C2 have been attributed to  
541 pedogenic (Maxbauer et al., 2016) or aeolian (Sagnotti et al., 2006; Font et al., 2018)  
542 magnetite particles, which raises the question of whether this component results from a  
543 combination of anthropogenic and natural sources as has been previously documented in  
544 other European cities (Spasov et al., 2004). Three lines of evidence support the  
545 hypothesis that C2 includes a mixture of natural and anthropogenic particles: 1) C2 has  
546 somewhat lower coercivities (19 to 42 mT) in samples with significant NAD loads,  
547 regardless of their type (Fig. 2, Table 1), when compared with standard urban and  
548 industrial samples (36-54 mT); 2) aeolian magnetite particles are often reported at urban  
549 sites in association with irregular magnetite/maghemite aggregates of anthropogenic  
550 origin (Fig. 5c); and 3) the somewhat large DP values associated to C2 are prone to result  
551 from the combination of more than one magnetic assemblage in a single component  
552 (Heslop et al., 2004). We interpret that in urban and industrial sites, aeolian magnetite is  
553 masked by a larger fraction of anthropogenic magnetite particles. The slightly higher  
554 coercivity of the latter would be linked to the presence of coarse MD magnetite grains  
555 rimmed by a layer of interacting, SP maghemite particles (Muxworthy et al., 2003;



556 Sagnotti et al., 2009; Saragnese et al., 2011; Revuelta et al., 2014). As with C1, the  
557 presence of a pervasive C2 in background sites suggest that emissions produced by traffic  
558 in urban and industrial environments reaches these sites via atmospheric mixing. When  
559 NAD loads are important, it is the larger, softer aeolian magnetite that masks the  
560 anthropogenic magnetite/maghemite even in urban sites, resulting in an overall shift to  
561 lower coercivities. Hysteresis (Fig. 3) and FORC (Fig. 4) data indicate that this aeolian  
562 magnetite occurs mainly as SD or vortex particles, in agreement with results from  
563 Revuelta et al. (2014).

564 IRM component C3 has a high coercivity (average of 740 mT) that is typically linked  
565 with aeolian hematite (Maxbauer et al., 2016; Font et al., 2018) and a grain size mostly  
566 spanning the SD range (Roberts et al., 2020). The presence of this component at all sites,  
567 including urban areas where production of local dust appears unlikely due to the dearth  
568 of bare soil, suggests a Saharan origin for this aeolian component, which is supported by  
569 its larger relative importance in samples with significant NAD loads (Fig. 2, Table 1).  
570 The main Saharan origin for this hematite component is apparent in the Ebro Basin, where  
571 the background magnetic signal appears to be enriched in high-coercivity minerals, as  
572 indicated by the excellent correlation observed in this area between HIRM and proxies  
573 for NAD abundances (Al and Ti) (Fig. 9b, Table 3).

574 With regard to IRM component C4, its exclusive association with significant NAD  
575 loads demonstrates a clear aeolian origin. Its coercivity (~125 mT) appears to be too hard  
576 for non-biogenic magnetite, but would be consistent with hematite grains larger (MD)  
577 than higher-coercivity hematite associated to C3 (Roberts et al., 2020). In this regard, it  
578 is important to note that C3 (mainly SD hematite) appears to be linked to background  
579 NAD fluxes identified in all the samples, whereas C4 (likely MD hematite) appears to be  
580 linked to only the major NAD outbreaks. This points to the arrival of hematite from a  
581 different Saharan source during these major periods. We tentatively link the background  
582 hematite of C3 to the coastal areas of the western Sahara (dust source area 2 of Scheuven  
583 et al., 2013) and the additional hematite of C4 to the interior areas of southern Algeria  
584 and northern Mali (dust source area 3 of Scheuven et al., 2013), where warmer and drier  
585 conditions favour enhanced hematite formation (Larrasoña et al., 2015) in surface silts  
586 that constitute the main dust source (Prospero et al., 2002).

587

#### 588 4.2. Environmental significance of bulk magnetic properties

589 In order to use bulk magnetic parameters as proxies for natural or antropogenic  
590 processes, their environmental significance needs to be determined. Of particular

591 importance to this study are IRM@0.3T and HIRM, which are candidate proxies for  
592 monitoring pollution and NAD fluxes via concentrations of anthropogenic  
593 magnetite/maghemite and natural (aeolian) hematite, respectively. The correlation  
594 between IRM@0.3T and proxies for anthropogenic pollution (Cu or Sb) is not good for  
595 all sites (Fig. 8, Table 3), and this might be explained by different factors. First, Cu and  
596 Sb loads are preferably incorporated into the soluble fraction of atmospheric deposition,  
597 whereas anthropogenic magnetite/maghemite concentrations correspond to PM  
598 accumulated on the filters. Second, Cu and Sb are mainly linked to vehicular traffic  
599 (Amato et al., 2009; Revuelta et al., 2014), whereas anthropogenic magnetite/maghemite  
600 particles reported here may also be sourced from industrial activities around urban areas.  
601 Third, C2 appears to be a mixture of natural magnetite and anthropogenic  
602 magnetite/maghemite particles, whose separation by means of the measured bulk  
603 magnetic properties is not possible. These circumstances indicate, in a context where  
604 NAD supply occurs throughout the year (Pey et al., 2013, Varga et al., 2014), that  
605 inferences concerning anthropogenic pollution based on IRM@0.3T or other proxies for  
606 magnetite abundance should be treated with caution.

607 With regards to hematite, no evidence for its occurrence as an anthropogenic mineral  
608 has been found in our study, in line with most environmental magnetic studies of PM  
609 (Petrovský and Ellwood, 1999; Muxworthy et al., 2002, 2003; Spassov et al., 2004;  
610 Sagnotti et al., 2006, 2009; Saragnese et al., 2011; Revuelta et al., 2014; Hofman et al.,  
611 2017; Tan et al., 2018). Therefore, the reliability of HIRM as a proxy for NAD flux relies  
612 on the choice of the magnetic fields used to impart the laboratory remanences from which  
613 HIRM is estimated. Thus, using IRM@1.2T and IRM@0.3T to calculate HIRM means  
614 that remanence carried by hematite with a coercivity <300 mT will be lost and result in  
615 underestimation of hematite content (Roberts et al., 2020). This will have occurred with  
616 the low-coercivity tail of C3 and the majority of C4 (see Fig. 2). Yet, the choice of a 0.3  
617 T field to estimate HIRM ensures that the contribution of the high-coercivity tail of C2 to  
618 HIRM, most likely associated with anthropogenic magnetite by virtue of its higher  
619 coercivity, is minimal so that the correlation of HIRM with NAD proxies is still apparent  
620 (Fig. 8, Table 3). Our results demonstrate that although HIRM based on IRM@0.3T  
621 misses part of the genuine NAD hematite contribution, it avoids contamination by a  
622 spurious, anthropogenic signal and, thus, can be used as a proxy for NAD flux.

623

624 4.3. Temporal variations in NAD and pollution

625 HIRM and Al contents indicate that the flux of NAD is somewhat higher during  
626 summer at most network sites compared with winter (Fig. 9a-c). Superimposed on this  
627 overall trend of background NAD are major dust outbreaks identified at different  
628 monitoring stations, mainly at the beginning and end of winter but also during summer.  
629 These events are delineated by sharp peaks in both hematite and elemental (Al and Ti)  
630 fluxes that span typically one (e.g. ORD, ADV) or two (FRI, BCN) samples (Fig. 9).  
631 Exceptions to this are background sites ENO and ORG, in the westernmost part of the  
632 Cantabrian-Pyrenean region, and industrial and urban sites MON and PMI, respectively.  
633 Very low NAD fluxes occurred at ENO and, to lesser extent, ORG, where only the two  
634 major NAD outbreaks identified in late fall and late winter are detected. At MON and  
635 PMI, most of the winter season is characterised by sustained high fluxes of hematite that  
636 are somehow decoupled from Al fluxes. A possible explanation for such decoupling  
637 would be an additional, anthropogenic source of hematite linked to industrial pollution in  
638 MON, and potentially leisure cruise activity in PMI (Cerro et al., 2020). Overall, HIRM  
639 values corroborate the larger NAD background reported in previous studies for the Iberian  
640 Peninsula during the summer, and also the preferential transport of NAD over the eastern  
641 half of the peninsula and the Mediterranean basin (Querol et al., 2009; Varga et al., 2014;  
642 Pey et al., 2013, 2020).

643 In contrast to HIRM, IRM@0.3T values show no significant variations throughout the  
644 year. At ORD, ZGZ and EDC, a slight increase in IRM@0.3T may occur during summer  
645 (Fig. 10). Such an increase appears to occur in parallel with an increase in hematite flux,  
646 raising the question of whether such a subtle rise in magnetite fluxes is linked, at least  
647 partially, to supply of background NAD rather than anthropogenic (pollution) fluxes. This  
648 may be the case because most of the major NAD outbreaks occurred during late autumn  
649 and late winter, which in many cases are associated with a noticeable increase in aeolian  
650 magnetite flux. Yet, the large Cu peaks associated with some (but not all) of these  
651 outbreaks suggest that part of the rise in magnetite abundances is linked to pollution,  
652 which might be explained by the compression of the mixing layer during NAD outbreaks  
653 and the resulting magnification of pollution (Pandolfi et al., 2014; Cerro et al., 2020). An  
654 alternative explanation to this parallel increase in magnetite and Cu might be entrainment  
655 of pollution by dust-laden desert winds as they pass over urban and industrial areas of  
656 NW Africa on their way to the Iberian Peninsula (Rodríguez et al., 2011). Regardless, the  
657 clearest indication that background IRM@0.3T values are ultimately linked to a  
658 dominant, anthropogenic source is that urban (PAM, ZGZ, BCN, PMI) and industrial  
659 (MON) sites have overall higher IRM@0.3T values when compared with their nearest

660 background locations (ORG, CHI, MSY, HJM). The rather constant IRM@0.3T values  
661 observed throughout the year at most sites point to a steady supply of pollution or,  
662 alternatively, indicate that magnetite/maghemite fluxes vary at timescales shorter (e.g.  
663 working days versus weekends) than our biweekly sampling period.

664

## 665 **5. Conclusions**

666 Our study has enabled identification of three main magnetic mineral components in  
667 the particulate fraction of atmospheric deposition collected at different environments in  
668 Spain. Components C1 and C2 have low coercivities and are associated mainly with  
669 coarse (PSD/vortex and MD) magnetite grains, likely coated with SP maghemite. These  
670 anthropogenic particles are derived mostly from vehicular traffic and, possibly also,  
671 industrial activities, which is consistent with findings for most European cities  
672 (Muxworthy et al., 2002, 2003; Spassov et al., 2004; Sagnotti et al., 2006, 2009;  
673 Saragnese et al., 2011; Sagnotti and Winkler, 2012; Revuelta et al., 2014, Hofman et al.,  
674 2017). Our results indicate that anthropogenic magnetite/maghemite particles are  
675 ubiquitous in pristine environments because of atmospheric transport. The third  
676 component, C3, is related to SD hematite of aeolian origin and appears to represent a  
677 baseline of NAD transported from the Sahara throughout most of the year. Such a natural  
678 hematite fraction has been identified in other large cities within the Mediterranean region  
679 of Europe (e.g. Rome, Sagnotti et al., 2006), and is likely to correspond to the subdued  
680 (and often overlooked) high-coercivity component reported in other southern European  
681 cities (Saragnese et al., 2011; Revuelta et al., 2014). When samples include distinctively  
682 intense events of NAD deposition, a fourth component associated with coarser aeolian  
683 hematite particles, possibly sourced from different areas within the Sahara, is also  
684 identified regardless the site type.

685 Our results demonstrate that part of the magnetite/maghemite associated to C2 appears  
686 to be of aeolian origin, a circumstance that has been reported in previous studies (Spassov  
687 et al., 2004; Sagnotti et al., 2006; Revuelta et al., 2014). This seems to explain, along with  
688 other factors, the somewhat moderate correlation observed between  
689 magnetite/maghemite abundances and proxies for vehicular traffic, and demonstrates the  
690 need for caution when interpreting environmental magnetic proxies for  
691 magnetite/maghemite abundances in terms of anthropogenic loads. This would be  
692 especially the case in southern European cities, where a steady supply of NAD occurs  
693 throughout most of the year, and possibly also in other regions (middle East, SE Asia,  
694 Australia, SW USA) located near deserts (see Prospero et al., 2002). In contrast, our

695 results show a reasonable correlation between hematite abundance and geochemical  
696 proxies for NAD supply (Al and Ti), and therefore indicate that environmental magnetic  
697 proxies of hematite abundances (HIRM) can be used to monitor the contribution of  
698 natural sources to atmospheric deposition.

699

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714

## 715 **Author contributions**

716 J.C. Larrasoña and J. Pey designed the study and performed the data analysis.  
717 Sampling and maintenance of the DONAIRE network was performed by J.C. Larrasoña  
718 (Navarra), J. Pey and T. Mochales (Aragón), N. Pérez (Cataluña), J.C. Cerro and M.L.  
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720 acquisition curves, hystereis and FORC measurements were performed by X. Zhao and  
721 D. Heslop, whereas bulk magnetic measurements were conducted by J.C. Larrasoña, E.  
722 Beamud and T.Mochales. J. Reyes and M.P. Mata were responsible for the obtention and  
723 interpretation of geochemical and SEM data, respectively. J.C. Larrasoña wrote the  
724 article with inputs from all the co-authors in different sections of the article.

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870

871 **Figure captions**

872 **Figure 1.** Sampling procedures. (A) Location of the studied sites. (B) Example of a  
873 collector installed on an air quality station. (C) Plastic containers partially filled with  
874 water after sampling. (D) Detail of the filtering process. (E) Example set of filters.

875

876 **Figure 2.** Unmixing of selected IRM acquisition curves using MAX UnMix (Maxbauer  
877 et al., 2016). Grey dots and the yellow curve represent the smoothed IRM data and  
878 modelled coercivity distribution, respectively. Shaded areas represent 95% confidence  
879 intervals associated with each component. (A) to (D) show representative samples from  
880 urban, industrial and agricultural sites. (E) to (H) show representative samples from  
881 background, urban and agricultural sites with a significant contribution from North  
882 African Dust (NAD).

883

884 **Figure 3.** Hysteresis loops of representative samples collected in urban (A), industrial  
885 (B) and agricultural (C) sites. (D) and (E) correspond to urban and background samples  
886 with a significant contribution from North African Dust (NAD). (F) Day diagram of the  
887 studied samples according to their type.

888

889 **Figure 4.** Sequence of different types of FORC diagrams measured for representative  
890 urban (A) and background samples with a significant contribution from North African  
891 Dust (NAD) (B), measured following Zhao et al. (2017). SF denotes the smoothing factor,  
892 and NF stands for non-linear factor in the colour scale.

893

894 **Figure 5.** Back-scattered electron microscope images of selected samples. (A)  
895 Anthropogenic iron oxide aggregate in urban sample BCN-6. (B) Anthropogenic,  
896 spherical iron oxide aggregate in industrial sample MON-10. (C) Mixture of  
897 anthropogenic iron-oxide aggregates (white arrows) and aeolian Fe-Ti oxide particles  
898 (dashed white arrows) in an urban sample with significant NAD load (PAM-12). (D)  
899 Large, anthropogenic iron-oxide particle (white arrow) in urban sample BCN-6. (E)  
900 Anthropogenic magnetite dendritic spherule from the sample shown in C9 (PAM-12). (F)  
901 Platy iron-oxides, often showing an hexagonal habit, attached to a large aeolian particle  
902 (white arrow) in a background sample with significant NAD load (ORD-13).

903

904 **Figure 6.** Bulk environmental magnetic parameters for the studied sites. (A) S-ratio. (B)  
905 Mass-specific magnetic susceptibility. (C) Mass-specific IRM@0.3T. (D) Mass-specific

906 HIRM. (E)  $IRM@1.2T/\chi$ . Red, black, orange and green symbols denote urban, industrial,  
907 agricultural and regional background sites, respectively.

908

909 **Figure 7.** Temporal variations in mass-specific environmental magnetic parameters for  
910 the studied urban sites, which are shown along with their nearest regional background  
911 site. Orange arrows indicate samples affected by significant NAD loads (Pey et al., 2020).

912

913 **Figure 8.** Scatter plots of  $IRM@0.3T$  versus Cu fluxes, and of HIRM versus Ti fluxes,  
914 for samples representative for urban (A), agricultural (B), industrial (C) and background  
915 (D) sites.

916

917 **Figure 9.** Temporal variations in the fluxes of HIRM and Ti for the studied sites in(A)  
918 the Cantabrian-Pyrenean region, (B) the Ebro Valley, and (C) the Mediterranean region.  
919 Orange arrows mark significant NAD breakouts as identified by Pey et al. (2020).

920

921 **Figure 10.** Temporal variations in  $IRM@0.3T$  and concentration of Cu for the studied  
922 sites. Both magnetizations and elemental concentrations are normalized by surface and  
923 day. Orange arrows mark significant NAD breakouts as identified by Pey et al. (2020).

924

925 **Table 1.** Characteristics of the studied sites, which include the abbreviations used  
926 throughout the paper, the type, geographical latitude and longitude, altitude, annual  
927 rainfall (Pey et al., 2020), and start and end of the monitoring period for each site.  
928 Asterisks mark the natural sites that act as the regional background for the urban and  
929 industrial sites listed immediately above.

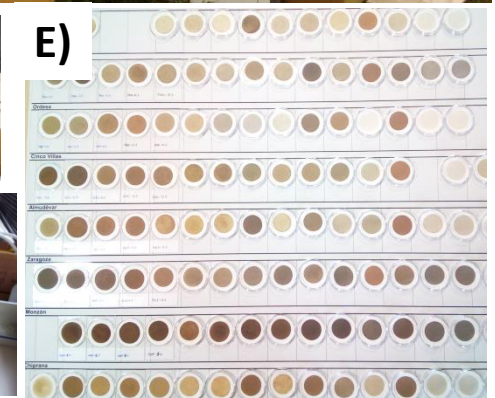
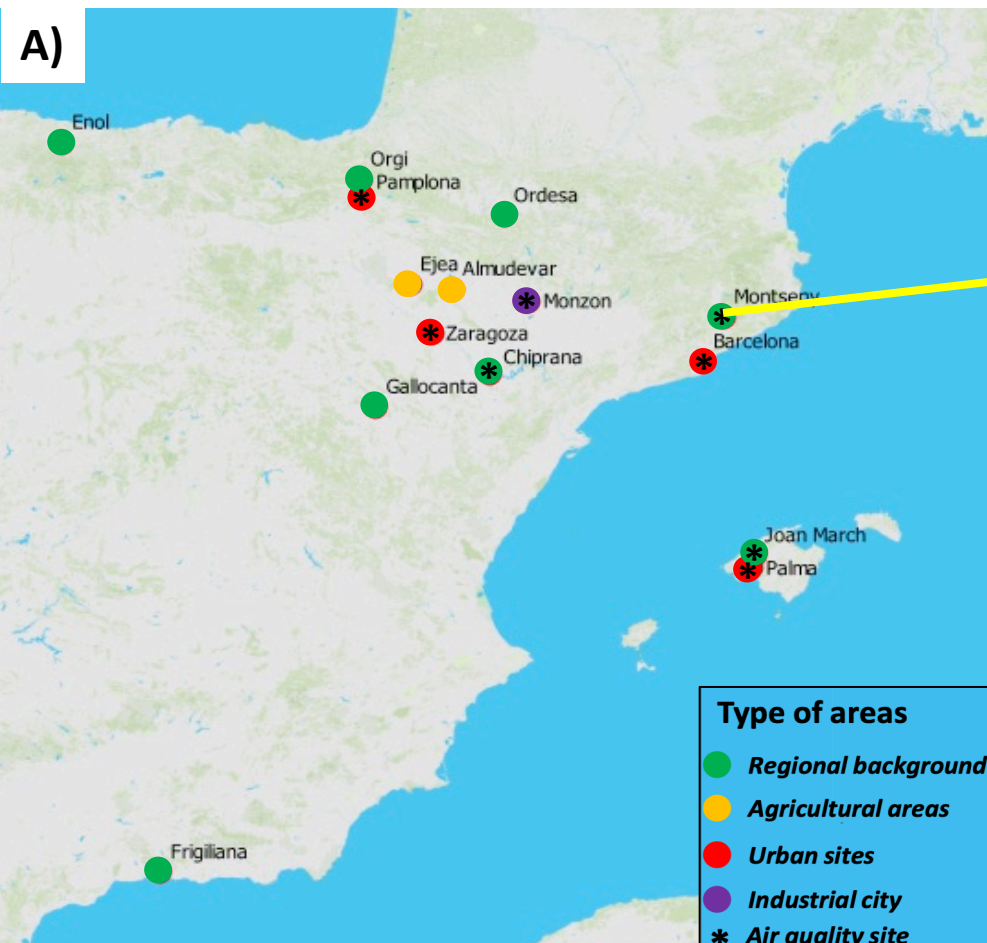
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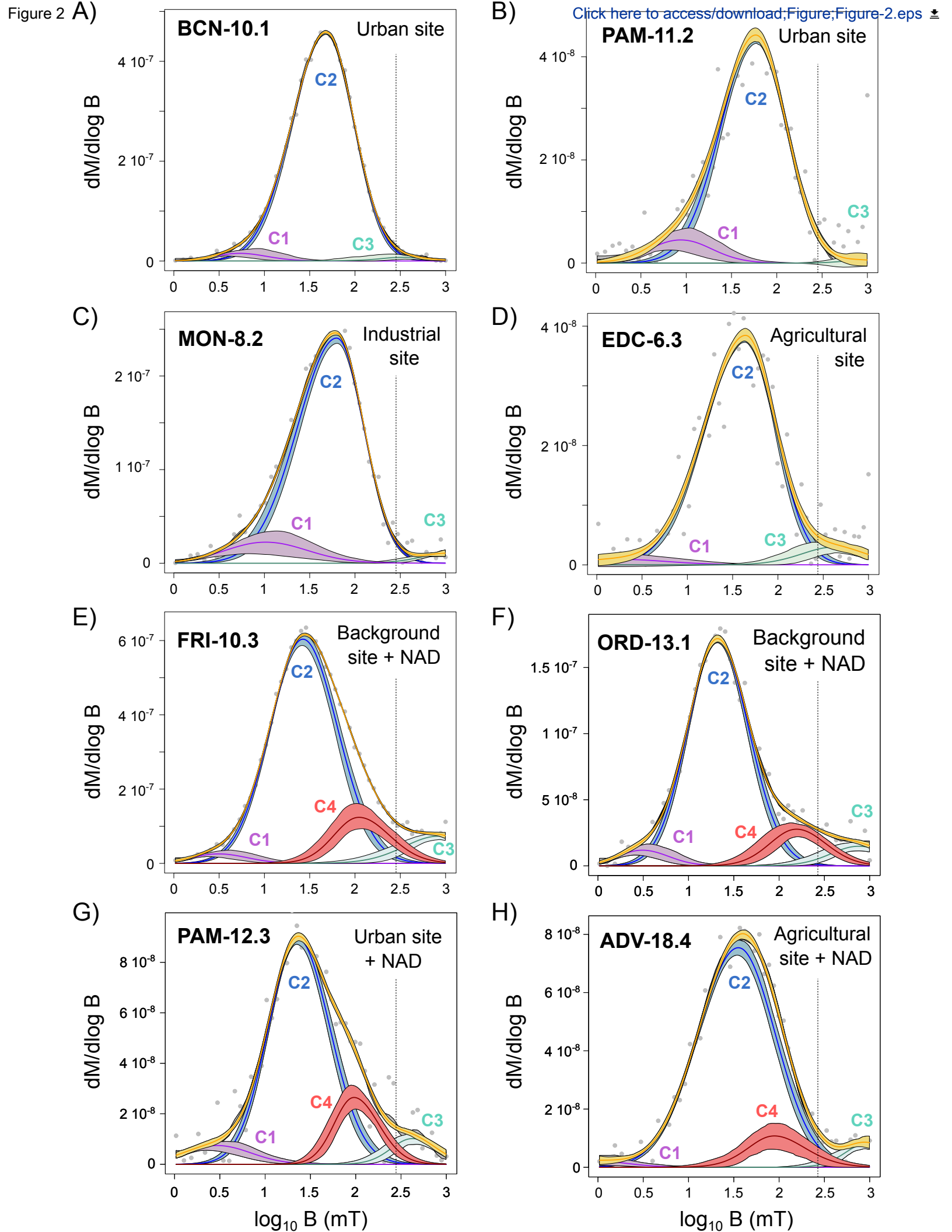
931 **Table 2.** Hysteresis and IRM unmixing data for the studied samples. *Bh*, *DP*, *S* and *EC*  
932 stand for the mean coercivity, dispersion parameter, skewness and extrapolated  
933 contribution associated to each component.

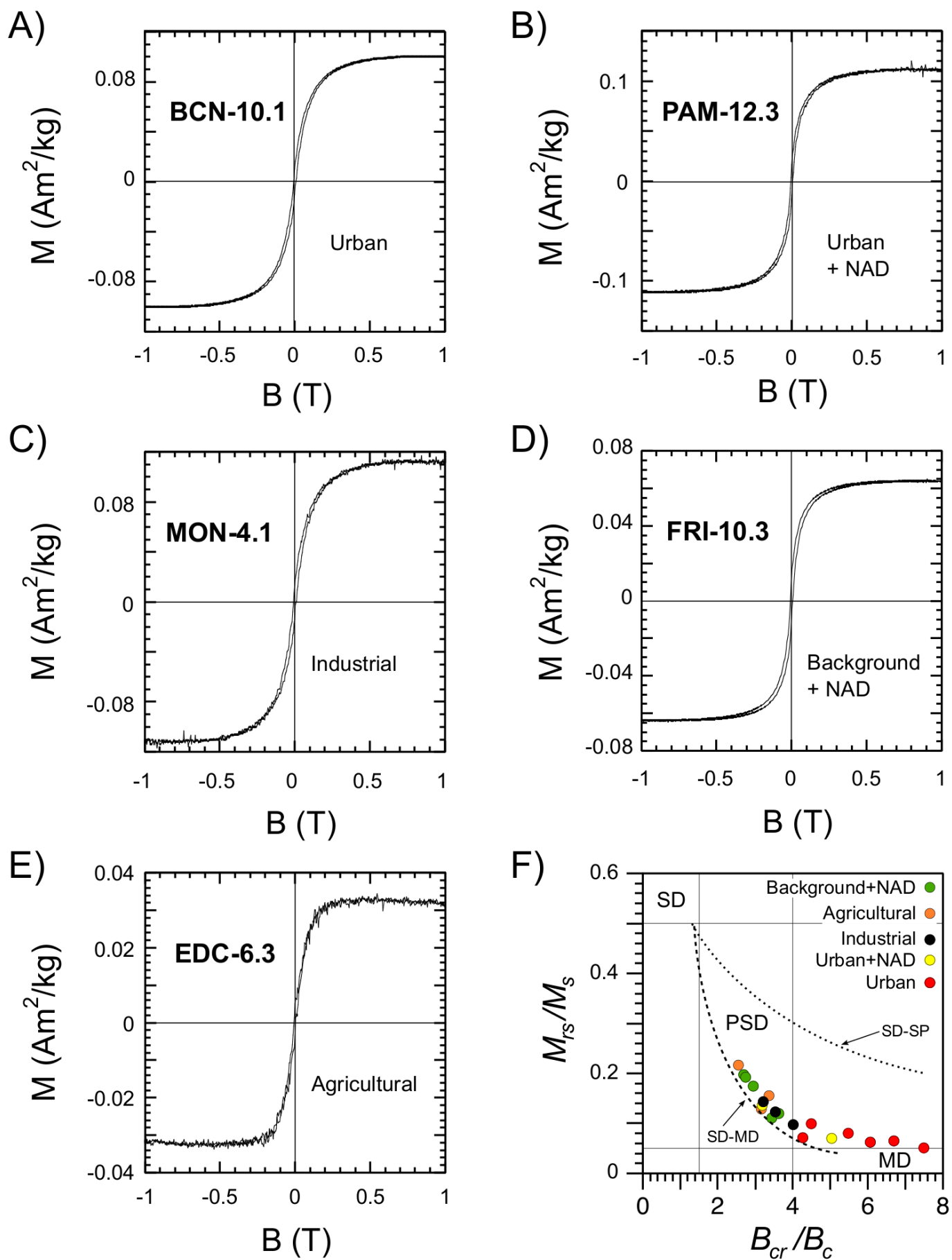
934

935 **Table 3.** Correlation coefficient (R) between  $IRM@0.3T$  and Cu,  $IRM@0.3T$  and Sb,  
936 HIRM and Al, and HIRM and Ti fluxes for the studied sites. Bold numbers denote R  
937 values larger than 0.7, considered to represent a strong correlation.

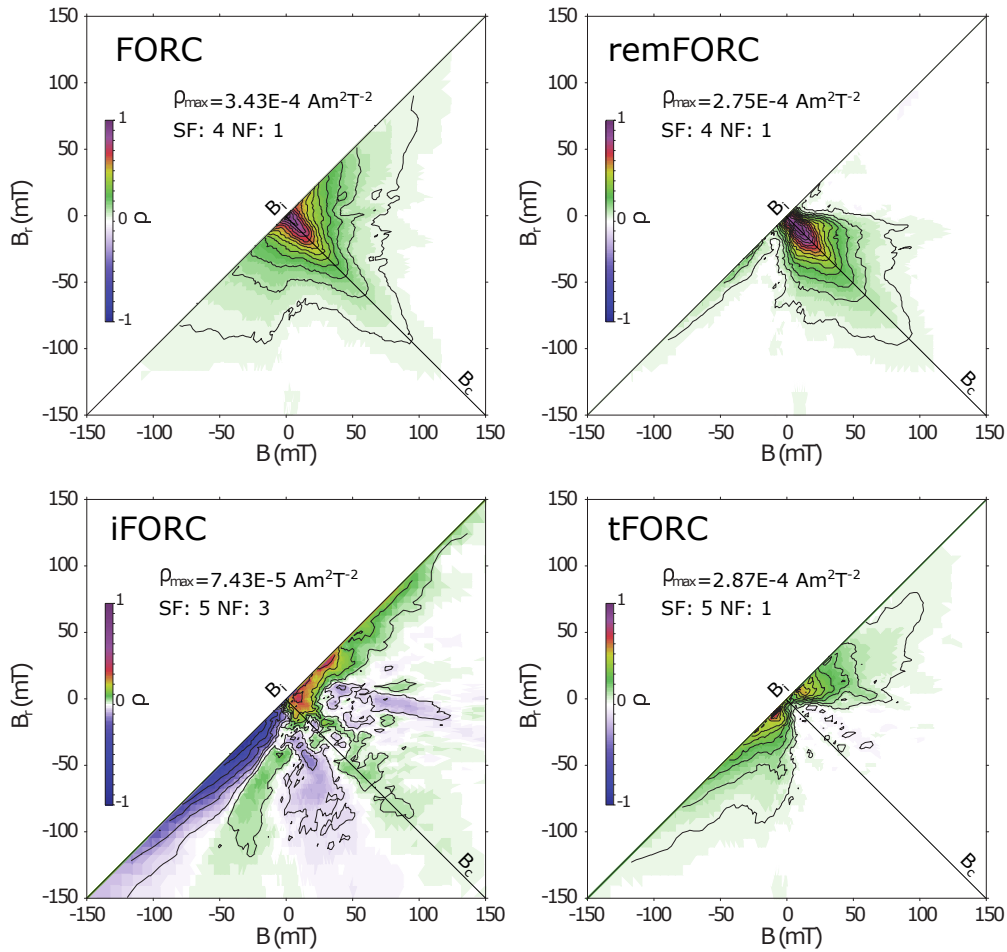
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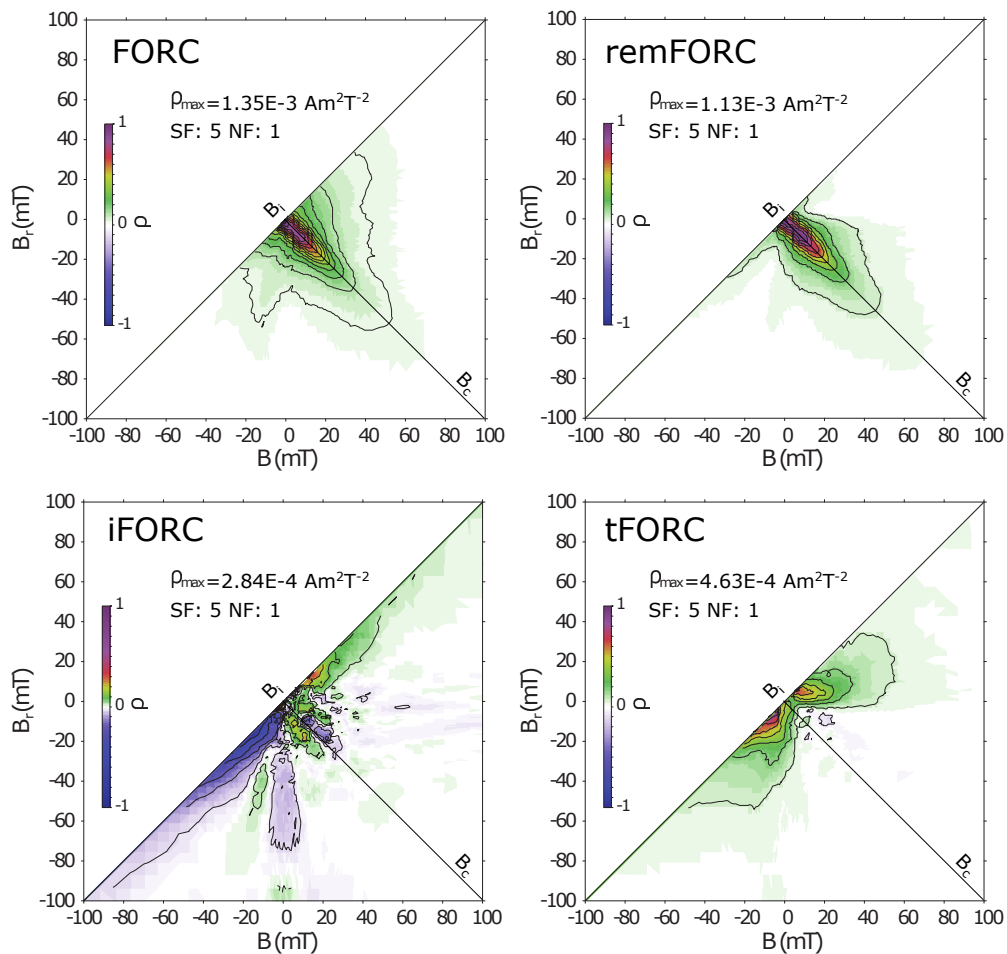




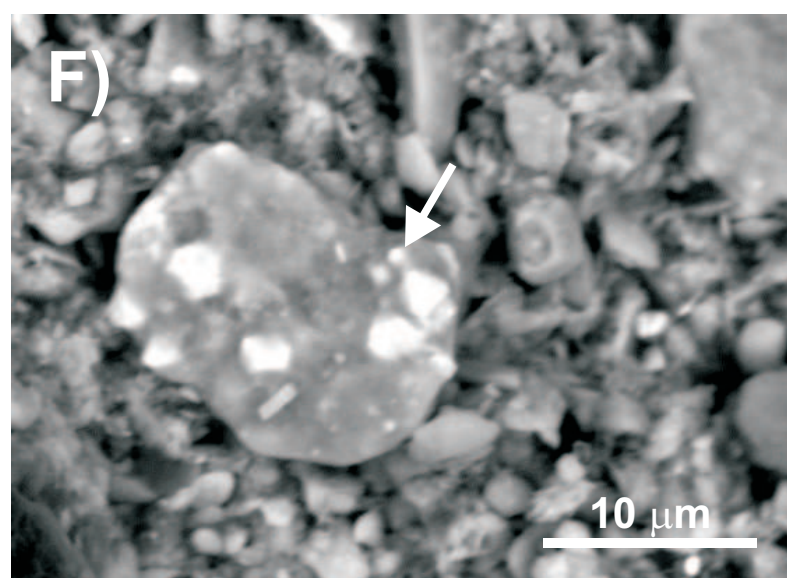
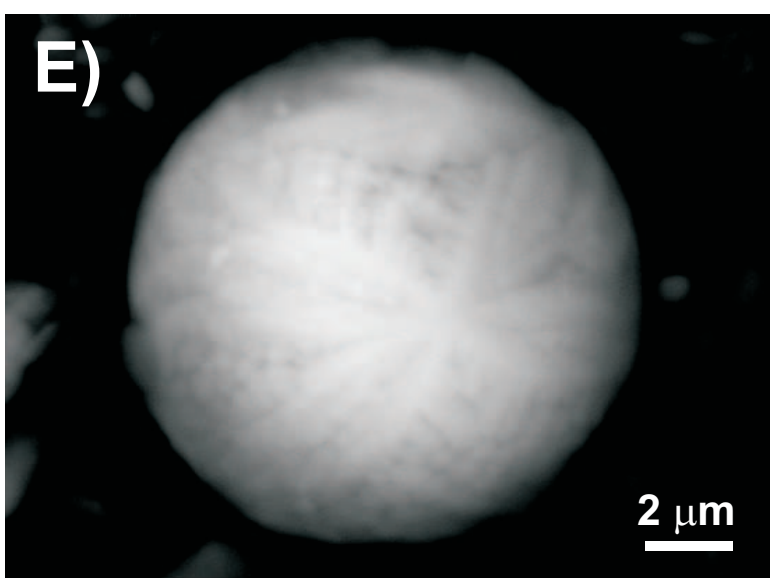
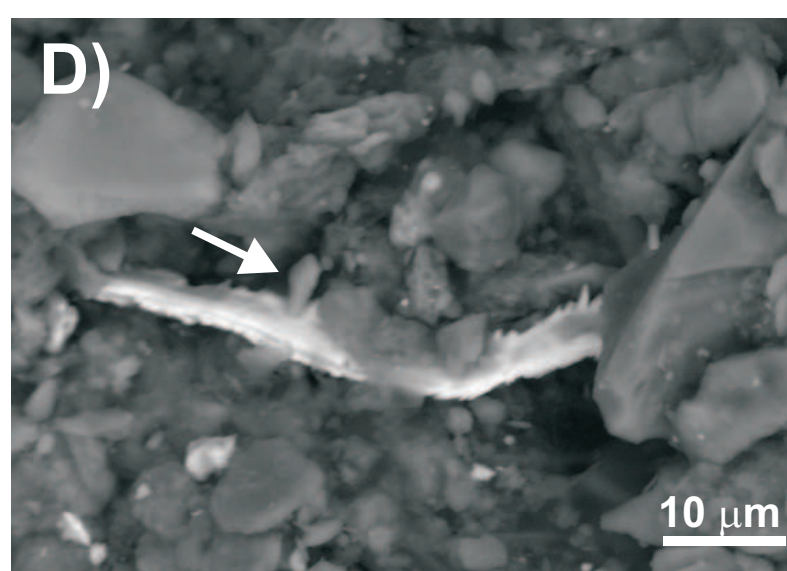
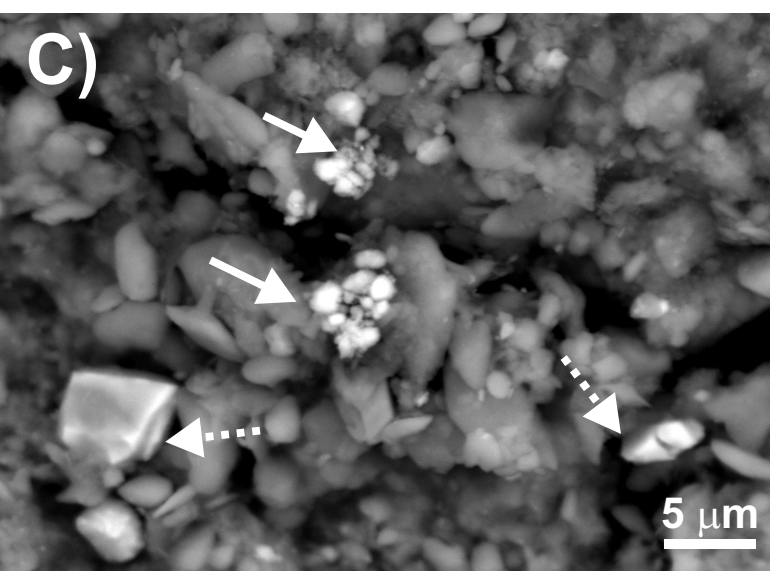
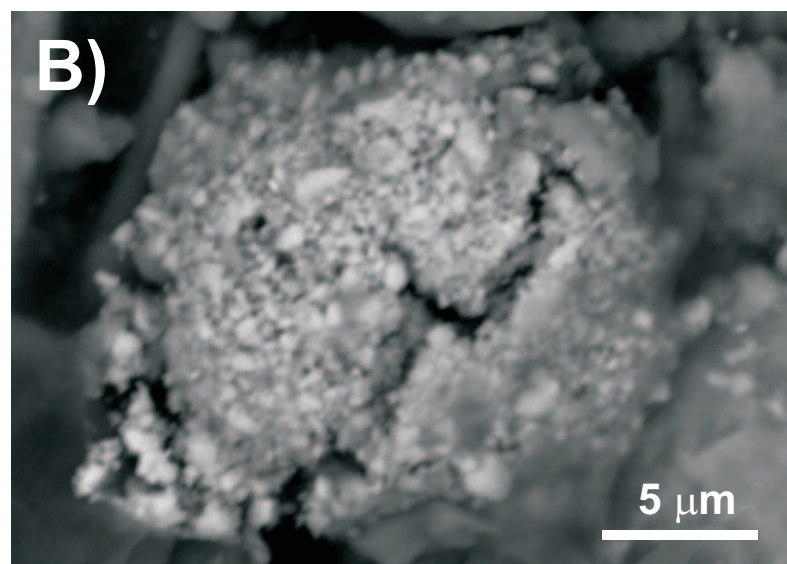
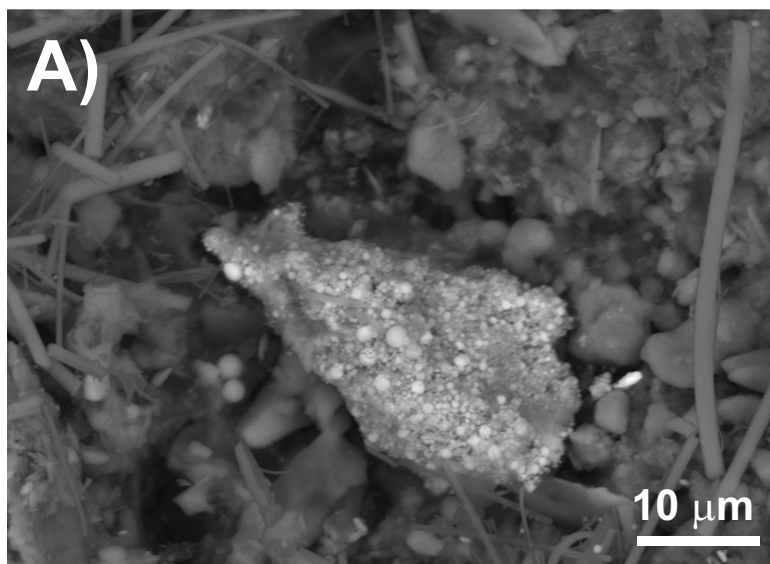
## A) BCN-10.1 urban site

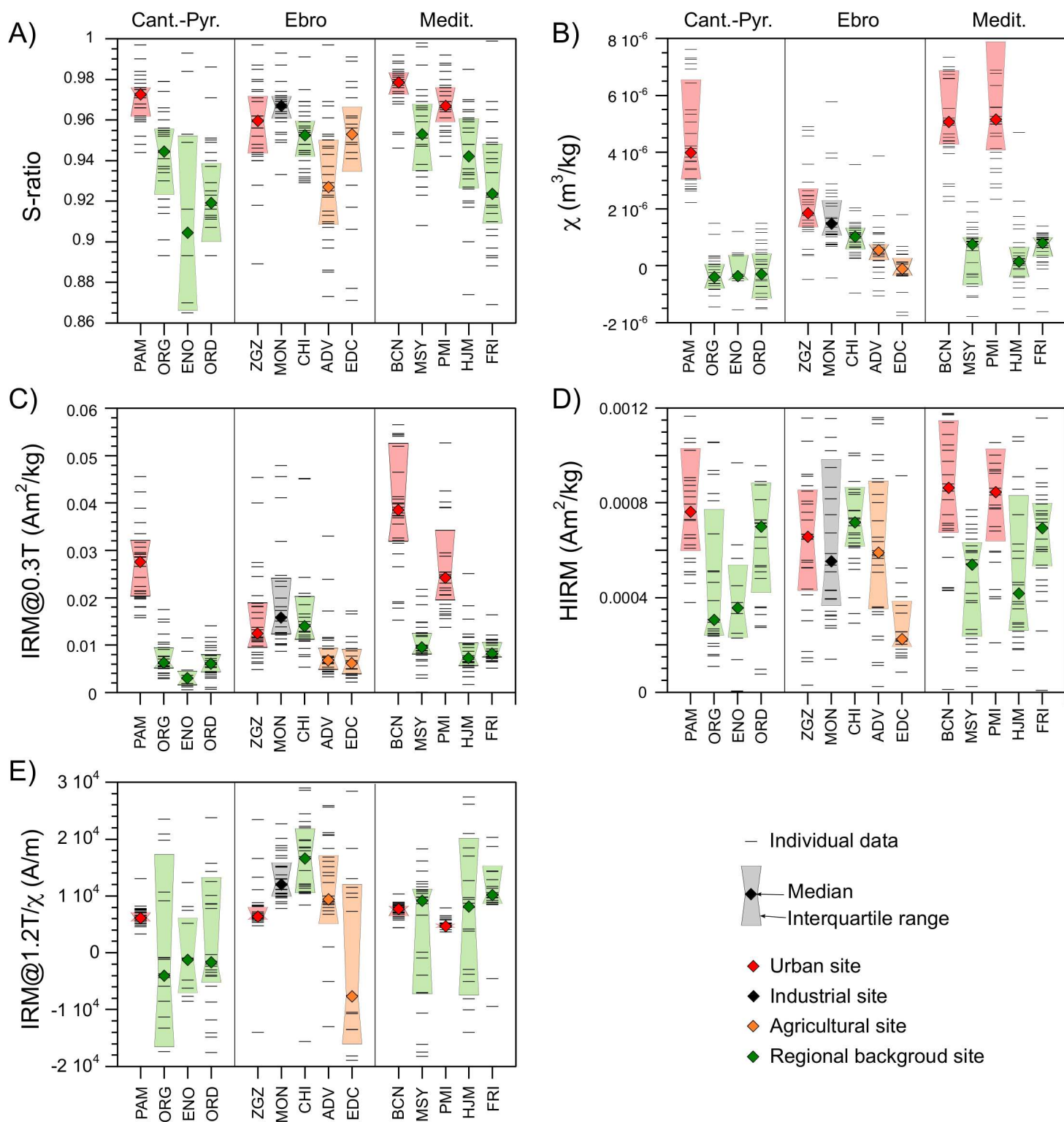


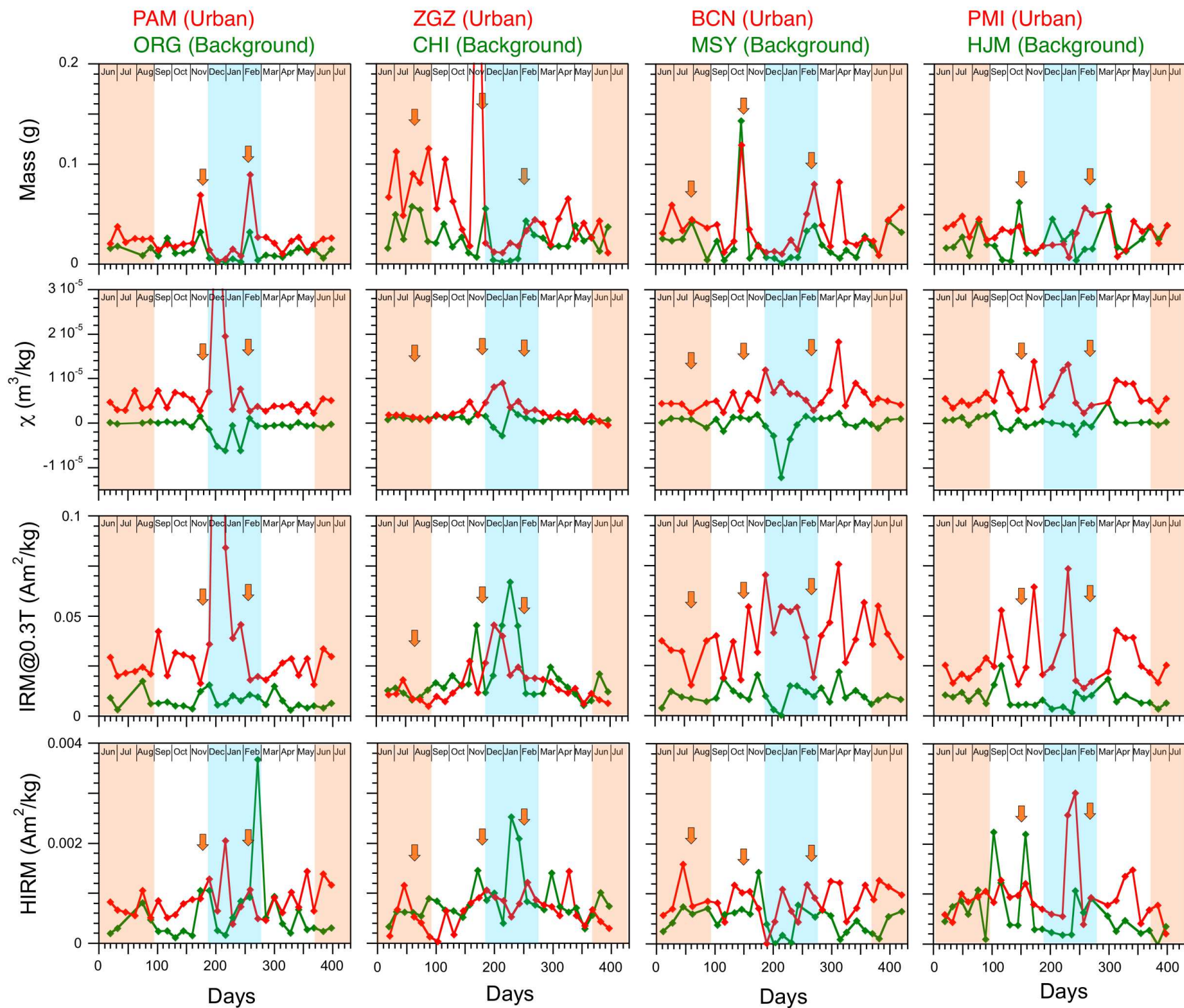
## B) FRI-10.3 Background site + NAD

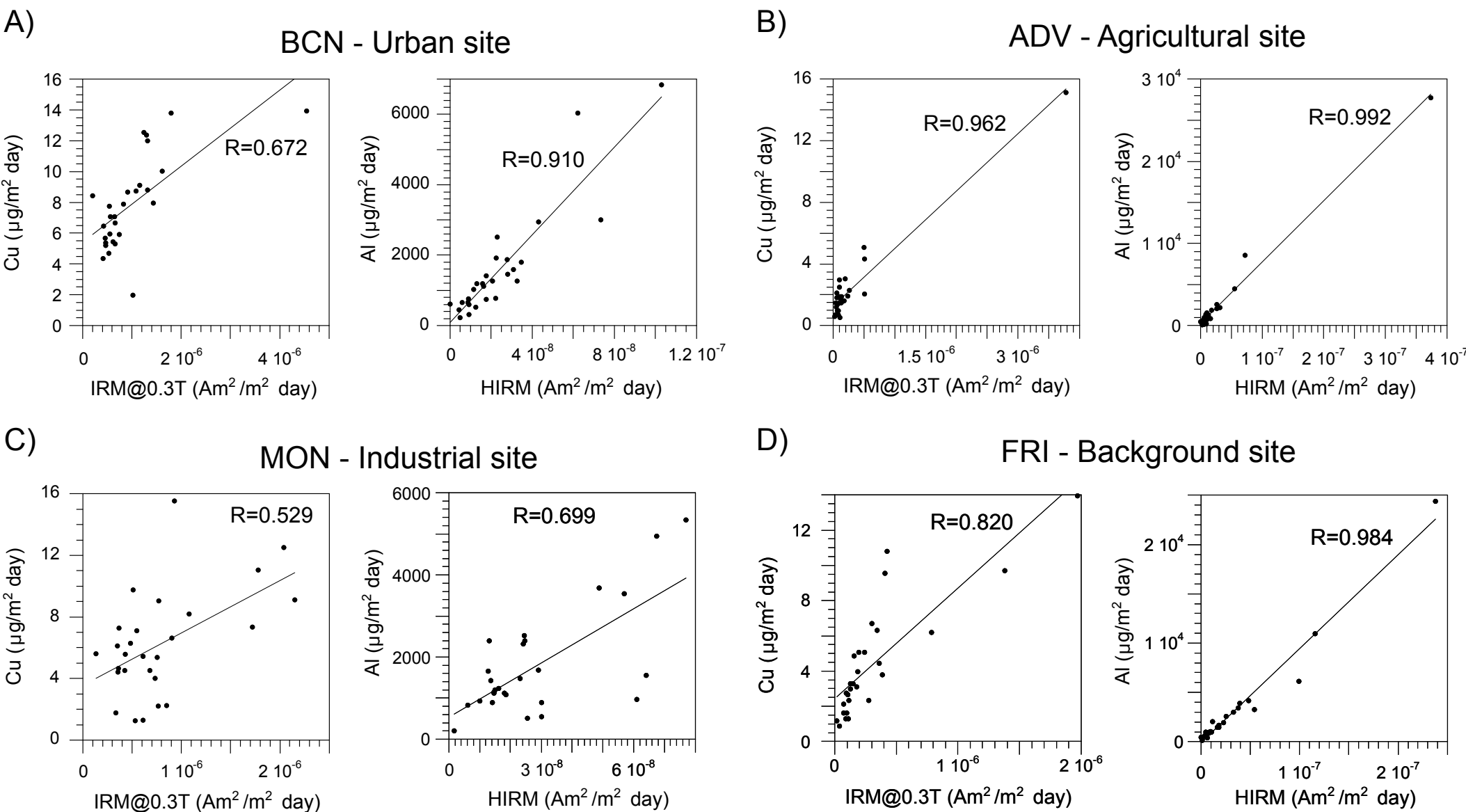


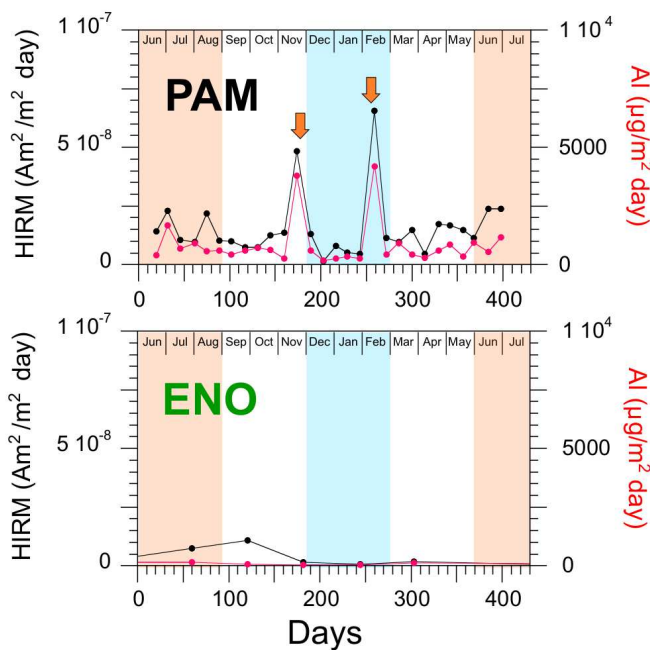




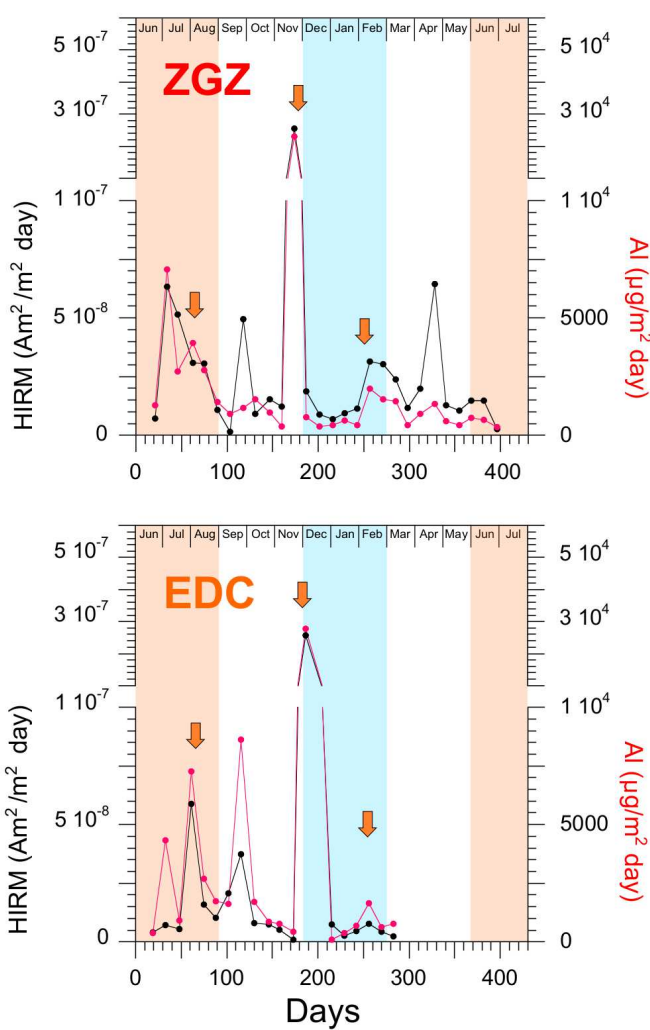
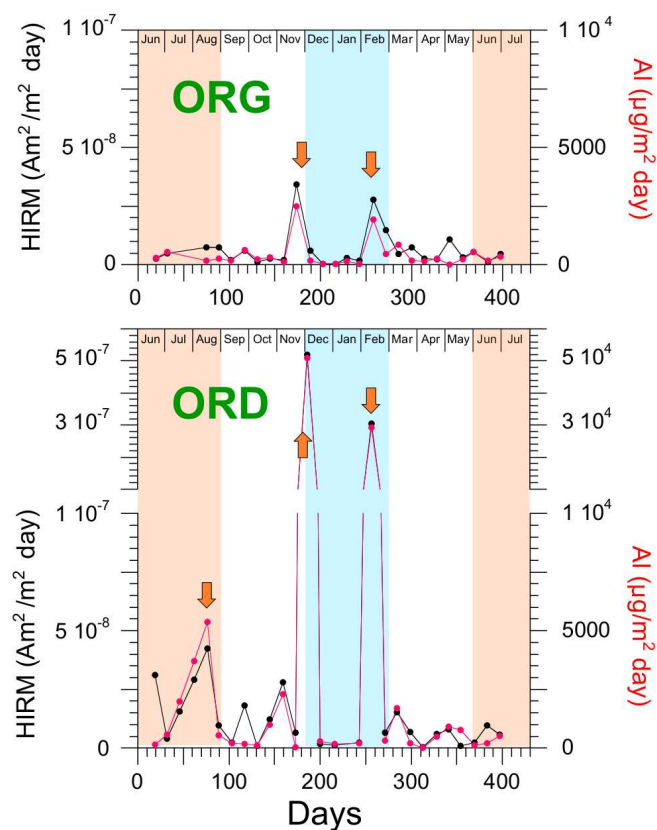




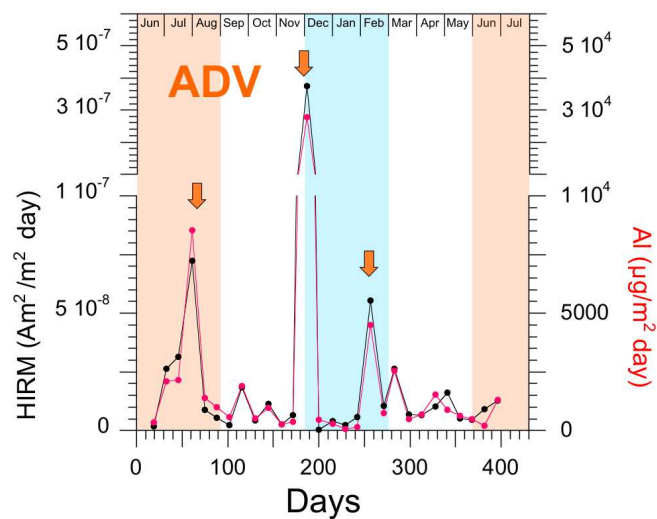
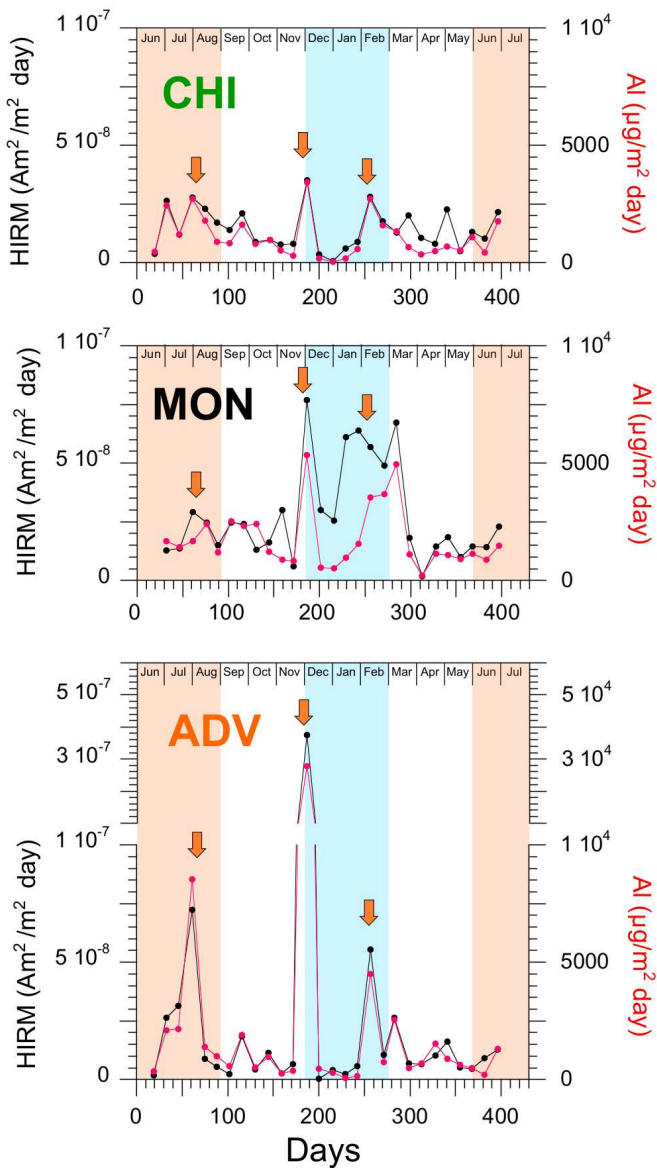


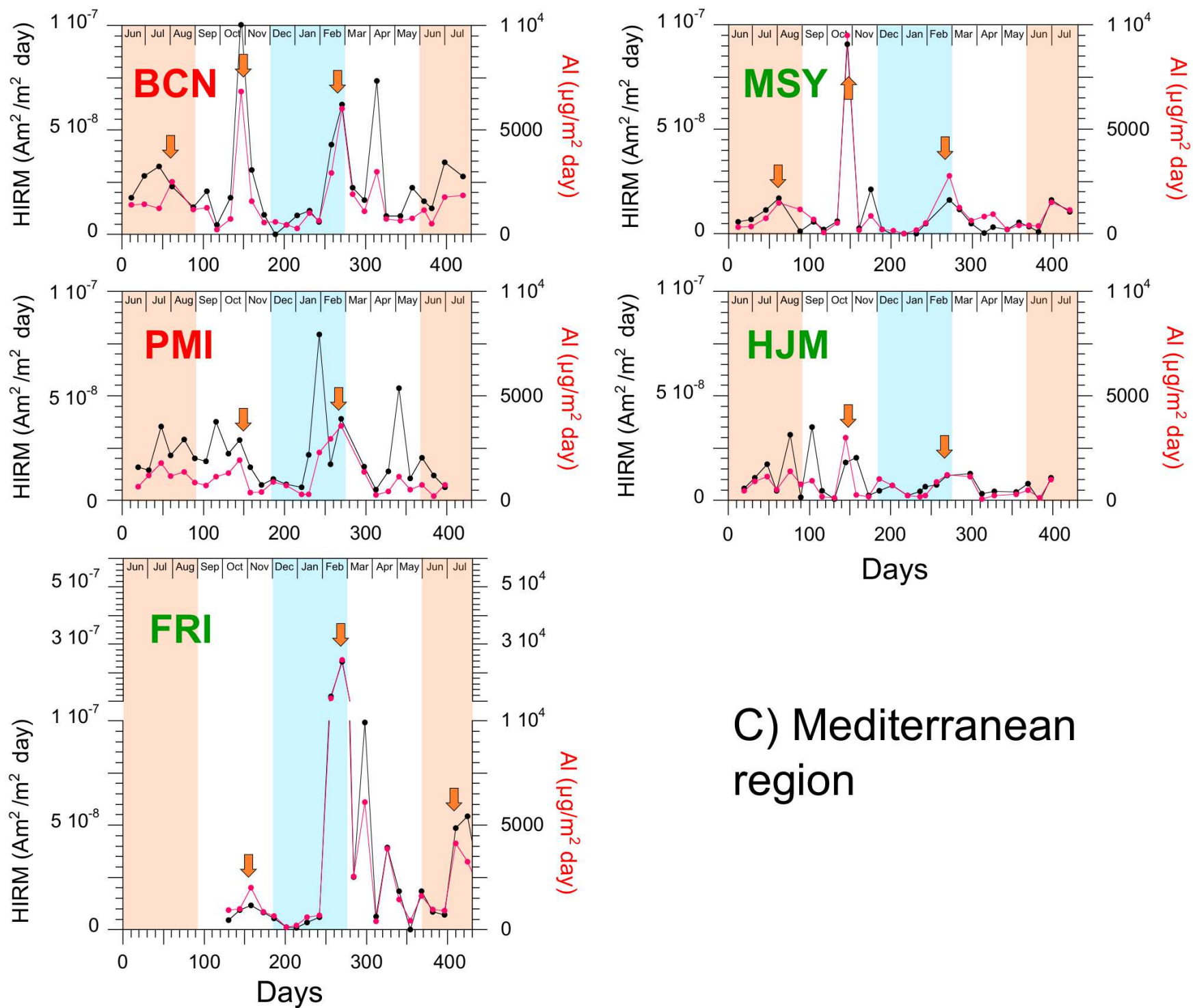


## A) Cantabrian-Pyrenean region

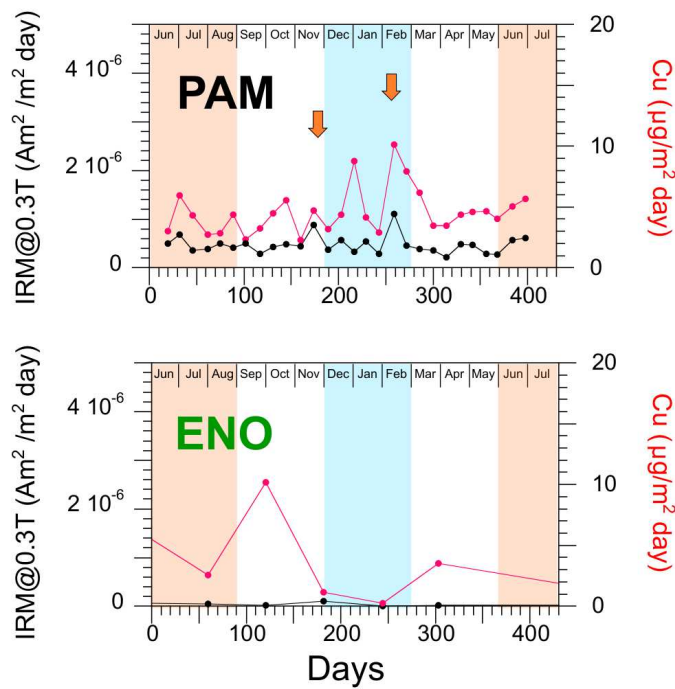


## B) Ebro Valley

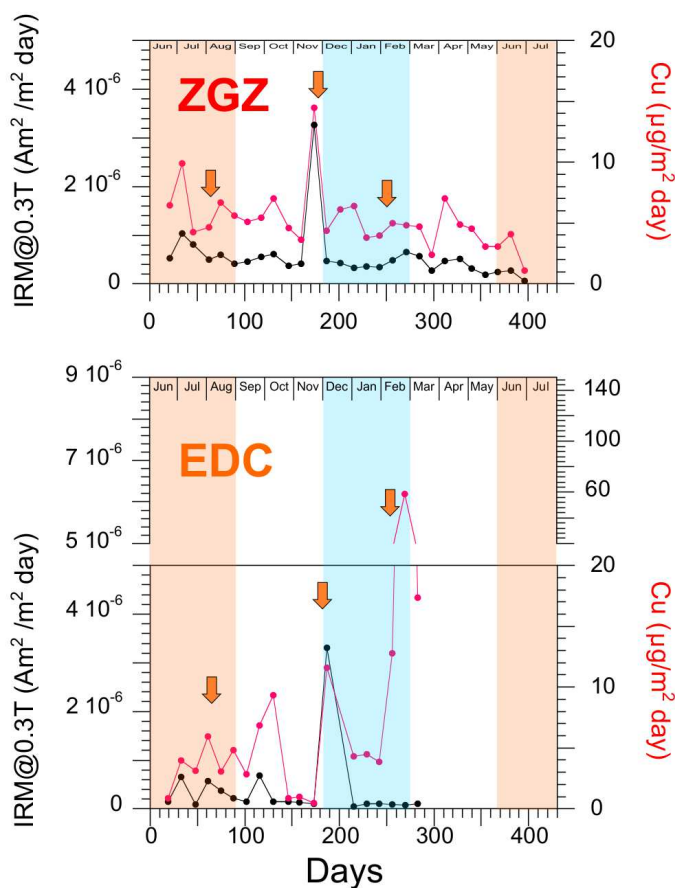
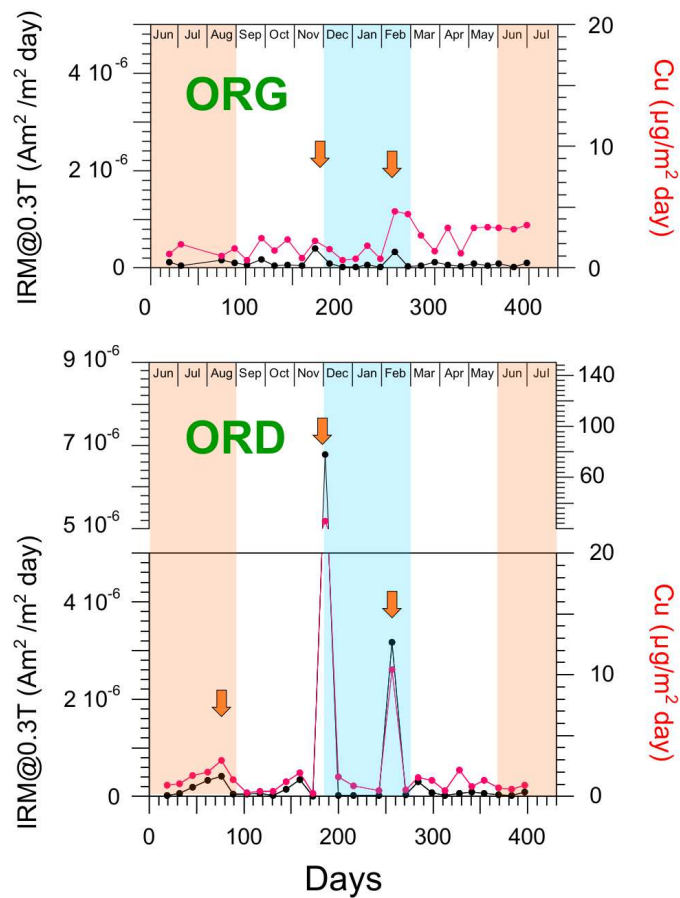




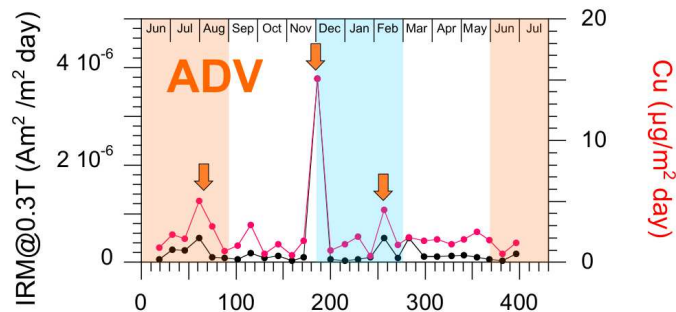
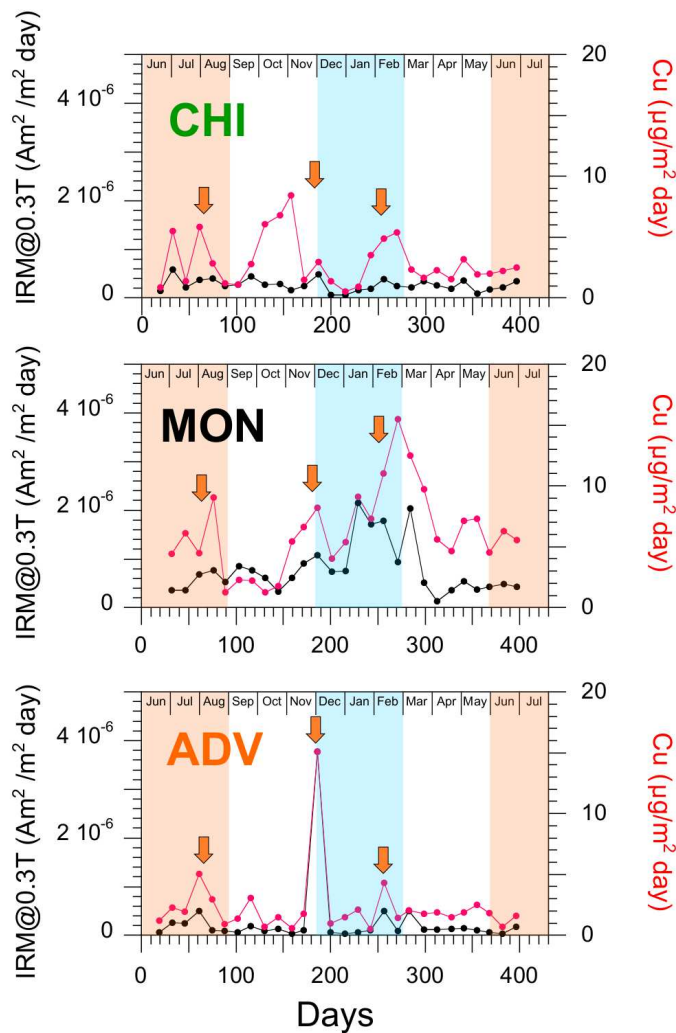
C) Mediterranean region

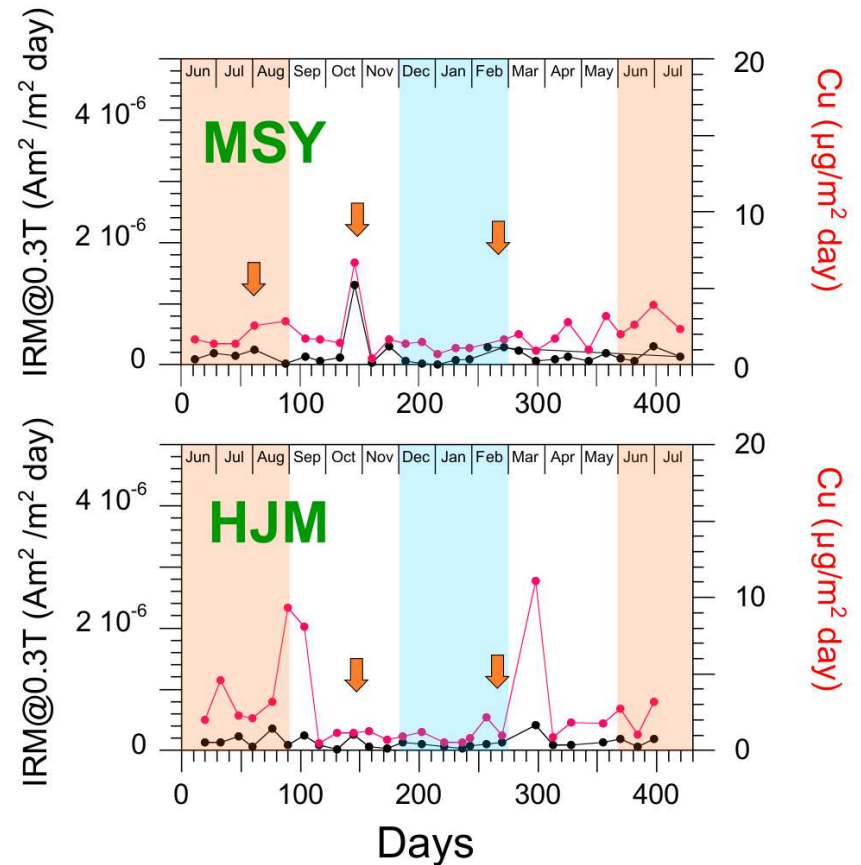
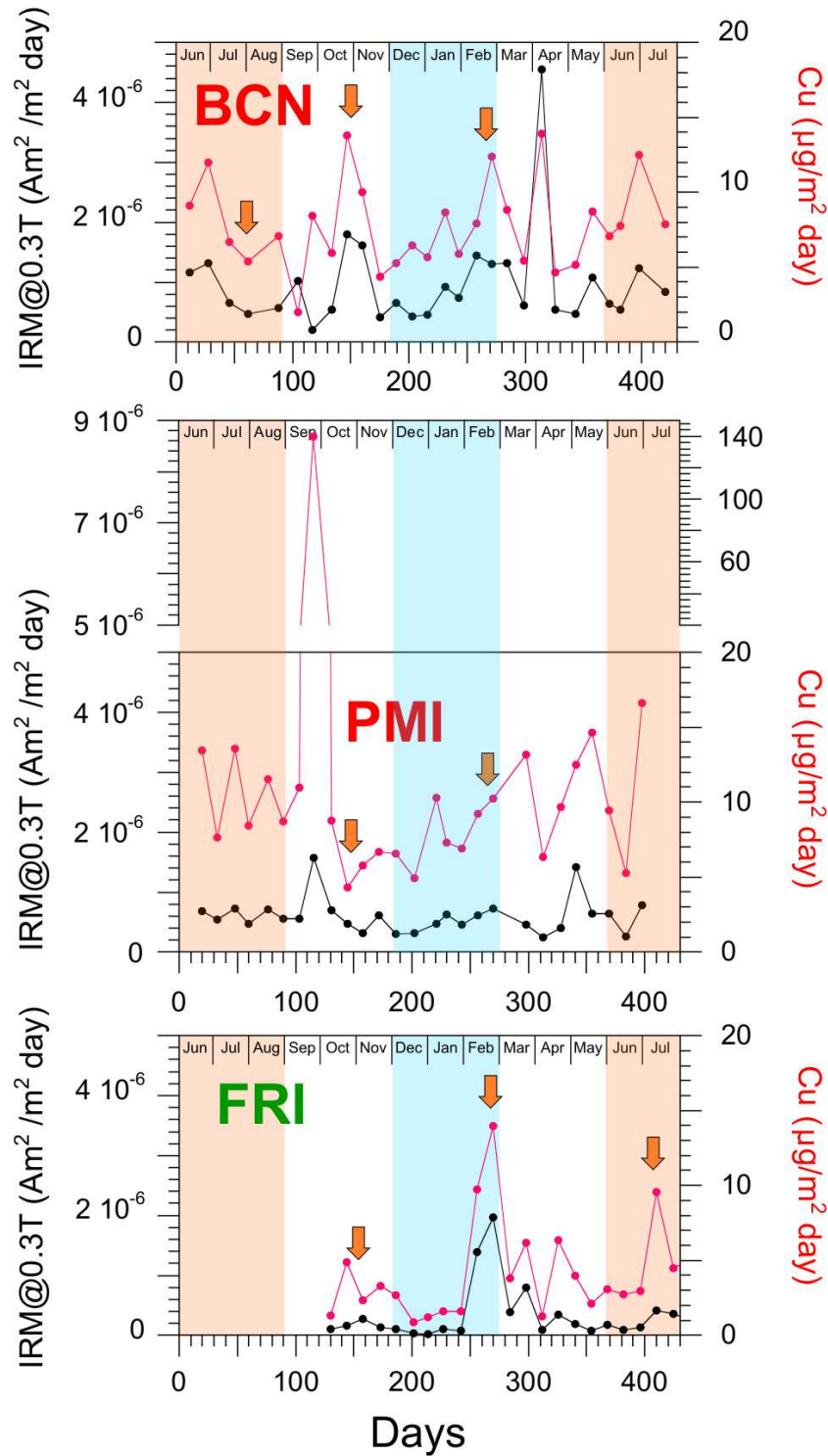


## A) Cantabrian-Pyrenean region



## B) Ebro Valley





C) Mediterranean region



| Site   | Abbreviation | Type                | Latitude | Longitude | Altitude<br>m a. s. l. | Annual rainfall<br>(l/m <sup>2</sup> ) | Start date | End date  |
|--|--------------|---------------------|----------|-----------|------------------------|--|------------|-----------|
| <b><i>Cantabrian and Pyrenean region</i></b> |              |                     |          |           |                        |  |            |           |
| Pamplona                                     | PAM          | Urban               | 42,800   | -1,650    | 450                    | 711                                    | Jun. 2016  | Jun. 2017 |
| Orgi *                                       | ORG          | Regional background | 42,960   | -1,680    | 512                    | 766                                    | Jun. 2016  | Jun. 2017 |
| Ordesa                                       | ORD          | Regional background | 42,650   | -0,090    | 1190                   | 1119                                   | Jun. 2016  | Jun. 2017 |
| Enol   | ENO          | Regional background | 43,272   | -4,991    | 1070                   | 1722                                   | Dec. 2015  | Jun. 2017 |
| <b><i>Ebro Valley</i></b>                    |              |                     |          |           |                        |  |            |           |
| Zaragoza                                     | ZGZ          | Urban               | 41,640   | -0,890    | 247                    | 295                                    | Jun. 2016  | Jun. 2017 |
| Monzón                                       | MON          | Industrial          | 41,920   | 0,190     | 279                    | 519                                    | Jun. 2016  | Jun. 2017 |
| Chiprana *                                   | CHI          | Regional background | 41,290   | -0,240    | 156                    | 287                                    | Jun. 2016  | Jun. 2017 |
| Almudévar                                    | ADV          | Agricultural        | 42,020   | -0,660    | 456                    | 473                                    | Jun. 2016  | Jun. 2017 |
| Ejea   | EDC          | Agricultural        | 42,060   | -1,130    | 336                    | 251                                    | Jun. 2016  | Mar. 2017 |
| <b><i>Mediterranean region</i></b>           |              |                     |          |           |                        |  |            |           |
| Barcelona                                    | BCN          | Urban               | 41,390   | 2,120     | 63                     | 563                                    | Jun. 2016  | Jun. 2017 |
| Montserrat *                                 | MSY          | Regional background | 41,780   | 2,350     | 720                    | 546                                    | Jun. 2016  | Jun. 2017 |
| Palma  | PMI          | Urban               | 39,570   | 2,660     | 12                     | 601                                    | Jun. 2016  | Jun. 2017 |
| Joan March *                                 | HJM          | Regional background | 39,680   | 2,690     | 172                    | 544                                    | Jun. 2016  | Jun. 2017 |
| Frigiliana                                   | FRI          | Regional background | 36,792   | -3,898    | 320                    | 506                                    | Oct. 2016  | Jul. 2017 |

| Sample   | Type               | $B_{cr}$ (mT) | $M_r/M_s$ | $B_{cr}/B_c$ | Component 1 |      |      |        | Component 2 |      |      |        | Component 3 |      |
|----------|--------------------|---------------|-----------|--------------|-------------|------|------|--------|-------------|------|------|--------|-------------|------|
|          |                    |               |           |              | $Bh$ (mT)   | DP   | S    | EC     | $Bh$ (mT)   | DP   | S    | EC     | $Bh$ (mT)   | DP   |
| PAM-11.2 | Urban              | 31.90         | 0.062     | 6.070        | 8.40        | 0.34 | 1.01 | 0.0885 | 53.53       | 0.36 | 0.93 | 0.9038 | 897.20      | 0.25 |
| ZGZ-10.3 | Urban              | 33.58         | 0.052     | 7.490        | 9.40        | 0.33 | 1.05 | 0.0658 | 51.40       | 0.36 | 0.83 | 0.8819 | 807.12      | 0.37 |
| BCN-2.3  | Urban              | 33.40         | 0.100     | 4.480        | 13.26       | 0.41 | 1.02 | 0.1101 | 51.91       | 0.37 | 0.90 | 0.8333 | 686.46      | 0.30 |
| BCN-10.1 | Urban              | 28.56         | 0.072     | 4.270        | 5.39        | 0.27 | 1.03 | 0.0270 | 42.45       | 0.34 | 0.91 | 0.9604 | 359.55      | 0.13 |
| PMI-2.3  | Urban              | 27.77         | 0.080     | 5.470        | 12.71       | 0.40 | 1.05 | 0.1755 | 51.82       | 0.37 | 0.90 | 0.7874 | 1338.26     | 0.29 |
| PMI-8.3  | Urban              | 30.89         | 0.066     | 6.700        | 9.34        | 0.34 | 0.93 | 0.1145 | 51.80       | 0.36 | 0.82 | 0.8763 | 493.02      | 0.25 |
| MON-4.1  | Industrial         | 32.14         | 0.122     | 3.540        | -           | -    | -    | -      | 41.44       | 0.42 | 0.83 | 0.9544 | 860.60      | 0.30 |
| MON-8.2  | Industrial         | 32.20         | 0.096     | 4.010        | 10.50       | 0.46 | 1.05 | 0.1000 | 49.09       | 0.37 | 0.83 | 0.8661 | 1101.13     | 0.33 |
| MON-13.1 | Industrial         | 24.82         | 0.142     | 3.210        | 2.31        | 0.41 | 1.07 | 0.0190 | 36.84       | 0.44 | 0.99 | 0.9465 | 791.32      | 0.26 |
| EDC-6.3  | Agricultural       | 25.68         | 0.155     | 3.360        | 0.50        | 0.50 | 1.05 | 0.0204 | 35.22       | 0.39 | 0.86 | 0.9125 | 422.44      | 0.36 |
| PAM-12.3 | Urban + NAD        | 20.00         | 0.136     | 3.180        | 2.90        | 0.37 | 1.00 | 0.0592 | 24.26       | 0.34 | 1.05 | 0.7230 | 450.76      | 0.26 |
| ZGZ-18.1 | Urban + NAD        | 28.75         | 0.070     | 5.030        | 3.16        | 0.45 | 1.12 | 0.0278 | 42.52       | 0.40 | 0.88 | 0.9241 | 951.25      | 0.35 |
| ORD-5.3  | Background + NAD   | 28.25         | 0.111     | 3.420        | 5.31        | 0.16 | 1.03 | 0.0244 | 34.82       | 0.38 | 1.02 | 0.7801 | 588.86      | 0.28 |
| ORD-13.1 | Background + NAD   | 17.86         | 0.197     | 2.670        | 3.15        | 0.27 | 1.04 | 0.0434 | 21.63       | 0.34 | 1.03 | 0.7740 | 795.11      | 0.33 |
| ORD-18.1 | Background + NAD   | 18.44         | 0.192     | 2.720        | 3.14        | 0.33 | 0.89 | 0.0474 | 22.57       | 0.33 | 1.03 | 0.7596 | 580.01      | 0.28 |
| HJM-19.1 | Background + NAD   | 26.56         | 0.120     | 3.600        | 2.82        | 0.18 | 1.03 | 0.0303 | 32.61       | 0.36 | 1.02 | 0.7665 | 547.26      | 0.28 |
| FRI-10.3 | Background + NAD   | 22.83         | 0.175     | 2.930        | 3.19        | 0.30 | 1.05 | 0.0249 | 27.52       | 0.37 | 1.03 | 0.7559 | 736.59      | 0.36 |
| ADV-8.4  | Agricultural + NAD | 26.49         | 0.129     | 3.160        | 1.04        | 0.33 | 1.13 | 0.0096 | 32.97       | 0.43 | 0.95 | 0.8406 | 1023.08     | 0.27 |
| ADV-13.1 | Agricultural + NAD | 17.82         | 0.217     | 2.560        | 2.37        | 0.23 | 1.03 | 0.0264 | 19.20       | 0.34 | 0.93 | 0.7026 | 666.47      | 0.37 |

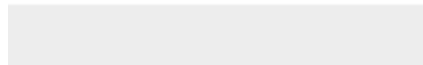
**Component 4**

| <i>S</i> | <i>EC</i> | <i>Bh (mT)</i> | <i>DP</i> | <i>S</i> | <i>EC</i> |
|----------|-----------|----------------|-----------|----------|-----------|
| 1.11     | 0.0077    | -              | -         | -        | -         |
| 1.10     | 0.0523    | -              | -         | -        | -         |
| 1.14     | 0.0565    | -              | -         | -        | -         |
| 1.15     | 0.0127    | -              | -         | -        | -         |
| 1.13     | 0.0371    | -              | -         | -        | -         |
| 1.19     | 0.0092    | -              | -         | -        | -         |
| 1.14     | 0.0456    | -              | -         | -        | -         |
| 1.06     | 0.0339    | -              | -         | -        | -         |
| 1.28     | 0.0345    | -              | -         | -        | -         |
| 1.03     | 0.0671    | -              | -         | -        | -         |
| 1.08     | 0.0624    | 105.81         | 0.25      | 1.13     | 0.0624    |
| 1.11     | 0.0481    | -              | -         | -        | -         |
| 0.98     | 0.0737    | 128.19         | 0.30      | 1.30     | 0.0737    |
| 1.05     | 0.0662    | 165.20         | 0.31      | 1.07     | 0.0662    |
| 1.06     | 0.0611    | 149.61         | 0.29      | 1.16     | 0.0611    |
| 1.01     | 0.0824    | 120.98         | 0.29      | 1.00     | 0.0824    |
| 0.89     | 0.0886    | 124.09         | 0.31      | 1.11     | 0.0886    |
| 1.11     | 0.0582    | 97.85          | 0.33      | 1.12     | 0.0582    |
| 1.14     | 0.0783    | 99.49          | 0.36      | 1.09     | 0.0783    |

| Site | IRM@0.3T Vs Cu | IRM@0.3T Vs Sb | HIRM Vs Al  | HIRM Vs Ti  |
|------|----------------|----------------|-------------|-------------|
| PAM  | 0.47           | 0.44           | 0.31        | 0.26        |
| ORG  | 0.27           | 0.45           | <b>0.89</b> | <b>0.79</b> |
| ORD  | <b>0.99</b>    | <b>0.99</b>    | <b>0.99</b> | <b>0.97</b> |
| ENO  | 0.19           | 0.02           | 0.45        | 0.06        |
| ZGZ  | <b>0.85</b>    | 0.29           | <b>0.96</b> | <b>0.88</b> |
| MON  | 0.52           | 0.61           | 0.66        | 0.62        |
| CHI  | 0.34           | 0.19           | <b>0.89</b> | <b>0.81</b> |
| ADV  | <b>0.96</b>    | 0.31           | <b>0.99</b> | <b>0.85</b> |
| EDC  | 0.01           | 0.4            | <b>0.97</b> | <b>0.83</b> |
| BCN  | 0.67           | 0.69           | <b>0.91</b> | <b>0.95</b> |
| MSY  | <b>0.81</b>    | <b>0.92</b>    | <b>0.96</b> | <b>0.97</b> |
| PMI  | 0.69           | <b>0.72</b>    | 0.57        | 0.44        |
| HJM  | 0.59           | 0.65           | 0.54        | 0.5         |
| FRI  | <b>0.82</b>    | 0.21           | <b>0.98</b> | <b>0.96</b> |



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**Supplementary Material**  
Supplementary Table 1.xlsx



We wish to confirm that there are no known conflicts of interest associated with this publication and there has been no significant financial support for this work that could have influenced its outcome.